## Reports of the Department of Geodetic Science

 Report No. 208
# CLOSED COVARIANCE EXPRESSIONS FOR GRAVITY ANOMALIES, GEOID UNDULATIONS, AND DEFLECTIONS OF THE VERTICAL IMPLIED BY ANOMALY DEGREE VARIANCE MODELS. 

by<br>C.C. Tscherning<br>and<br>Richard H. Rapp



The Ohic State University Department of Geodetic Science<br>1958 Neil Avenue<br>Columbus, Ohio 43210

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This report first develops a new anomaly degree variance model by considering potential coeffic ient information to degree 20, and updated values of the point anomaly variance ( $1795 \mathrm{mgal}^{2}$ ), the $1^{\circ}$ block variance ( $920 \mathrm{mgal}^{2}$ ) and the $5^{\circ}$ block variance ( $302 \mathrm{mgal}^{2}$ ), the variances being given with respect to the Geodetic Reference System 1967. This new model was computed assuming that anomaly information was given on a sphere of radius 6371 km with the radius of the best fitting Bjerhammer sphere found to be 6369.8 km .

This new model and several other models were used to develop closed expressions for the covariance and cross-covariance functions between gravity anomalies, geoid undulations (or height anomalies), and deflections of the vertical. It is shown how these global covariance expressions can be modified for use as local covariances and for use when mean anomalies are being considered. A Fortran subroutine is provided for the determination of the covariance values implied by the recommended anomaly degree variance model.

## FOREWORD

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## 1. Introduction

In carrying out the estimation of gravimetric dependent quantities using the methods of least squares collocation (Moritz, 1972) we need to have an analytical function that can be used to determine the covariance functions for such quantities as anomalies, deflections of the vertical, geoid undulations etc. Generally speaking a numerical covariance function for anomalies can be determined from anomaly data. The resultant function can be considered by determining a model for the anomaly degree variances. Tscherning (1972) has shown how such anomaly degree variance models may be used to determine the covariance models for several gravimetric quantities. Since we need the best estimates of our covariance models for the application of least squares collocation, it is appropriate that we use the latest available data in determining our models. In addition we are now at a stage where refinements in anomaly degree variance modeling, beyond that used by Rapp (1973) can be considered.

The purpose of this report is to describe recent computations made and subsequent analytical work that leads to improved analytical covariance models.

## 2. Preliminary Equations

In this section some of the relevant formulas to be used in later sections will be presented.

We first consider our covariance function which for the purposes of this report will be considered as stationary and isotropic. Then we can follow the standard definition (Heiskanen and Moritz, 1967) of the anomaly covariance as the mean product (at a given distance) of the anomaly pair $\Delta g_{p}, \Delta g_{Q}$. Thus:

$$
\begin{equation*}
\mathrm{C}(\mathrm{P}, \mathrm{Q})=\operatorname{cov}\left(\Delta \mathrm{g}_{\mathrm{P}}, \Delta \mathrm{~g}_{\mathrm{Q}}\right)=\mathrm{M}\left(\Delta \mathrm{~g}_{\mathrm{P}}, \Delta \mathrm{~g}_{\mathrm{Q}}\right) \tag{1}
\end{equation*}
$$

On a plane the distance, or anomaly separation is usually specified by some linear distance (such as 20 km ). If we deal with data on a sphere we usually considered the distance to be defined as $\psi$ a spherical arc so that we are interested in values of $C(\psi)$. At $\psi=0, C(\psi)$ becomes the anomaly variance. For the estimation of $C(\psi)$ from anomaly data given on the surface of a sphere, we can write (Heiskanen and Moritz, 1967, p. 258):

$$
\begin{equation*}
C(\psi)=\frac{1}{4 \pi} \int_{\lambda=0}^{2 \pi} \int_{\theta=0}^{\pi} \frac{1}{2 \pi} \int_{\alpha=0}^{2 \pi} \Delta g(\theta, \lambda) \Delta g\left(\theta^{\prime}, \lambda^{\prime}\right) \sin \theta \mathrm{d} \theta \mathrm{~d} \lambda \mathrm{~d} \alpha \tag{2}
\end{equation*}
$$

where $\theta$ is a polar angle ( 0 at the north pole), $\lambda$ is the longitude and $\alpha$ is an azimuth.

We will obtain from (2), a point anomaly covariance function if the $\Delta g$ values are point anomalies or we will obtain a mean anomaly covariance function (for a specific block size) if the $\Delta g$ values are mean anomalies. In practice the sphere is not completely covered by anomalies so that an expression that may be used to compute the covariance between any two functions $f_{j}$ and $f_{k}$ given in blocks on the sphere whose area is $A_{y}$ and $A_{k}$ respectively may be written: (Kaula, 1966a, p. I. B. 7).

$$
\begin{equation*}
C(\psi)=\frac{\sum A_{\jmath} A_{k} f_{\mathrm{g}} f_{k}}{\sum A_{\jmath} A_{k}} \tag{3}
\end{equation*}
$$

In our case $f_{j}=\overline{\Delta g}(\theta, \lambda)$ and $f_{k}=\overline{\Delta g}\left(\theta^{\prime}, \lambda^{\prime}\right)$ where the overbar signifies a mean anomaly. If the anomalies are given in equal area block (3) becomes:

$$
\begin{equation*}
C(\psi)=\frac{\sum \mathrm{f}_{\mathrm{y}} \mathrm{f}_{\mathrm{k}}}{\mathrm{n}} \tag{4}
\end{equation*}
$$

where $n$ is the number of products taken at a given spherical distance $\psi$. In practice the distance $\psi$ to which a special product at distance $\psi_{j k}$ is determined by the equation:

$$
\begin{equation*}
\psi-\frac{\Delta \psi}{2}<\psi_{j k}<\psi+\frac{\Delta \psi}{2} \tag{5}
\end{equation*}
$$

where $\Delta \psi$ is a suitably chosen range. In our numerical results to be discussed later, $\Delta \psi$ was specified to be $P$.

A more fundamental covariance function than that of the gravity anomalies is that of the disturbing potential, $\mathrm{K}(\mathrm{P}, \mathrm{Q})$. We generally do not estimate $\mathrm{K}(\mathrm{P}, \mathrm{Q})$ from numerical data, but rather consider the following series representation for it: (Moritz, 1972, p. 88):

$$
\begin{equation*}
K(P, Q)=\sum_{\ell=0}^{\infty} \sigma_{2}\left(\frac{R^{2}}{\mathrm{rr}^{\prime}}\right)^{\ell+1} P_{2}(\cos \psi) \tag{6}
\end{equation*}
$$

where: $\quad \sigma_{l}$ are the degree variances of the anomalous potential;
$R$ is the radius of the Bjerhammar sphere;
$r, r^{\prime}$ are the geocentric radii to points $P$ and $Q$ which are separated by a spherical radius $\psi$.

For convenience we let:

$$
\begin{equation*}
\mathrm{s}=\frac{\mathrm{R}^{2}}{\mathrm{rr}^{7}} \tag{7}
\end{equation*}
$$

In the case that we are dealing with information at the approximate surface of the earth, it is convenient to take $r^{\prime}=R_{e}^{2}$ where $R_{0}$ is a mean earth radius. Then:

$$
\begin{equation*}
\mathrm{s}=\left(\frac{\mathrm{R}}{\mathrm{R}_{\theta}}\right)^{2} \tag{8}
\end{equation*}
$$

We then can write:

$$
\begin{equation*}
K(P, Q)=\sum_{\ell=0}^{\infty} \sigma_{\ell} s^{\ell+1} P_{\ell}(\cos \psi) \tag{9}
\end{equation*}
$$

We can also write the anomaly covariances in a series expression as Moritz, 1972, p. 89):

$$
\begin{equation*}
\mathrm{C}(\mathrm{P}, \mathrm{Q})=\sum_{\ell=0}^{\infty} \mathrm{c}_{\ell} \mathrm{s}^{\ell+z^{2}} \mathrm{P}_{\ell}(\cos \psi) \tag{10}
\end{equation*}
$$

where $c_{\ell}$ are the anomaly degree variances. As written, equation (10) would yield a point anomaly covariance. In order to obtain a mean anomaly covariance we can use the $R_{\ell}$ functions of Meissl (1970, p. 23) or the $q_{l}$ functions of Pellinen (1966). Using $\beta_{\ell}$, the modification of equation (10) yields:

$$
\begin{equation*}
\overline{\mathrm{C}}(\mathrm{P}, \mathrm{Q})=\sum_{\ell=0}^{\infty} \beta_{\ell}^{2} \mathrm{c}_{\ell} \mathrm{s}^{\ell+2} \mathrm{P}_{\ell}(\cos \psi) \tag{11}
\end{equation*}
$$

where $P$ and $Q$ now refer to anomaly blocks. $\beta_{\ell}$ is defined as follows: (Meissl, 1970, p. 24):

$$
\begin{equation*}
B_{\ell}=\frac{1}{1-\cos \psi_{0}} \quad \frac{1}{2 \ell+1}\left[P_{\ell-1}\left(\cos \psi_{b}\right)-P_{\ell+1}\left(\cos \psi_{0}\right)\right] \tag{12}
\end{equation*}
$$

where $\psi_{0}$ is the circular cap radius of the mean anomaly block whose covariance is to be computed. We have (for example):

$$
\begin{align*}
& \beta_{0}=1  \tag{12~A}\\
& \beta_{1}=\frac{1}{2} \sin \psi_{6} \cot \frac{\psi_{0}}{2} \tag{12B}
\end{align*}
$$

Since we usually deal with rectangular blocks of dimension $s^{\circ}$, the corresponding $\psi_{0}^{\circ}$ can be found simply by equating the areas of the circular cap and the square blocks. Assuming a plane figure we write (for small blocks only):

$$
\begin{equation*}
\psi_{0}^{0}=s^{\circ} / \sqrt{\pi}=0.564 s^{\circ} \tag{13}
\end{equation*}
$$

As $\psi_{0} \rightarrow 0, B_{\ell} \rightarrow 1$
Since gravity anomalies are related to the disturbing potential by the following equation (valid in a spherical approximation which is the case considered here):

$$
\begin{equation*}
\Delta g=\frac{-\partial T}{\partial r}-\frac{2}{R} T \tag{14}
\end{equation*}
$$

where $T$ is the disturbing potential, we can relate the anomaly degree variances ( $c_{\ell}$ ) and the degree variances of the anomalous potential $\left(\sigma_{\ell}\right)$ by:

$$
\begin{equation*}
\sigma_{\ell}=\frac{R^{2}}{(\ell-1)^{2}} c_{\ell} \tag{15}
\end{equation*}
$$

Analytic models for either $\sigma_{\ell}$, or $\mathrm{c}_{\ell}$ have been described by Lauritzen (1973), Tscherning (1972), by Rapp (1973a) and implicitly by Kaula (1966b, p. 98).

The inverse of equation (10) is:

$$
\begin{equation*}
\mathrm{c}_{\ell}=\frac{2 \ell+1}{2} \mathrm{~s}^{-\left(\ell^{+z}\right)} \int_{0}^{\pi} C(\psi) P_{\ell}(\cos \psi) \sin \psi d \psi \tag{16}
\end{equation*}
$$

Equation (16) is written assuming $C(\psi)$ is a point anomaly covariance function referring to a sphere whose radius is $R_{e}$. If $C(\psi)$ is a point anomaly covariance function, then (16) with $\mathrm{C}(\psi)$ replaced by $\overline{\mathrm{C}}(\psi)$ will yield a mean anomaly degree variance $\overline{\mathrm{c}}_{\ell}$, which is related to $c_{\ell}$ through the $b_{\ell}$ equations:

$$
\begin{equation*}
\overline{\mathrm{c}}_{\ell}=\beta_{\ell}^{2} \mathrm{c}_{\ell} \tag{17}
\end{equation*}
$$

Thus, knowing $\overline{\mathrm{C}}(\psi)$ we can find $\overline{\mathrm{c}}_{\ell}$ from (16) and $\mathrm{c}_{\ell}$ from (17) knowing the size of the anomaly blocks to which $\overline{\mathrm{C}}(\psi)$ refers. Specifically we can write:

$$
\begin{equation*}
c_{\ell}=\frac{2 \ell+1}{2} \frac{1}{\beta_{\ell}^{2} s^{(\ell+2)}} \int_{0}^{\pi} \overline{\mathrm{C}}(\psi) \mathrm{P}_{\ell}(\cos \psi) \sin \psi \mathrm{d} \psi \tag{16A}
\end{equation*}
$$

## 3. Numerical $1^{\circ}$ Covariance Functions

We first start our numerical determinations by the estimation of the covariance function for $1^{\circ}$ (approximately) equal area anomalies. One degree covariance functions have been previously estimated for $\psi$ values from $0^{\circ}$ to $7^{\circ}$ by Kaula (1966c) and by

Rapp (1972). The values found in the past studies were based on analyzing $1^{\circ}$ anomalies within a $5^{\circ}$ equal area anomaly so that product pairs in adjacent $5^{\circ}$ blocks were not computed nor were product pairs for distances greater than $\psi$ approximately $7^{\circ}$ were considered. In addition, a programming error made the results of Kaula and Rapp somewhat erroncous.

Because of the limitations of previous estimations of the $1^{\circ}$ covariance function it was decided that it was appropriate to compte a global $1^{\circ}$ covariance function. The starting point was a recent collection of $29960,1^{\circ} \times 1^{\circ}$ equiangular mean free-air anomalies that was obtained by updating a $1^{\circ} \times 1^{\circ}$ mean anomaly set supplied by the Defense Mapping Agency - Aerospace Center. The updating was carried out using additional data along the lines of a previous update as described in Rapp (1972). These anomalies were all referred to the gravity formula of the Geodetic Reference System 1967. The $1^{\circ} \times 1^{\circ}$ equiangular tape was then converted to a set of 21828 (approximately) equal area anomalies. The subdivisions of these anomalies was such that the latitude increment was $1^{\circ}$ while the longitude increment was some integer degree of such size that the block was approximately equal in area to a $1^{\circ} \times 1^{\circ}$ block at the equator. The covariances were computed using equation (3) with the $\Delta \psi$ in equation (5) of $1^{\circ}$. The results of this computation are given in Table $A$ of the appendix. In this table the following quantities are given: number of product pairs, average $\psi$ (in degrees), covariance ( $\mathrm{mgal}^{2}$ ). For further use the 181 values given in Table A were interpolated to determine a covariance at 0.5 degree intervals. This interpolation was carried out using an Aitken-Lagrange interpolation using 20 points as implemented through subroutine DALI (and DATSG) of the IBM System/360 Scientific Subroutine Package (H20-0205-3), Version III. The resultant 361 values are given in Table One, being identified as the unmodified $\overline{\mathrm{C}}(\psi)$ values. The plot of this covariance function is shown in Figure One.

From these unmodified $\overline{\mathrm{C}}(\psi)$ values we can compute the smoothed anomaly degree variances from equation (16). Such values are shown for degree 0 through 10 in Table Two where $s$ and $B$ are taken to be one (causing a maximum error of less than $5 \%$ ). In addition values of $c_{\ell}$ from the recommended set of potential coefficients given by Rapp (1973b) are given for comparison purposes.


One Degree Anomaly Covariance Function

|  | $\overline{\mathrm{C}}(\psi)$ | $\overline{\mathrm{C}}(\psi)$ |  | $\overline{\mathrm{C}}(\psi)$ | $\overline{\mathrm{C}}(\psi)$ |
| :---: | :---: | :---: | :---: | :---: | :---: |
| $\psi^{\circ}$ | Modified | Unmodified | $\psi^{\circ}$ | Modified | Unmodified |
| 0.0 | 919.66 | 996.66 | 25.00 | 10.51 | 29.16 |
| 0.50 | 671.64 | 748.60 | 25.50 | 8.90 | 26.48 |
| 1.00 | 493.43 | 570.27 | 26.00 | 6.63 | 23.19 |
| 1.50 | 368.24 | 444.87 | 26.50 | 4.67 | 20.26 |
| 2.00 | 285.35 | 361.70 | 27.00 | 2.99 | 17.65 |
| 2.50 | 236.09 | 312.07 | 27.50 | 1.65 | 15.42 |
| 3.00 | 211.42 | 286.95 | 28.00 | 0.47 | 13.40 |
| 3.50 | 200.69 | 275.70 | 28.50 | -0.86 | 11.28 |
| 4.00 | 193.37 | 267.78 | 29.00 | -2. 23 | 9.16 |
| 4.50 | 176.88 | 250.62 | 29.50 | -2.86 | 7.83 |
| 5.00 | 155.86 | 228.85 | 30.00 | -3.29 | 6.73 |
| 5.50 | 146.38 | 218.55 | 30.50 | -4.57 | 4.84 |
| 6.00 | 141.39 | 212.67 | 31.00 | -5.84 | 2.99 |
| 6.50 | 133.39 | 203.72 | 31.50 | -6.17 | 2.12 |
| 7.00 | 124.92 | 194.23 | 32.00 | -6.14 | 1.65 |
| 7.50 | 119.86 | 188.09 | 32.50 | -6.46 | 0.87 |
| 8.00 | 117.38 | 184.48 | 33.00 | -7.38 | -0.48 |
| 8.50 | 116.89 | 182.80 | 33.50 | -8.88 | -2.37 |
| 9.00 | 115.10 | 179.77 | 34.00 | -10.61 | -4.46 |
| 9.50 | 107.02 | 170.40 | 34.50 | -11.85 | -6.03 |
| 10.00 | 96.52 | 158.56 | 35.00 | -12.71 | -7.19 |
| 10.50 | 92.45 | 153.11 | 35.50 | -13.52 | -8.28 |
| 11.00 | 90.54 | 149.79 | 36.00 | -14.16 | -9.17 |
| 11.50 | 82.86 | 140.66 | 36.50 | -14.36 | -9.60 |
| 12.00 | 74.57 | 130.89 | 37.00 | -14.49 | -9.93 |
| 12.50 | 72.40 | 127.21 | 37.50 | -15.11 | -10.74 |
| 13.00 | 71.77 | 125.06 | 38.00 | -15.73 | -11.53 |
| 13.50 | 66.85 | 118.59 | 38.50 | -15.39 | -11.35 |
| 14.00 | 59.75 | 109.92 | 39.00 | -14.70 | -10.81 |
| 14.50 | 53.61 | 102.21 | 39.50 | -14.94 | -11.18 |
| 15.00 | 49.30 | 96.32 | 40.00 | -15.41 | -11.78 |
| 15.50 | 47.01 | 92.44 | 40.50 | -14.59 | -11.08 |
| 16.00 | 45.95 | 89.79 | 41.00 | -13.60 | -10.20 |
| 16.50 | 44.04 | 86.29 | 41.50 | -13.46 | -10.18 |
| 17.00 | 41.61 | 82.29 | 42.00 | -13.36 | -10.19 |
| 17.50 | 39.27 | 78.38 | 42.50 | -12.28 | -9.22 |
| 18.00 | 36.96 | 74.51 | 43.00 | -11.10 | -8.15 |
| 18.50 | 34.87 | 70.88 | 43.50 | -10.96 | -8.13 |
| 19.00 | 32.97 | 67.46 | 44.00 | -11.32 | -8.61 |
| 19.50 | 31.09 | 64.08 | 44.50 | -11.53 | -8.95 |
| 20.00 | 29.33 | 60.85 | 45.00 | -11.94 | -9.49 |
| 20.50 | 27.89 | 57.96 | 45.50 | -12.80 | -10.49 |
| 21.00 | 26.35 | 55.01 | 46.00 | -13.75 | -11.60 |
| 21.50 | 24.11 | 51.38 | 46.50 | -14.14 | -12.15 |
| 22.00 | 21.59 | 47.51 | 47.00 | -14.26 | -12.44 |
| 22.50 | 19.81 | 44.42 | 47.50 | -14.75 | -13.12 |
| 23.00 | 18.06 | 41.40 | 48.00 | -15.25 | -13.81 |
| 23.50 | 14.99 | 37.09 | 48.50 | -15.06 | -13.83 |
| 24.00 | 11.78 | 32.69 | 49.00 | -14.93 | -13.92 |
| 24.50 | 10.79 | 30.54 | 49.50 | -16.34 | -15.56 |


| 50.00 | -17.93 | -17.39 |
| :--- | :--- | :--- |
| 50.50 | -18.22 | -17.94 |
| 51.00 | -18.07 | -18.06 |
| 51.50 | -18.45 | -18.72 |
| 52.00 | -19.03 | -19.59 |
| 52.50 | -19.34 | -20.21 |
| 53.00 | -19.19 | -20.37 |
| 53.50 | -18.43 | -19.94 |
| 54.00 | -17.53 | -19.37 |
| 54.50 | -17.29 | -19.47 |
| 55.00 | -17.42 | -19.95 |
| 55.50 | -17.28 | -20.17 |
| 56.00 | -17.00 | -20.25 |
| 56.50 | -16.64 | -20.26 |
| 57.00 | -16.03 | -20.02 |
| 57.50 | -14.87 | -19.23 |
| 58.00 | -13.61 | -18.35 |
| 58.50 | -12.86 | -17.98 |
| 59.00 | -12.44 | -17.93 |
| 59.50 | -12.37 | -18.24 |
| 60.00 | -12.46 | -18.71 |
| 60.50 | -12.51 | -19.13 |
| 61.00 | -12.38 | -19.37 |
| 61.50 | -11.79 | -19.14 |
| 62.00 | -10.88 | -18.59 |
| 62.50 | -9.92 | -17.99 |
| 63.00 | -9.23 | -17.64 |
| 63.50 | -9.17 | -17.92 |
| 64.00 | -9.42 | -18.51 |
| 64.50 | -9.34 | -18.75 |
| 65.00 | -9.14 | -18.87 |
| 65.50 | -9.21 | -19.24 |
| 66.00 | -9.30 | -19.63 |
| 66.50 | -8.92 | -19.53 |
| 67.00 | -8.50 | -19.39 |
| 67.50 | -8.50 | -19.65 |
| 68.00 | -8.82 | -20.23 |
| 68.50 | -9.26 | -20.91 |
| 69.00 | -9.53 | -21.42 |
| 69.50 | -9.29 | -21.40 |
| 70.00 | -8.84 | -21.16 |
| 70.50 | -8.71 | -21.24 |
| 71.00 | -8.73 | -21.45 |
| 71.50 | -9.00 | -21.90 |
| 72.00 | -9.12 | -22.20 |
| 72.50 | -8.62 | -21.86 |
| 73.00 | -7.71 | -21.10 |
| 73.50 | -6.77 | -20.31 |
| 74.00 | -6.07 | -19.75 |
| 74.50 | -5.86 | -19.67 |
| 75.00 | -6.12 | -20.06 |
| 75.50 | -6.72 | -20.77 |
| 76.00 | -7.34 | -21.50 |
| 76.50 | -7.60 | -21.87 |
| 77.00 | -7.61 | -21.97 |


| 77.50 | -7.66 | -22.12 |
| ---: | ---: | ---: |
| 78.00 | -7.55 | -22.09 |
| 78.50 | -6.83 | -21.45 |
| 79.00 | -5.97 | -20.67 |
| 79.50 | -5.87 | -20.64 |
| 80.00 | -5.91 | -20.75 |
| 80.50 | -5.93 | -20.83 |
| 81.00 | -6.02 | -20.98 |
| 81.50 | -6.71 | -21.72 |
| 82.00 | -7.14 | -22.20 |
| 82.50 | -6.80 | -21.90 |
| 83.00 | -6.47 | -21.60 |
| 83.50 | -6.15 | -21.31 |
| 84.00 | -5.83 | -21.02 |
| 84.50 | -6.01 | -21.22 |
| 85.00 | -6.04 | -21.26 |
| 85.50 | -5.17 | -20.39 |
| 86.00 | -4.16 | -19.37 |
| 86.50 | -3.54 | -18.74 |
| 87.00 | -3.09 | -18.26 |
| 87.50 | -2.75 | -17.89 |
| 88.00 | -2.55 | -17.64 |
| 88.50 | -2.57 | -17.60 |
| 89.00 | -2.51 | -17.47 |
| 89.50 | -2.12 | -17.00 |
| 90.00 | -1.78 | -16.56 |
| 90.50 | -1.80 | -16.46 |
| 91.00 | -1.70 | -16.23 |
| 91.50 | -1.10 | -15.48 |
| 92.00 | -0.52 | -14.73 |
| 92.50 | -0.47 | -14.49 |
| 93.00 | -0.25 | -14.07 |
| 93.50 | 0.45 | -13.14 |
| 94.00 | 1.03 | -12.31 |
| 94.50 | 1.19 | -11.88 |
| 95.00 | 1.47 | -11.31 |
| 95.50 | 2.13 | -10.34 |
| 96.00 | 2.83 | -9.30 |
| 96.50 | 3.36 | -8.41 |
| 97.00 | 3.82 | -7.56 |
| 97.50 | 4.24 | -6.73 |
| 98.00 | 4.83 | -5.71 |
| 98.50 | 5.65 | -4.43 |
| 99.00 | 6.56 | -3.03 |
| 99.50 | 7.42 | -1.67 |
| 100.00 | 7.97 | -0.59 |
| 100.50 | 8.07 | 0.06 |
| 101.00 | 8.42 | 0.99 |
| 101.50 | 9.67 | 2.83 |
| 102.00 | 10.61 | 4.39 |
| 102.50 | 10.21 | 4.62 |
| 103.00 | 9.79 | 4.85 |
| 103.50 | 10.04 | 5.77 |
| 104.00 | 10.26 | 6.68 |
| 104.50 | 10.02 | 7.14 |


| 105.00 | 9.38 | 7.21 |
| :---: | :---: | :---: |
| 105.50 | 8.38 | 6.93 |
| 106.00 | 7.64 | 6.92 |
| $\underline{106.50}$ | 7.71 | 7.73 |
| 107.00 | 7.90 | 8.66 |
| 107.50 | 7.82 | 9.33 |
| 108.00 | 7.85 | 10.11 |
| $\therefore 08.50$ | 8.19 | 11.19 |
| -09.00 | 8.60 | 12.34 |
| .09 .50 | 8.92 | 13.40 |
| 110.00 | 9.20 | 14.41 |
| $\geq 10.50$ | 9.28 | 15.21 |
| $\vdots 11.00$ | 9.61 | 16.24 |
| 111.50 | 10.34 | 17.67 |
| 112.00 | 10.95 | 18.95 |
| 112.50 | 11.10 | 19.76 |
| 113.00 | 10.80 | 20.09 |
| 113.50 | 10.20 | 20.10 |
| 114.00 | 9.87 | 20.36 |
| 114.50 | 10.31 | 21.36 |
| 115.00 | 10.51 | 22.09 |
| $\pm 15.50$ | 9.69 | 21.76 |
| 116.00 | 8.95 | 21.49 |
| 116.50 | 8.85 | 21.82 |
| 117.00 | 8.69 | 22.06 |
| 117.50 | 8.01 | 21.74 |
| 118.00 | 7.38 | 21.43 |
| 118.50 | 6.63 | 20.96 |
| 119.00 | 6.24 | 20.82 |
| 119.50 | 6.58 | 21.36 |
| 120.00 | 6.96 | 21.90 |
| 120.50 | 7.53 | 22.59 |
| 121.00 | 7.87 | 23.02 |
| 121.50 | 7.80 | 22.99 |
| 122.00 | 7.26 | 22.45 |
| 122.50 | 6.29 | 21.44 |
| 123.00 | 5.32 | 20.38 |
| 123.50 | 4.62 | 19.57 |
| 124.00 | 4.55 | 19.34 |
| 124.50 | 5.22 | 19.81 |
| 125.00 | 5.63 | 20.00 |
| 125.50 | 5.11 | 19.21 |
| 126.00 | 4.67 | 18.47 |
| 126.50 | 4.86 | 18.34 |
| 127.00 | 5.34 | 18.46 |
| 127.50 | 5.89 | 18.63 |
| 128.00 | 6.64 | 18.97 |
| 128.50 | 7.62 | 19.52 |
| 129.00 | 8.77 | 20. 21 |
| 129.50 | 10.03 | 21.00 |
| 130.00 | 10.83 | 21.32 |
| 130.50 | 10.78 | 20.76 |
| 131.00 | 10.11 | 19.58 |
| 131.50 | 9.12 | 18.07 |
| $\therefore 32.00$ | 8.14 | 16.56 |


| 132.50 | 7.45 | 15.34 |
| :--- | ---: | ---: |
| 133.00 | 7.40 | 14.76 |
| 133.50 | 8.11 | 14.93 |
| 134.00 | 8.88 | 15.17 |
| 134.50 | 8.98 | 14.75 |
| 135.00 | 8.83 | 14.08 |
| 135.50 | 8.89 | 13.64 |
| 136.00 | 8.15 | 12.40 |
| 136.50 | 6.01 | 9.78 |
| 137.00 | 4.15 | 7.46 |
| 137.50 | 3.76 | 6.62 |
| 138.00 | 3.65 | 6.08 |
| 138.50 | 3.13 | 5.15 |
| 139.00 | 2.63 | 4.26 |
| 139.50 | 2.36 | 3.63 |
| 140.00 | 1.84 | 2.77 |
| 140.50 | 0.87 | 1.48 |
| 141.00 | 0.18 | 0.50 |
| 141.50 | 0.22 | 0.27 |
| 142.00 | 0.23 | 0.04 |
| 142.50 | 0.28 | -0.13 |
| 143.00 | 0.12 | -0.48 |
| 143.50 | -0.35 | -1.12 |
| 144.00 | -0.87 | -1.78 |
| 144.50 | -1.24 | -2.27 |
| 145.00 | -1.76 | -2.88 |
| 145.50 | -2.56 | -3.75 |
| 146.00 | -3.42 | -4.66 |
| 146.50 | -4.20 | -5.48 |
| 147.00 | -4.92 | -6.21 |
| 147.50 | -5.60 | -6.88 |
| 148.00 | -5.99 | -7.25 |
| 148.50 | -5.94 | -7.17 |
| 149.00 | -6.01 | -7.19 |
| 149.50 | -6.67 | -7.79 |
| 150.00 | -6.72 | -7.77 |
| 150.50 | -5.28 | -6.26 |
| 151.00 | -3.91 | -4.81 |
| 151.50 | -3.45 | -4.27 |
| 152.00 | -3.48 | -4.27 |
| 152.50 | -3.75 | -4.40 |
| 153.00 | -3.97 | -4.54 |
| 153.50 | -4.01 | -4.51 |
| 154.00 | -4.55 | -4.98 |
| 154.50 | -6.07 | -6.44 |
| 155.00 | -6.94 | -7.26 |
| 155.50 | -5.86 | -6.15 |
| 156.00 | -4.85 | -5.11 |
| 156.50 | -4.94 | -5.19 |
| 157.00 | -4.74 | -5.00 |
| 158.00 | -3.69 | -3.97 |
| 159.50 | -6.48 | -7.04 |


| 160.00 | -6.07 | -6.75 |
| :--- | ---: | ---: |
| 160.50 | -6.19 | -7.02 |
| 161.00 | -7.31 | -8.30 |
| 161.50 | -9.38 | -10.55 |
| 162.00 | -10.89 | -12.27 |
| 162.50 | -10.66 | -12.26 |
| 163.00 | -10.96 | -12.81 |
| 163.50 | -13.54 | -15.66 |
| 164.00 | -15.49 | -17.89 |
| 164.50 | -15.45 | -18.15 |
| 165.00 | -14.70 | -17.72 |
| 165.50 | -14.05 | -17.40 |
| 166.00 | -14.10 | -17.80 |
| 166.50 | -14.96 | -19.01 |
| 167.00 | -15.07 | -19.49 |
| 167.50 | -13.53 | -18.33 |
| 168.00 | -12.14 | -17.32 |
| 168.50 | -12.07 | -17.64 |
| 169.00 | -12.91 | -18.88 |
| 169.50 | -14.92 | -21.28 |
| 170.00 | -17.18 | -23.93 |
| 170.50 | -19.16 | -26.30 |
| 171.00 | -19.96 | -27.49 |
| 171.50 | -19.48 | -27.39 |
| 172.00 | -19.95 | -28.23 |
| 172.50 | -23.66 | -32.30 |
| 173.00 | -27.41 | -36.39 |
| 173.50 | -29.62 | -38.93 |
| 174.00 | -30.53 | -40.16 |
| 174.50 | -30.66 | -40.59 |
| 175.00 | -30.88 | -41.08 |
| 175.50 | -31.99 | -42.45 |
| 176.00 | -33.69 | -44.38 |
| 176.50 | -36.24 | -47.14 |
| 177.00 | -40.40 | -51.48 |
| 177.50 | -47.02 | -58.26 |
| 178.00 | -54.21 | -65.58 |
| 178.50 | -54.64 | -66.11 |
| 179.00 | -41.72 | -53.26 |
| 179.50 | -37.87 | -49.46 |
| 180.00 | -72.83 | -84.43 |
|  |  |  |
| 17 |  |  |

Table Two
Smoothed Anomaly Degree Variances ( $\mathrm{c}_{\ell}$ ) As
Computed With $1^{\circ}$ Free Air Anomalies
and From a Current Potential Coefficient Set
$\left(\mathrm{mgal}^{2}\right)$

| Degree | from $1^{\circ}{ }^{\mathrm{c}_{\ell}}$ anomaly data | $\mathrm{c}_{\ell}$ <br> from potential coefficients <br> (Rapp, 1973b) |
| :---: | :---: | :---: |
| 0 |  | -- |
| 1 | 0.07 | -- |
| 2 | 2.3 | 7.5 |
| 3 | 26.2 | 33.9 |
| 4 | 58.3 | 19.2 |
| 5 | 16.0 | 21.6 |
| 6 | 26.3 | 18.9 |
| 7 | 36.0 | 18.8 |
| 8 | 22.8 | 10.4 |
| 9 | 12.6 | 11.1 |
| 10 | 20.0 | 11.4 |

If the gravity formula were that of a mean earth ellipsoid, the zeroth degree variance should be zero. This is essentially the case here with the fact that the $\gamma_{\theta}$ and the flattening of the GRS67 are quite close to be current best estimates of these parameters (Rapp, 1974). The anomalies taken on a global scale should have no first degree anomaly degree variance. The non-global $1^{\circ}$ anomalies that we have imply through the covariance function a small one of $2.3 \mathrm{mgal}^{2}$.

The anomaly degree variances from the potential coefficients should be reliable at the lower degrees because of the accurate determination of low degree potential coefficients through satellite orbital analysis. Comparison of these values with that implied by the covariance function indicates poor agreement for degrees $2,3,6$ and 9 . This disagreement may be related to the fact that the $1^{\circ}$ anomalies cover only $50 \%$ of the earth's surface and we cannot hope to find good low degree information from such limited coverage.

However, for future analysis it is important that we use a $1^{\circ}$ covariance function that is characteristic of the real world especially at low degrees. To develop such a covariance function we modify the covariance function computed from the anomalies by imposing on the modified function the $c_{\ell}$ values to degree 10 as listed in Table Two
(as computed from potential coefficients). To do this we first remove the effect of the $\overline{\mathrm{c}}_{\ell}$ values listed in Table Two and then add back the covariance contribution from the $c_{2}$ values, in both cases using equation (11) setting $R_{l}$ and $s$ equal to one. In effect we carry out the following modification to obtain a modified $1^{\circ}$ covariance function:

The modified covariance function is shown in Table One being labeled Modified $\bar{C}(\psi)$. This modified covariance function is plotted in Figure One.

Smoothed anomaly degree variances were developed from this modified covariance function where were then converted to the actual degree variances using equation (16A). These results and values of $\beta_{\ell}$ for one degree blocks and $s^{-(l+2)}$ are given in Table B of the appendix.

From Table One, using the modified covariance function of the current estimate for the variance of a $1^{\circ}$ anomaly is $919.66 \mathrm{mgal}^{2}$, or a root mean square value of $\pm 30.3$ mgals with respect to the gravity formula of the Geodetic Reference System 1967.

## 4. A Five Degree Anomaly Variance

For purposes of obtaining models of anomaly degree variance using procedures such as described in Rapp (1973a) we need to estimate the variance of the $5^{\circ}$ anomalies. This can be done in two ways. The first procedure is by the numercal integration of the $1^{\circ}$ modified covariance function according to equation (7-82) of Heiskanen and Moritz ( p .270 ). This leads to an estimate of $305 \mathrm{mgal}^{2}$. The second procedure is to compute the variance directly from the $5^{\circ}$ anomalies. This was done by first predicting $5^{\circ}$ equal area anomalies using the methods described in Rapp (1972) but with the more current $1^{\circ} \times 1^{\circ}$ set. The variance computed by this procedure from the 1354 predicted anomalies was $298 \mathrm{mgal}^{2}$. We adopt for further use the variance of 5 degree anomalies as $302 \mathrm{mgal}^{2}$ with respect to the gravity formula of the Geodetic Reference System 1967.

## 5. The Point Anomaly Variance

The value of $\mathrm{C}_{0}$ is an important quantity as it is a scaling factor for many representations of the point anomaly covariance function. $C_{0}$ has been treated as both a local or regional quantity, or a global quantity. On a regional basis $\mathrm{C}_{0}$ is the variance of the point anomalies in some defined area. Thus, it will change from area to area. The global $C_{0}$ value is considered to be representative of the gravity field of the whole earth. The estimation of $\mathrm{C}_{0}$ on a giobal bas is is not straight
forward since we do not have global gravity coverage. The only global point covariance function numerically estimated is that given by Kaula (1959) where he used gravity data that was current to 1958. During the 16 years since the compliation of gravity data as used by Kaula, a considerable amount of additional data has become available. Thus, a new computation of global point covariance seems appropriate and is needed. Such a computation can only be done through some organization that has access to the gravity data holdings. For this report we do not have the facilities or funds to carry out a computation of a point covariance function. However, we can use several procedures to determine $\mathrm{C}_{0}$, the quantity so fundamental to the analytical representation of a point covariance function.

### 5.1 Method One

One method to estimate $\mathrm{C}_{0}$ is to consider the relationship between a point covariance function ( $\mathrm{C}(\mathrm{d})$ ) and the variance $\left(\mathrm{G}_{\mathrm{so}}^{2}\right)$ of anomalies given in blocks of size $s^{\circ}$. One convenient relatıonsh ip is given by Hirvonen (1962):

$$
\begin{equation*}
\mathrm{G}_{\mathrm{s}}^{2} \mathrm{c}=\int_{0}^{\sqrt{2}} \mathrm{WC}(\mathrm{~d}) \mathrm{dr} \tag{19}
\end{equation*}
$$

where

$$
\begin{aligned}
& \mathrm{d}=\mathrm{rs}^{\circ} \\
& \mathrm{W}=\left(2 \pi-8 \mathrm{r}+2 \mathrm{r}^{2}\right) \mathrm{r} \text { when } 0<\mathrm{r}<1 \\
& \mathrm{~W}=\left(2 \pi-4-2 \mathrm{r}^{2}+8 \sqrt{\mathrm{r}^{2}-1}-8 \tan ^{-1} \sqrt{\mathrm{r}^{2}-1}\right) \mathrm{r} \text { where } 1<\mathrm{r}<\sqrt{2}
\end{aligned}
$$

If we represent $C(d)$ in the form of:

$$
\begin{equation*}
C(d)=C_{0} f(d) \tag{20}
\end{equation*}
$$

we can solve (19) and (20) for $\mathrm{C}_{0}$ :

$$
\begin{equation*}
\mathrm{C}_{0}=\frac{\mathrm{G}_{8^{\circ}}^{2}}{\int_{0}^{\sqrt{2}} \mathrm{Wf}(\mathrm{~d}) \mathrm{dr}} \equiv \frac{\mathrm{G}_{\mathrm{g}^{\circ}}^{2}}{\mathrm{I}} \tag{21}
\end{equation*}
$$

The value of I can be obtained for various representative $f(d)$.
Many representations of the point covariance function have been suggested. Many of these representations are summarized in papers by Groten (1966), Lauer (1971), and Jordan (1972). For the purposes of this paper we have used three models. These are:

$$
\begin{align*}
& \text { (1) } C(d)=C_{0} \underbrace{e^{-o d}}_{f_{1}(d)} \\
& \text { (2) } C(d)=C_{0} \underbrace{\left(1-\frac{d}{2 c_{2}}\right) e^{-d / c_{2}}}_{f_{2}(d)}  \tag{23}\\
& \text { (3) } C(d)=C_{0}\left(1+d\left(a_{1}+d\left(a_{2}+d\left(a_{3}+d\left(a_{4}+d\left(a_{5}\right)\right)\right)\right)\right)\right) \tag{24}
\end{align*}
$$

The $c_{1}$ and $c_{2}$ values were obtained from fitting the Kaula (1959) point covariance curve to a distance of $1.5^{\circ}$. We found $c_{2}=0^{\circ} .897$ and $c_{2}=1^{\circ} .88$. Beyond a distance of $1.5^{\circ}$, the point covariance would not be represented well by equations (22) and (23) with the above constants. The constants of equation (24) were obtained by a least squares polynomial fit using the Kaula point covariance function to $8^{\circ}$. We found:

$$
\begin{aligned}
& a_{1}=-.9816195 \\
& a_{2}=.4894498 \\
& a_{3}=-.1149583 \\
& a_{4}=.0126057 \\
& a_{5}=-.000523222
\end{aligned}
$$

For these models, the root mean square fit to the observed covariance function was $\pm 30 \mathrm{mgal}^{2}, \pm 75 \mathrm{mgal}^{2}$, and $\pm 28 \mathrm{mgal}^{2}$ for models 1,2 and 3 respectively. For $\mathrm{s}^{\circ}=1^{\circ}$, values of $I$ (computed by numerical integration), and $C_{0}$ (taking $G_{8}^{\&}=919.66$ from Table One) are given for each of the models in Table Three.

| Table Three |  |  |
| :---: | :---: | :---: |
| Estimation of Co from $1^{\circ}$ |  |  |
| Anomaly Block Variances |  |  |
| Model | I | $\mathrm{C}_{0}$ |
| 1, equation (22) | .64185 | $1433 \mathrm{mgal}^{2}$ |
| 2, equation (23) | .66491 | $1383 \mathrm{mgal}^{2}$ |
| 3, equation (24) | .62595 | $1469 \mathrm{mgal}^{2}$ |

Using weights based on the root mean square fits to the point covariance curve, the estimated $C_{0}$ from this analys is is $1447 \mathrm{mgal}^{2}$.

### 5.2 Method Two

A more direct method for determining $\mathrm{C}_{0}$ is through the analysis of the actual point gravity anomalies. Such an analysis is not a straight forward one since the anomaly data is not uniformly distributed over the earth. Since certain areas (such as land areas) have, in general, denser anomaly coverage than ocean areas, and since free-air anomalies are correlated with land elevations or ocean depth, special care needs to be taken in the analys is of a set of point gravity anomalies for $\mathrm{C}_{0}$.

In our analys is we basically considered a point variance by elevation range, and then converted these individual variances into a global estimate of $\mathrm{C}_{0}$ by forming a weighted mean with weights being based on the percentage of the earth's surface lying within the elevation range.

As the first step in this procedure the Defense Mapping Agency Aerospace Center considered a set of $2,253,122$ point free-air anomalies whose elevation or depth was known. Elevation ranges of 100 meter increment were chosen. For all anomalies falling within each range, the mean anomaly, the mean square anomaly and the mean elevation from the points, was determined. The mean square anomaly was computed as the sum of the square of the anomalies with the elevation range divided by the number of anomalies within the range. In subsequent discussions this quantity will be referred to as the variance of the range. This terminology is not specifically correct as a variance is usually defined with respect to a quantity whose mean is zero. In fact, the anomaly mean within a range will not be zero, but it will be zero or close to it on a global basis. This data by ranges is shown in Table Four.

In order to form a global estimate of $\mathrm{C}_{0}$, we now need to know how elevations are distributed on the actual earth. To do this we considered mean elevations in 1654 $5^{\circ}$ equal area blocks and $64800,1^{\circ} \times 1^{\circ}$ mean elevations. From this data the preentage of the earth's surface within a given elevation range could be found. The results found for the $5^{\circ}$ and $1^{\circ}$ data are shown as the last two columns in Table Four. The $5^{\circ}$ results are shown as a matter of interest only, as the $5^{\circ}$ subdivision is too large for the purposes needed here. We should note that all 0.0 's given in Table Four with the exception of the mean anomaly for the 100 to 200 meter range indicate no data was available for the quantity. The $1^{\circ}$ subdivision is also not sufficiently small for the most accurate work as can be seen from the fact that certain elevation ranges for which there was point elevations data were not represented in the data from the $1^{0}$ mean elevation data.

The weighted variance (or $\mathrm{C}_{0}$ ) was then determined as follows:

$$
\begin{equation*}
C_{0}=\frac{\sum_{1} P_{1}\left(C_{0}\right)_{1}}{\sum_{1} P_{1}} \tag{25}
\end{equation*}
$$

Table Four
Anomaly Variance and Related Information by Elevation Range

| (meters) |  |
| :---: | :---: |
| -14100 | -14000 |
| -11200 | -11100 |
| -10800 | -10700 |
| -10700 | -10600 |
| -10600 | -10500 |
| -10500 | - 10400 |
| -10400 | -10300 |
| -10300 | -10200 |
| -10200 | -10100 |
| -10100 | -10000 |
| -10000 | -9900 |
| -9900 | -9800 |
| -9800 | -9700 |
| -9700 | -9600 |
| -9600 | -9500 |
| -9500 | -9400 |
| -9400 | -9300 |
| -9300 | -9200 |
| -9200 | -9100 |
| -9100 | -9000 |
| -9000 | -8900 |
| -8900 | -8800 |
| -8800 | -8700 |
| -8700 | -8600 |
| -8600 | -8500 |
| -8500 | -8400 |
| -8400 | -8300 |
| -8300 | -8200 |
| -8200 | -8100 |
| -8100 | -8000 |
| -8000 | -7900 |
| -7900 | -7800 |
| -7800 | -7700 |
| -7700 | -7600 |
| -7600 | -7500 |
| -7500 | -7400 |
| -7400 | -7300 |
| -7300 | -7200 |
| -7200 | -7100 |
| -7100 | -7000 |
| -7000 | -6900 |
| -6900 | -6800 |
| -6800 | -6700 |
| -6700 | -6600 |
| -6600 | -6500 |
| -6500 | -6400 |
| -6400 | -6300 |
| -6300 | -6200 |
| -6200 | -6100 |
| -6100 | -6000 |

Point anom.
No. of point Anomalies

Mean sq. (mgals) (mal ${ }^{2}$ ) $0 \quad 0.0 \quad 0.0$ 5 1 2 13
$-213.1$
$-277.3$
$-285.0$
$-282.4$
$-270.6$
$-290.3$
$-283.4$
$-282.3$
$-279.8$
$-276.4$
$-273.5$
$-260.5 \quad 70629.0$
$-267.7 \quad 73083.8$
$-248.864161 .0$
$-241.9 \quad 63265.9$
$-259.2 \quad 68675.0$
-231.9 56997.3
$-243.2 \quad 61142.7$
$-236.8 \quad 57878.9$
$-242.4 \quad 61454.6$
-222.6 52069.2
$-227.8$
$-222.1$
$\begin{array}{ll}-225.5 & 52099.5 \\ -249.9 & 63345.2\end{array}$
$\begin{array}{ll}-249.9 & 63345.2 \\ -220.6 & 50153.0\end{array}$
$-214.7 \quad 47255.5$
$-221.0 \quad 51087.8$
$-231.6 \quad 57988.7$
-211.2 47870.1
$-207.7 \quad 46967.6$
$-204.2 \quad 45255.4$
$-205.3 \quad 46034.8$
-214.8 51871.4
$-180.2 \quad 35771.0$
$-168.6$
$-158.3$
31597.3
28832.5
38391.0
27775.9
30243.8
21597.5
18588.8
18717.8
15090.3
12701.3
9847.8
6841.8
4002.0

Average of
Percentage of earth's pt. elevations surface within range

| (meters) | $1^{\circ}$ data | 5 data |
| :---: | :---: | :---: |
| 0.0 | 0.002 | 0.0 |
| -11113.0 | 0.0 | 0.0 |
| -10750.6 | 0.0 | 0.0 |
| -10674.0 | 0.0 | 0.0 |
| -10592.0 | 0.0 | 0.0 |
| -10425.2 | 0.0 | 0.0 |
| -10353.8 | 0.0 | 0.0 |
| -10228.4 | 0.0 | 0.0 |
| -10149.4 | 0.0 | 0.0 |
| -10065.4 | 0.0 | 0.0 |
| -9947.7 | 0.0 | 0.0 |
| -9858.6 | 0.0 | 0.0 |
| -9743.8 | 0.0 | 0.0 |
| -9645.9 | 0.0 | 0.0 |
| -9546.7 | 0.0 | 0.0 |
| -9446.1 | 0.0 | 0.0 |
| -9347.8 | 0.0 | 0.0 |
| -9247.8 | 0.0 | 0.0 |
| -9147.5 | 0.0 | 0.0 |
| -9041.5 | 0.0 | 0.0 |
| -8956.7 | 0.0 | 0.0 |
| -8853.9 | 0.0 | 0.0 |
| -8764.4 | 0.0 | 0.0 |
| -8667.9 | 0.0 | 0.0 |
| -8551.8 | 0.0 | 0.0 |
| -8440.7 | 0.0 | 0.0 |
| -8360.1 | 0.0 | 0.0 |
| -8259.2 | 0.0 | 0.0 |
| -8148.5 | 0.0 | 0.0 |
| -8048.0 | 0.0 | 0.0 |
| -7952.1 | 0.0 | 0.0 |
| -7856.2 | 0.002 | 0.0 |
| -7759.5 | 0.0 | 0.0 |
| -7654.5 | 0.002 | 0.0 |
| -7551.2 | 0.002 | 0.0 |
| -7446.8 | 0.0 | 0.0 |
| -7337.9 | 0.008 | 0.0 |
| -7250.2 | 0.003 | 0.0 |
| -7151.5 | 0.002 | 0.0 |
| -7054.5 | 0.014 | 0.0 |
| -6949.7 | 0.007 | 0.0 |
| -6851.3 | 0.002 | 0.0 |
| -6744.5 | 0.005 | 0.0 |
| -6658.1 | 0.009 | 0.0 |
| -6546.2 | 0.024 | 0.0 |
| -6454.0 | 0.010 | 0.0 |
| -6349.6 | 0.027 | 0.0 |
| -6246.4 | 0.054 | 0.0 |
| -6146.7 | 0.153 | 0.0 |
| -6049.2 | 0.321 | 0.0 |


| -6000 | -5900 | 4024 | -27.3 | 3215.6 | -5946.8 | 0.292 | 0.120 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| -5900 | -5800 | 7166 | -22.5 | 1786.2 | -5847.2 | 0.502 | 0.181 |
| -5800 | -5700 | 8170 | -17.8 | 1319.7 | -5750.8 | 0.697 | 0.403 |
| -5700 | -5600 | 9542 | -14.2 | 1287.4 | -5652.4 | 0.983 | 0.562 |
| -5600 | -5500 | 10061 | -13.0 | 1143.5 | -5550.1 | 1.449 | 0.833 |
| -5500 | -5400 | 12429 | $-11.7$ | 1106.4 | -5450.9 | 1.170 | 1.059 |
| -5400 | -5300 | 14512 | -12.0 | 1113.8 | -5350.6 | 1.478 | 1.688 |
| -5300 | -5200 | 14791 | -12.2 | 1245.1 | -5248.2 | 2.130 | 1.620 |
| -5200 | -5100 | 15930 | -13.6 | 1243.7 | -5147.5 | 3.033 | 2.111 |
| -5100 | -5000 | 16687 | $-12.7$ | 1231.6 | -5051.1 | 2.766 | 1.738 |
| -5000 | -4900 | 16185 | -11.9 | 1370.6 | -4950.5 | 2.222 | 2.816 |
| -4900 | -4800 | 16178 | -14.6 | 1401.8 | -4849.1 | 2.395 | 2.085 |
| -4800 | -4700 | 16412 | -13.9 | 1350.3 | $-4750.2$ | 1.815 | 2.459 |
| -4700 | -4600 | 16153 | -13.2 | 1433.0 | -4651.0 | 2.027 | 2. 300 |
| -4600 | -4500 | 16000 | $-12.2$ | 1243.0 | -4551.8 | 2.115 | 1.613 |
| -4500 | -4400 | 14261 | $-12.3$ | 1431.1 | -4452.9 | 2.102 | 2.059 |
| -4400 | -4300 | 12350 | -10.6 | 1550.4 | -4352.9 | 2.013 | 2.866 |
| -4300 | -4200 | 12568 | $-12.1$ | 1588.8 | -4251.1 | 2.578 | 2.051 |
| -4200 | -4100 | 11377 | -15.1 | 2402.4 | -4149.5 | 2.366 | 2.514 |
| -4100 | -4000 | 11111 | -11.0 | 1709.1 | -4050.4 | 2.435 | 2.053 |
| -4000 | -3900 | 11122 | -8.7 | 1476.6 | -3949.8 | 1.649 | 2.129 |
| -3900 | -3800 | 10887 | -10.3 | 1705.2 | -3851.0 | 1.930 | 2. 204 |
| -3800 | -3700 | 9948 | -7.6 | 1732.4 | -3751.4 | 1.626 | 1.929 |
| -3700 | -3600 | 9982 | -7.4 | 1861.9 | -3650.7 | 1.404 | 2.398 |
| -3600 | -3500 | 10272 | -7.9 | 2086.7 | -3550.7 | 1.647 | 1.299 |
| -3500 | -3400 | 9574 | -6. 8 | 2061.3 | -3451.0 | 1.282 | 2.092 |
| -3400 | -3300 | 9899 | -7.1 | 1974.4 | -3350.6 | 1.220 | 1.449 |
| -3300 | -3200 | 11356 | -5.0 | 1920.2 | -3250.4 | 1.339 | 1.818 |
| -3200 | -3100 | 11342 | -5.0 | 1884.6 | -3151.1 | 1.219 | 1.167 |
| -3100 | -3000 | 11490 | -7.6 | 2039.3 | -3050.1 | 1.673 | 1.091 |
| -3000 | -2900 | 12194 | $-10.2$ | 2075.9 | -2950.7 | 0.829 | 0.914 |
| -2900 | -2800 | 13157 | -10.6 | 2138.0 | -2848.7 | 0.765 | 0.967 |
| -2800 | -2700 | 13830 | -9.4 | 2080.9 | -2748.9 | 0.874 | 0.714 |
| -2700 | -2600 | 13583 | $-12.9$ | 2353.4 | -2651.9 | 0.453 | 0.837 |
| -2600 | -2500 | 19490 | -2.2 | 2107.7 | -2544.1 | 0.570 | 0.558 |
| -2500 | -2400 | 10858 | -17.7 | 3254.9 | -2451.0 | 0.909 | 1.047 |
| -2400 | -2300 | 10449 | -15.4 | 3024.0 | $-2351.8$ | 0.421 | 0.789 |
| -2300 | -2200 | 9355 | -15.0 | 3135.1 | -2251.8 | 0.439 | 1.332 |
| -2200 | -2100 | 12206 | -9.5 | 2133.6 | -2147.7 | 1.018 | 0.543 |
| -2100 | -2000 | 12676 | -2.8 | 1803.5 | -2048.7 | 0.559 | 0.615 |
| -2000 | -1900 | 12963 | 1.2 | 1713.2 | -1951.6 | 0.288 | 0.538 |
| -1900 | -1800 | 11371 | -0.3 | 2125.1 | $-1851.6$ | 0.682 | 0.322 |
| -1800 | -1700 | 9321 | 0.0 | 2522.1 | -1751.1 | 0.330 | 0.659 |
| -1700 | -1600 | 9475 | 5.2 | 2308.1 | -1649.8 | 0.302 | 0.438 |
| -1600 | -1500 | 10452 | 6.1 | 1964.9 | -1549.9 | 0.563 | 0.790 |
| -1500 | -1400 | 10102 | 4.9 | 2023.2 | -1453.0 | 0.241 | 0.673 |
| -1400 | -1300 | 10371 | 5.4 | 2321.7 | -1350.6 | 0.252 | 0.547 |
| -1300 | -1200 | 10581 | 3.8 | 2407.0 | -1252.6 | 0.504 | 0.721 |
| -1200 | -1100 | 9917 | 6.1 | 2397.7 | -1151.8 | 0.251 | 0.431 |
| -1100 | -1000 | 9722 | 7.8 | 2246.5 | -1051.8 | 0.439 | 0.892 |
| -1000 | -900 | 9816 | 11.5 | 2221.2 | -949.4 | 0.383 | 0.766 |
| -900 | -800 | 11381 | 9.2 | 2640.5 | -848.5 | 0.331 | 0.720 |
| -800 | -700 | 10731 | 11.2 | 2043.0 | -752.6 | 0.288 | 0.661 |
| -700 | -600 | 9038 | 14.2 | 2294.5 | -650.6 | 0.400 | 0.570 |
| -600 | -500 | 10338 | 16.9 | 2636.4 | -547.9 | 0.459 | 0.604 |


| -500 | -400 | 12816 | 15.4 | 2470.5 | -447.7 | 0.394 | 0.779 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| -400 | -300 | 16341 | 12.9 | 2490.7 | -348.5 | 0.696 | 0.371 |
| -300 | -200 | 19910 | 12.9 | 1938.9 | -249.7 | 0.772 | 0.528 |
| -200 | -100 | 37357 | 13.6 | 1756.5 | -141.3 | 1.073 | 1.091 |
| -100 | 0 | 85482 | 8.7 | 1713.2 | -49.1 | 3.151 | 1.925 |
| 0 | 100 | 404177 | 3.1 | 1345.0 | 40.6 | 3.557 | 2.825 |
| 100 | 200 | 227862 | 0.0 | 807.1 | 147.8 | 3.961 | 3.477 |
| 200 | 300 | 172980 | -0.6 | 801.3 | 245.2 | 3.431 | 3.921 |
| 300 | 400 | 106121 | -0.8 | 970.9 | 347.6 | 2.960 | 2.447 |
| 400 | 500 | 86419 | 0.5 | 1054.6 | 448.2 | 2.424 | 2.855 |
| 500 | 600 | 51225 | 3.3 | 1345.5 | 546.7 | 1.812 | 2.197 |
| 600 | 700 | 35994 | 1.4 | 1580.4 | 647.9 | 1.497 | 0.990 |
| 700 | 800 | 29210 | -2. 2 | 1654.3 | 748.6 | 1.210 | 1.272 |
| 800 | 900 | 26750 | 1.8 | 1540.9 | 349.5 | 1.109 | 0.717 |
| 900 | 1000 | 23329 | 2.4 | 1540.5 | 948.9 | 1.054 | 1.277 |
| 1000 | 1100 | 23078 | 5.6 | 1416.9 | 1048.9 | 0.889 | 0.898 |
| 1100 | 1200 | 26348 | -0.6 | 1193.7 | 1154.8 | 0.773 | 0.734 |
| 1200 | 1300 | 26176 | 0.7 | 1214.8 | 1251.9 | 0.689 | 0.559 |
| 1300 | 1400 | 30036 | $-1.2$ | 930.2 | 1348.5 | 0.506 | 0.538 |
| 1400 | 1500 | 23156 | 1.4 | 1165.4 | 1448.3 | 0.451 | 0.234 |
| 1500 | 1600 | 17911 | 1.6 | 1557.4 | 1548.4 | 0.368 | 0.058 |
| 1600 | 1700 | 15296 | 3.1 | 1671.9 | 1648.0 | 0.272 | 0.443 |
| 1700 | 1800 | 12868 | 7.3 | 1610.2 | 1749.1 | 0.238 | 0.129 |
| 1800 | 1900 | 11550 | 10.5 | 1842.9 | 1849.7 | 0.219 | 0.061 |
| 1900 | 2000 | 12138 | 12.3 | 1683.3 | 1951.7 | 0.183 | 0.183 |
| 2000 | 2100 | 13163 | 12.2 | 1638.8 | 2049.5 | 0.171 | 0.173 |
| 2100 | 2200 | 10544 | 19.1 | 1886.8 | 2146.8 | 0.153 | 0.058 |
| 2200 | 2300 | 8208 | 33.7 | 2766.9 | 2247.9 | 0.091 | 0.183 |
| 2300 | 2400 | 4939 | 37.4 | 3644.2 | 2346.1 | 0.078 | 0.0 |
| 2400 | 2500 | 4006 | 42.5 | 4236.7 | 2450.8 | 0.062 | 0.0 |
| 2500 | 2600 | 3547 | 50.8 | 5041.9 | 2547.8 | 0.068 | 0.0 |
| 2600 | 2700 | 2661 | 55.4 | 5229.0 | 2647.4 | 0.075 | 0.061 |
| 2700 | 2800 | 2150 | 57.5 | 6384.3 | 2748.3 | 0.044 | 0.058 |
| 2800 | 2900 | 1721 | 58.3 | 7472.7 | 2846.7 | 0.033 | 0.067 |
| 2900 | 3000 | 1331 | 74.4 | 8846.3 | 2947.7 | 0.035 | 0.058 |
| 3000 | 3100 | 1098 | 78.4 | 10269.7 | 3048.9 | 0.037 | 0.067 |
| 3100 | 3200 | 869 | 87.3 | 11715.8 | 3147.4 | 0.030 | 0.0 |
| 3200 | 3300 | 771 | 88.8 | 11760.3 | 3249.9 | 0.022 | 0.0 |
| 3300 | 3400 | 654 | 94.5 | 13293.2 | 3348.9 | 0.033 | 0.0 |
| 3400 | 3500 | 596 | 80.5 | 11545.6 | 3449.6 | 0.026 | 0.125 |
| 3500 | 3600 | 362 | 103.0 | 16509.2 | 3549.2 | 0.022 | 0.0 |
| 3600 | 3700 | 585 | 82.6 | 10158.4 | 3660.7 | 0.027 | 0.0 |
| 3700 | 3800 | 566 | 91.5 | 11237.4 | 3743.8 | 0.022 | 0.0 |
| 3800 | 3900 | 680 | 93.5 | 10670.7 | 3844.9 | 0.031 | 0.0 |
| 3900 | 4000 | 406 | 102.6 | 15102.7 | 3944.7 | 0.028 | 0.0 |
| 4000 | 4100 | 281 | 105.8 | 13003.0 | 4052.3 | 0.040 | 0.0 |
| 4100 | 4200 | 234 | 117.9 | 16503.9 | 4148.8 | 0.034 | 0.061 |
| 4200 | 4300 | 149 | 134.6 | 23227.6 | 4242.6 | 0.024 | 0.0 |
| 4300 | 4400 | 208 | 151.4 | 26309.9 | 4344.9 | 0.024 | 0.0 |
| 4400 | 4500 | 136 | 114.6 | 16198.7 | 4447.5 | 0.029 | 0.061 |
| 4500 | 4600 | 101 | 137.4 | 20794.3 | 4548.0 | 0.031 | 0.0 |
| 4600 | 4700 | 87 | 149.5 | 24966.3 | 4637.7 | 0.027 | 0.0 |
| 4700 | 4800 | 20 | 163.6 | 29051.9 | 4736.6 | 0.034 | 0.0 |
| 4800 | 4900 | 8 | 198.3 | 43451.1 | 4834.6 | 0.029 | 0.051 |
| 4900 | 5000 | 4 | 111.5 | 46485.9 | 4961.1 | 0.037 | 0.0 |


| 5000 | 5100 | 1 | 252.5 | 63756.2 | 5018.6 | 0.037 | 0.0 |
| ---: | ---: | ---: | ---: | ---: | ---: | ---: | ---: |
| 5100 | 5200 | 4 | 82.2 | 18374.4 | 5163.8 | 0.028 | 0.061 |
| 5200 | 5300 | 1 | 268.4 | 72038.6 | 5235.8 | 0.037 | 0.0 |
| 5300 | 5400 | 0 | 0.0 | 0.0 | 0.0 | 0.020 | 0.0 |
| 5400 | 5500 | 0 | 0.0 | 0.0 | 0.0 | 0.016 | 0.0 |
| 5500 | 5600 | 0 | 0.0 | 0.0 | 0.0 | 0.012 | 0.0 |
| 5700 | 5800 | 0 | 0.0 | 0.0 | 0.0 | 0.004 | 0.0 |
| 5800 | 5900 | 0 | 0.0 | 0.0 | 0.0 | 0.002 | 0.0 |
| 5900 | 6000 | 0 | 0.0 | 0.0 | 0.0 | 0.002 | 0.0 |
| 7000 | 7100 | 0 | 0.0 | 0.0 | 0.0 | 0.002 | 0.0 |
| 8900 | 9000 | 0 | 0.0 | 0.0 | 0.0 | 0.002 | 0.0 |

where $\left(C_{0}\right)_{1}$ is the variance for each of the elevation ranges and $P_{1}$ is the percentage of the earth's surface area having that elevation range as estimated from the $1^{\circ}$ mean elevation data. Values of $C_{0}$ as estimated from (25) using all the data, and data from just the positive and negative elevations are given in Table Five.

| Estantare |  |
| :---: | :---: |
| Method | $\mathrm{C}_{\mathrm{O}}\left(\underline{m g a n}{ }^{2}\right)$ |
| Kaula (1959) | 1201 |
| Table Three | 1447 |
| Equation (25), all data | 1795 |
| Equation (25), negative elevations | 1772 |
| Equation (25), positive elevations | 1860 |
| Based on all anomalies without | 1644 |

For our future needs we select the $C_{0}=1795 \mathrm{mgal}^{2}$ as the best estimate. A truer value may even be larger than this as certain high variance values found in certain elevation ranges are not represented in the 1795 figures as our elevation data was not sufficiently detailed to tell us what percentage of the earth's surface lies within these elevation ranges. The 1795 value should be more reliable than the value of 1447 estimated from Table Three, as a certain smoothing has taken place in deriving
 the shape of the covariance curve in deriving the values for Table Three.

## 6. Anomaly Degree Variance Modeling

At this point we will develop a model for the anomaly degree variance which in turn will prove of value in deriving a closed expression for the covariance function of the disturbing potential and other gravimetric quantities. The basic procedures for this modeling have been discussed by Rapp (1973a). However, we introduce for this paper the s term and the $\beta_{\ell}$ term.

We first postulate an anomaly degree variance model of the following form:

$$
\begin{equation*}
c_{\ell}=\frac{A(\ell-1)}{(\ell-2)(\ell+B)} \tag{25~A}
\end{equation*}
$$

This model had originally been suggested by Tscherning. Best estimates for the A and $B$ parameters are to be found subject to the following data:

## 1. Anomaly Degree Variances Determined From Potential Coefficients

The values of $c_{l}$ that are used here are for $\ell=3$ to 20 are those found from the least squares collocation solution for potential coeffic ients as described in Rapp (1973b). These values are given in Table Five.

| Table Five |  |  |  |
| :---: | :---: | :---: | :---: |
| Anomaly Degree Variances From Potential Coefficients (Rapp, 1973b) ( $\mathrm{mgal}^{2}$ ) |  |  |  |
|  |  |  |  |
| $\ell$ | $\mathrm{c}_{\ell}$ | $\ell$ | $\mathrm{c}_{\ell}$ |
| 3 | 33.9 | 12 | 4.8 |
| 4 | 19.2 | 13 | 11.7 |
| 5 | 21.6 | 14 | 5.5 |
| 6 | 18.9 | 15 | 7.3 |
| 7 | 18.8 | 16 | 6.5 |
| 8 | 10.4 | 17 | 5.7 |
| 9 | 11.1 | 18 | 10.7 |
| 10 | 11.4 | 19 | 11.0 |
| 11 | 8.4 | 20 | 8.9 |

No formal standard deviations were attached to these values of $c_{\ell}$.
These values of $c_{\ell}$ can be directly used with (25A).

## 2. Anomaly Block and Point Variances

We have previously determined the block variances for $1^{\circ}$ and $5^{\circ}$ equal area blocks. These values can be related to $c_{\ell}$ values through equation (11) which is rewritten for the variance (i.e. $\psi=0$ ) as:

$$
\begin{equation*}
\overline{\mathrm{C}}(\psi=0)=\sum_{\ell=0}^{\infty} B_{\ell}^{z} \mathrm{c}_{\ell} \mathrm{s}^{\ell+z} \tag{26}
\end{equation*}
$$

Equation (26) is also valid for point anomalies recalling that in this case $B_{\ell}$ equals one.
In (26) the summation is started from $\ell=0$ but in fact we are trying to model $c_{\ell}$ from degree 3. Thus, we carry out the summation to degree 3 but we must modify our point and block variances by essentially removing the $c_{2}$ value. From Rapp (1973b) $c_{2}=7.5 \mathrm{mgal}^{2}$. The modified data is shown in Table Six.

Table Six

| Modified* Point and Block Variances <br> For Anomaly Degree Variance Fitting |  |
| :---: | :---: |
| Size | Modified |
|  | Variance |
| Point | 1788 mgal $^{2}$ |
| $1^{\circ}$ | 912 |
| $5^{\circ}$ | 295 |

*to refer to a complete second degree field
The adjustment procedure was carried out by first trying to determine best estimates of $A$ and $B$ for equation (25) by using the data of Table Five and the block variances of Table Six. The value of $\beta_{\ell}$ needed in (26) was computed using a $\psi$ value determined from equation (13). Tests indicated the summation to $\infty$ in (26) could sa fely be replaced by a summation to (4) $\left(180^{\circ}\right) / \theta^{\circ}$ or to $720 / \theta^{\circ}$. Various runs were made with different $s$ values to determine a proper value such that the summation to $\infty$ (or in practice a high number such as 50,000 or 100,000 ) would come close to the modified point variance of $1788 \mathrm{mgal}^{2}$. (It was found that for an accuracy of 0.1 mgals it was sufficient to carry out the point anomaly summation to $\ell=16000$ while for a 0.001 mgal accuracy the summation should be carried to about $\ell=30000$ ).

For theoretical reasons to be seen later, the $B$ unknown in equation (25A) should be an integer. To produce such an unknown we first made an adjustment letting A and B adjust freely. The resultant B found was 24.03 . We then repeated the adjustment, fixing $B$ at 24 exactly. In this adjustment the two block variances were given weights of $1 / 100$. All anomaly degree variances except for degree 3 and 4 were given weights of $1 / .64$. At degree 3 a weight of $1 / .08$ was used while at degree 4 a weight of $1 / .16$ was used. These weight assignments were made only to assure a reasonable fit to the data and were not based on relative accuracy considerations of the data.

We give in Table Seven the parameters of the final model.

> | Table Seven |
| :---: |
| Parameters of Anomaly Degree Variance Model |
| $\begin{array}{l}\mathrm{A}\end{array}=425.28 \mathrm{mgal}^{2}$ |
| $\mathrm{~B}=24$ (exact) |
| $\mathrm{S}=0.999617$ |

We give in Table Eight a comparison of the anomaly degree variances from Table Five and those as computed from Equation (25A)using the A and B values given in Table Seven.

Table Eight

| Anomaly Degree Variances (mgal $\left.{ }^{2}\right)$ |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: |
|  | Original <br> Table 5 | Equation <br> $(25 A)$ |  | Original | Equation <br> $(25 A)$ |
|  |  |  |  |  |  |
| 3 | 33.9 | 31.5 | 12 | 4.8 | 13.0 |
| 4 | 19.2 | 22.8 | 13 | 11.7 | 12.5 |
| 5 | 21.6 | 19.6 | 14 | 5.5 | 12.1 |
| 6 | 18.9 | 17.7 | 15 | 7.3 | 11.7 |
| 7 | 18.8 | 16.5 | 16 | 6.5 | 11.4 |
| 8 | 10.4 | 15.5 | 17 | 5.7 | 11.1 |
| 9 | 11.1 | 14.7 | 18 | 10.7 | 10.8 |
| 10 | 11.4 | 14.1 | 19 | 11.0 | 10.5 |
| 11 | 8.4 | 13.5 | 20 | 8.9 | 10.2 |

The root mean square difference between the original and adjusted values was $\pm 4.0 \mathrm{mgal}{ }^{2}$. The $1^{\circ}$ residual block variance from the adjusted model is $841 \mathrm{mgal}^{2}$ with the $5^{\circ}$ residual jlock variance being $360 \mathrm{mgal}^{2}$ as compared to the corresponding values of $912 \mathrm{mgal}^{2}$ and $295 \mathrm{mgal}^{2}$ as given in Table Six. By summing (26) with $\beta_{8}=1$ to a sufficiently high degree (50000) the point variance implied by this model is $1788 \mathrm{mgal}^{2}$. If we wished, at this point, the covariance functions implied by this new anomaly degree variance model could be computed by substitution of the model into equation (10) or (11). This :ype of computation will be postponed until the discussion of the closed covariance function expressions.
7. Relationship Between the Covariance Function of the Anomalous Potential and Covariance Functions of Gravity Anomalies or Deflections of the Vertical

As explained e.g. in Moritz (1972, p. 97), covariance functions of quantities related to the anomalous potential can be derived from the covariance function of the anomalous potential $K(P, Q)$. The covariance between two quantities $A$ and $B$, derived by applying a certain operation on $T$ can be found by applying the same operation on $K(P, Q)$. Moritz calls this fact "the law of propagation of covariances". We have above used the law to derive (15), and thereby the relation between $K(P, Q)$ and $C(P, Q)$. In the following we will derive the relationship between $K(P, Q)$ and the covariances of or between the height anomaly $\zeta$, the free-air gravity anomaly $\Delta g$ and the two deflection components $\xi$ and $\eta$.

We will use the same notation for the covariance functions as used in Moritz (1972), i.e. $\operatorname{cov}(A, B)$ for the covariance of the two quantities A and B. The relationship between the gravity anomaly and the anomalous potential is given above in (14). For the three other quantities we have the well known relations:

$$
\begin{align*}
& \zeta=\mathrm{T} / \gamma,  \tag{27}\\
& \xi=-\frac{1}{\gamma \cdot \mathrm{r}} \cdot \frac{\partial \mathrm{~T}}{\partial \varphi} \text { and }  \tag{28}\\
& \eta=-\frac{1}{\cos \varphi \cdot \gamma \cdot \mathrm{r}} \cdot \frac{\partial \mathrm{~T}}{\partial \lambda}, \tag{29}
\end{align*}
$$

where $\gamma$ is the reference gravity, $r$ the distance from the center of the Earth, of the latitude and $\lambda$ the longitude. It will for most purposes be sufficient to work in spherical approximation. But we will not restrict ourselves to consider only points on the surface of the Earth.

On the surface of the Earth $r$ is substituted by a mean Earth radius ( $\mathrm{R}_{\mathrm{e}}$ ), $\gamma$ by a mean gravity value (G), and $\varphi$ by the geocentric latitude. For a point outside (or inside) the surface of the Earth, we will substitute for $r$ the radius of a sphere e.g. including the same volume as an ellipsoid confocal with the adopted reference ellipsoid and passing through the considered point. (Thus, we will still call this quantity r). The reference gravity can then be substituted by $\mathrm{kM} / \mathrm{r}^{3}$ and $\varphi$ again with the proper geocentric latitude. (In practice $\varphi$ is just treated as if it was equal to the geocentric latitude).

We will introduce a more compact notation for the partial derivative with respect to an independent variable e.g. r:

$$
D_{r}=\frac{\partial}{\partial r},
$$

and for the second order partial derivative with respect to $r$ and $t$ :

$$
D_{r t}^{\partial}=\frac{\partial^{\partial}}{\partial r \partial t}
$$

The formulae (27), (14), (28) and (29) becomes then:

$$
\begin{align*}
\zeta & =\mathrm{T} / \mathrm{G}  \tag{30}\\
\Delta \mathrm{~g} & =-\mathrm{D}_{\mathrm{r}} \mathrm{~T}-\frac{2}{\mathrm{r}} \mathrm{~T}  \tag{31}\\
\xi & =-\frac{1}{\mathrm{G} \cdot \mathrm{r}} \mathrm{D}_{\varphi} \mathrm{T} \text { and }  \tag{32}\\
\eta & =-\frac{1}{\mathrm{G} \cdot \mathrm{r} \cdot \cos \varphi} \mathrm{D}_{\lambda} \mathrm{T} . \tag{33}
\end{align*}
$$

Using the law of propagation of covariances given by Moritz (1972, p. 97) applied to equations (30) - (33) we find:

$$
\begin{align*}
& \operatorname{cov}\left(\mathrm{T}_{\mathrm{P}}, \mathrm{~T}_{\mathrm{Q}}\right)=\mathrm{K}(\mathrm{P}, \mathrm{Q})  \tag{34}\\
& \operatorname{cov}\left(\Delta \mathrm{g}_{\mathrm{P}}, \Delta \mathrm{~g}_{\mathrm{Q}}\right)=\mathrm{C}(\mathrm{P}, \mathrm{Q})=\mathrm{D}_{\mathrm{r}} \mathrm{D}_{\mathrm{r}}{ }^{\prime} \mathrm{K}(\mathrm{P}, \mathrm{Q})+\frac{2}{\mathrm{r}} \cdot \mathrm{D}_{\mathrm{r}}{ }^{\prime} \mathrm{K}(\mathrm{P}, \mathrm{Q})+ \\
& \frac{2}{\mathrm{r}^{\prime}} \mathrm{D}_{\mathrm{r}} \mathrm{~K}(\mathrm{P}, \mathrm{Q})+\frac{4}{\mathrm{rr}^{\prime}} \mathrm{K}(\mathrm{P}, \mathrm{Q}),  \tag{35}\\
& \operatorname{cov}\left(\Delta g_{P}, \zeta_{Q}\right)=\left(-D_{r} K(P, Q)-\frac{2}{r} K(P, Q)\right) \cdot \frac{1}{\mathrm{G}^{r}},  \tag{36}\\
& \operatorname{cov}\left(\zeta_{P}, \zeta_{Q}\right)=K(P, Q) /\left(G \cdot G^{\prime}\right),  \tag{37}\\
& \operatorname{cov}\left(\xi_{P}, \zeta_{Q}\right)=-D_{\varphi} K(P, Q) /\left(G \cdot G^{\prime} \cdot r\right),  \tag{38}\\
& \operatorname{cov}\left(\eta_{P}, \zeta_{Q}\right)=-D_{\lambda} K(P, Q) /\left(G^{\prime} G \cdot r \cdot \cos \varphi\right)  \tag{39}\\
& \operatorname{cov}\left(\xi_{p}, \xi_{Q}\right)=D_{\varphi} D_{\varphi}{ }^{\prime} K(P, Q) /\left(G^{\prime} \cdot G \cdot r \cdot r^{\prime}\right)=D_{\varphi \varphi^{\prime}}^{2} K(P, Q) /\left(G \cdot G^{\prime} \cdot r^{\prime}\right),  \tag{40}\\
& \operatorname{cov}\left(\xi_{p}, \eta_{Q}\right)=D_{\varphi}^{2} \lambda^{\mathrm{K}}(\mathrm{P}, \mathrm{Q}) /\left(\mathrm{G}^{\prime} \cdot \mathrm{r}^{\prime} \cdot \cos _{\varphi}{ }^{\prime} \cdot \mathrm{r} \cdot \mathrm{G}\right),  \tag{41}\\
& \operatorname{cov}\left(\eta_{\mathrm{P}}, \eta_{\mathrm{Q}}\right)=\mathrm{D}_{\lambda}^{2} \lambda^{\mathrm{K}} \mathrm{~K}(\mathrm{P}, \mathrm{Q}) /\left(\mathrm{G}^{\prime} \cdot \mathrm{G} \cdot \mathrm{rr}^{\prime} \cos \varphi \cdot \cos \varphi^{\prime}\right),  \tag{42}\\
& \operatorname{cov}\left(\Delta g_{P}, \xi_{Q}\right)=-\mathrm{D}_{\varphi}\left(\operatorname{cov}\left(\Delta \mathrm{g}_{\mathrm{P}}, \zeta_{Q}\right)\right) / \mathrm{r}^{\prime}=\mathrm{D}_{\varphi^{\prime}}\left(\mathrm{D}_{\mathrm{r}} \mathrm{~K}(\mathrm{P}, \mathrm{Q})+\frac{2}{\mathrm{r}} \mathrm{~K}(\mathrm{P}, \mathrm{Q})\right) /\left(\mathrm{G}^{\prime} \mathrm{r}^{\prime}\right)  \tag{43}\\
& \operatorname{cov}\left(\Delta \mathrm{g}_{\mathrm{P}}, \eta_{\mathrm{Q}}\right)=-\mathrm{D}_{\lambda^{\prime}}\left(\operatorname{cov}\left(\Delta \mathrm{g}_{\mathrm{g}}, 5_{\mathrm{Q}}\right)\right) /\left(\mathrm{r}^{\prime} \cdot \cos \rho^{\prime}\right)=\mathrm{D}_{\lambda^{\prime}}\left(\mathrm{D}_{\mathrm{r}} \mathrm{~K}(\mathrm{P}, \mathrm{Q})+\right.  \tag{44}\\
& \left.\frac{2}{\mathrm{r}} \mathrm{~K}(\mathrm{P}, \mathrm{Q})\right) /\left(\mathrm{G}^{\prime} \mathrm{r}^{\prime} \cos \mathrm{c}^{\prime}\right),
\end{align*}
$$

- Eere the quantities marked with an apostrophe refer to $Q$ and the unmarked quantities $こ こ=r$ to $P$.

The covariances involving the deflections components ((38)-(44)) are most easily expressed and computed) as derivatives with respect to the cosine of the spherical distance $\psi$ between $P$ and $Q$. (We will from now on only regard isotropic covariance functions $K(P, Q)$, i.e. so that (9) always is valid and hence $K(P, Q)$ only depends on $\psi, r$ and $r^{\prime}$ ).

Putting $t=\cos \psi, D_{t} K(P, Q)=K^{\prime}$ and $D_{t}^{2} K(P, Q)=K^{\prime \prime}$ we get:

$$
\begin{aligned}
& D_{\varphi} K=D_{\varphi} t \cdot K^{\prime} \\
& D_{\lambda} K=D_{\lambda} t \cdot K^{\prime}
\end{aligned}
$$

Hence

$$
\begin{align*}
& \mathrm{D}_{\varphi \subset \rho}^{2}, \mathrm{~K}=\mathrm{D}_{\varphi \rho} \mathrm{t} \cdot \mathrm{D}_{\varphi}^{\prime \prime} \mathrm{t} \cdot \mathrm{~K}^{\prime \prime}+\mathrm{D}_{\varphi \varphi^{2}}^{2} \mathrm{t} \cdot \mathrm{~K}^{\prime}  \tag{45}\\
& \left.D_{\varphi \lambda^{\prime}}^{a} K=D_{\varphi} t \cdot D_{\lambda^{\prime}} t \cdot K^{\prime \prime}+D_{\varphi \lambda^{\prime}}^{a} t \cdot K^{\prime}\right)  \tag{46}\\
& D_{\lambda \lambda^{\prime}}^{2} K=D_{\lambda} t \cdot D_{\lambda^{\prime}} t \cdot K^{\prime \prime}+D_{\lambda \lambda^{\prime}}^{2} t \cdot K^{\prime} .  \tag{47}\\
& -\mathrm{D}_{\varphi^{\prime}}\left(\operatorname{cov}\left(\Delta \mathrm{g}_{P}, \mathrm{~T}_{\mathrm{Q}}\right)\right)=\mathrm{D}_{\varphi^{\prime}} \mathrm{t} \cdot\left(\mathrm{D}_{\mathrm{r}}^{2} \mathrm{~K}(\mathrm{P}, \mathrm{Q})+\frac{2}{\mathrm{r}} \mathrm{D}_{\mathrm{t}} \mathrm{~K}(\mathrm{P}, \mathrm{Q})\right)  \tag{48}\\
& -D_{\lambda^{\prime}}\left(\operatorname{cov}\left(\Delta g_{f}, T_{Q}\right)\right)=D_{\lambda^{\prime}} t \cdot\left(D_{\mathrm{rt}^{2}}^{2} K(P, Q)+\frac{2}{r} D_{t} K(P, Q)\right) \tag{49}
\end{align*}
$$

Note, the common factors $K^{\prime}$ and $K^{\prime \prime}$ in (47), (48) and (49), i. e., the three covariance functions $\operatorname{cov}\left(\xi_{P}, \xi_{Q}\right), \operatorname{cov}\left(\xi_{P}, \eta_{Q}\right)$ and $\operatorname{cov}\left(\eta_{P}, \eta_{Q}\right)$ can easily be computed at the same time. The covariance functions (38) - (44) are used in actual prediction computations involving deflections either as observed quantities or as quantities to be predicted. These covariance functions are not anymore isotropic. Then for theoretical discussions it is more convenient to regard the covariances, where one or both of the quantities are either the longitudinal ( $\ell$ ) or the transverse component ( m ) of the deflection of the vertical. This type of covariance function will be isotropic and will have a simple relation to $\mathrm{K}(\mathrm{P}, \mathrm{Q})$.


Figure 2.
Epherical triangle (Pole, Q, P) with the deflection components ( $\xi, \eta$ ) and ( $\ell, \mathrm{m}$ ) shown as vectors.

In Moritz (1972) the relationships between $K(P, Q)$ and the covariance functions are expressed in terms of derivatives with respect to $\psi$. We will express the relations in terms of derivatives with respect to $t=\cos \psi$.

Let the azimuth between $P$ and $Q$ be $\alpha$. Then we have (cf. figure 2):

$$
\begin{align*}
& \ell_{p}=\cos \alpha \cdot\left(-\xi_{p}\right)+\sin \alpha_{0}\left(-\eta_{p}\right) \text { and }  \tag{50}\\
& m_{p}=\sin \alpha \cdot\left(-\xi_{p}\right)-\cos \alpha \cdot\left(-\eta_{p}\right)
\end{align*}
$$

Using (38) and (39) and the law of propagation of covariances, we get:

$$
\begin{aligned}
& \operatorname{cov}\left(\ell_{p}, \zeta_{Q}\right)=\left(\cos \alpha \cdot D_{\varphi} t \cdot K^{\prime}+\sin \alpha \cdot D_{\lambda} t \cdot K^{\prime} \cdot \frac{1}{\cos \varphi}\right) /\left(G \cdot G^{\prime} \cdot r\right) \text { and } \\
& \operatorname{cov}\left(m_{p}, \zeta_{Q}\right)=\left(\sin \alpha \cdot D_{\varphi} t \cdot K^{\prime}-\cos \alpha \cdot D_{\lambda} t \cdot K^{\prime} \cdot \frac{1}{\cos 0}\right) /\left(G \cdot G^{\prime} \cdot r\right)
\end{aligned}
$$

Because

$$
\mathrm{t}=\sin \varphi^{\prime} \cdot \sin \varphi^{\prime}+\cos \infty \cdot \cos \theta^{\prime} \cdot \cos \left(\lambda^{\prime}-\dot{\lambda}\right)
$$

we have

$$
D_{\varphi} t=\cos \varphi \cdot \sin \varphi^{\prime}-\sin \varphi \cdot \cos \varphi^{\prime} \cdot \cos \left(\lambda^{\prime}-\lambda\right)=\sin \psi \cdot \cos \alpha \text { and }
$$

$$
D_{\lambda} t=\cos \varphi \cdot \cos \varphi^{\prime} \sin \left(\lambda^{\prime}-\lambda\right)=\cos \varphi \cdot \sin \psi \cdot \sin \alpha
$$

hence

$$
\begin{equation*}
\operatorname{cov}\left(l_{p}, \zeta_{Q}\right)=\sin \psi \cdot \mathrm{K}^{\prime} /\left(\mathrm{G} \cdot \mathrm{G}^{\prime} \cdot \mathrm{r}\right) \text { and } \tag{51}
\end{equation*}
$$

$$
\begin{equation*}
\operatorname{cov}\left(m_{P}, \zeta_{Q}\right)=0 \tag{52}
\end{equation*}
$$

For the covariance with the gravity anomaly we get in the same way (using the law of propagation of covariations and (14))

$$
\begin{align*}
& \operatorname{cov}\left(\ell_{p}, \Delta \mathrm{~g}_{Q}\right)=+\left(\mathrm{D}_{\mathrm{r}} \quad \mathrm{~K}^{\prime}+\frac{2}{\mathrm{r}} \mathrm{~K}^{\prime}\right) \cdot \sin \psi /(\mathrm{G} \cdot \mathrm{r})  \tag{53}\\
& \operatorname{cov}\left(\mathrm{m}_{\rho}, \Delta \mathrm{g}_{Q}\right)=0 \tag{54}
\end{align*}
$$

The expressions for $\operatorname{cov}\left(\ell_{p}, \ell_{Q}\right), \operatorname{cov}\left(\ell_{p}, m_{Q}\right)$ and $\operatorname{cov}\left(m_{p}, m_{Q}\right)$ are derived in a very simple way in Moritz (1972, p. 109). We repeat the results expressed as derivatives of $t$.

$$
\begin{align*}
& \operatorname{cov}\left(l_{p}, l_{Q}\right)=-D_{\psi}^{2} K /\left(G \cdot G^{\prime} \cdot r \cdot r^{\prime}\right)=\left(t \cdot K^{\prime}-\sin ^{2} \psi \cdot K^{\prime \prime}\right) /\left(G \cdot G^{\prime} \cdot r \cdot r^{\prime}\right)  \tag{55}\\
& \operatorname{cov}\left(l_{p}, m_{Q}\right)=0 \text { and }  \tag{56}\\
& \operatorname{cov}\left(m_{p}, m_{Q}\right)=-D_{\psi} K /\left(\sin \psi \cdot G \cdot G^{\prime} \cdot r \cdot r^{\prime}\right)=K^{\prime} /\left(G \cdot G^{\prime} \cdot r \cdot r^{\prime}\right) \tag{57}
\end{align*}
$$

From the formulae (51) - (57) several intersting consequences of the imposed isotropic property can be seen. The deflection components at P are independent of the height anomaly and the gravity anomaly in $P$. The transverse component of the deflection in $P, m_{p}$ is independent of $\zeta_{Q}, \Delta g_{Q}$ and $\ell_{Q}$. For $\xi_{p}$ this implies, that $\xi_{p}$ is independent of $\eta_{Q}$ for $\varphi=\varphi^{\prime}$ and $\eta_{P}$ independent of $\xi_{Q}$ for $\lambda=\lambda^{\prime}$.

Finally we will conclude that the basic quantities to be computed in the evaluation of the expressions (34)-(49) and (51) - (57) are $\mathrm{K}, \mathrm{K}^{\prime}, \mathrm{K}^{\prime \prime}, \mathrm{D}_{\mathrm{r}} \mathrm{K}^{\prime}+\frac{2}{\mathrm{r}} \mathrm{K}$ and $\operatorname{cov}\left(\Delta \mathrm{g}_{\mathrm{P}}, \Delta \mathrm{g}_{\mathrm{q}}\right)$.

## 8. Closed covariance function expressions.

In this section we will consider different models for the degree-variances and explain how closed expressions for corresponding covariance functions can be obtained. We will distinguish between different types of degree-variances and hence between different covariance functions models. Thus we will still consider only isotropic models. A subscript k will be used to distinguish between the models. Then we can define $\sigma_{k}, \ell(A, B)$ to be the degree-variances of degree $\ell$ in the $k^{\prime}$ th degree-variance model, i. e. so that the corresponding covariance function becomes:

$$
\begin{equation*}
\left.\operatorname{cov}_{k}(\mathrm{~A}, \mathrm{~B})=\left(\frac{\mathrm{R}}{\mathrm{r}}\right)^{\prime}\left(\frac{\mathrm{R}}{\mathrm{r}}\right)^{\nu} \sum_{\ell=0}^{\infty} \sigma_{\mathrm{k}, \ell(\mathrm{~A}}, \mathrm{B}\right) \mathrm{s}^{\ell+1} \mathrm{P}_{\ell}(\mathrm{t}), \tag{58}
\end{equation*}
$$

Where $I$ and $J$ are either 0 or 1 . (Note, that for $I=J=1$ we have $\left(\frac{R}{r}\right) \cdot\left(\frac{R}{r^{\prime}}\right)=s$ ). For the already introduced quantities $c_{\ell}$ and $\sigma_{l}$ we then have:

$$
\begin{aligned}
& \mathrm{c}_{\ell}=\sigma_{\mathrm{k}, \ell}(\Delta \mathrm{~g}, \Delta \mathrm{~g}) \text { and } \\
& \sigma_{\ell}=\sigma_{\mathrm{k}, \ell}(\mathrm{~T}, \mathrm{~T}) .
\end{aligned}
$$

The corresponding covariance functions become, using (9) and (10)

$$
\begin{align*}
\operatorname{cov}_{\mathrm{k}}\left(\mathrm{~T}_{\mathrm{P}}, \mathrm{~T}_{\mathrm{Q}}\right) & =\sum_{\ell=0}^{\infty} \sigma_{\mathrm{k}, \ell(\mathrm{~T}, \mathrm{~T}) \mathrm{s}^{\ell+1} \mathrm{P}_{\ell}(\mathrm{t}) \text { and }}  \tag{59}\\
\operatorname{cov}_{\mathrm{k}}\left(\Delta \mathrm{~g}_{\mathrm{P}}, \Delta \mathrm{~g}_{\mathrm{q}}\right) & =\sum_{\ell=0}^{\infty} \sigma_{\mathrm{k}, \ell}(\Delta \mathrm{~g}, \Delta \mathrm{~g}) \mathrm{s}^{\ell+2} \mathrm{P}_{\ell}(\mathrm{t})  \tag{60}\\
& =\left(\frac{\mathrm{R}}{\mathrm{r}}\right)\left(\frac{\mathrm{R}}{\mathrm{r}}\right) \sum_{\ell=0}^{\infty} \sigma_{\mathrm{k}, \ell}(\Delta \mathrm{~g}, \Delta \mathrm{~g}) \mathrm{s}^{\ell+1} \mathrm{P}_{\ell}(\mathrm{t})
\end{align*}
$$

The relationship (15) becomes:

$$
\begin{equation*}
\sigma_{\mathrm{k}, \ell}(\mathrm{~T}, \mathrm{~T})=\frac{\mathrm{R}^{2}}{(\ell-1)^{2}} \sigma_{\mathrm{k}, \ell}(\Delta \mathrm{~g}, \Delta \mathrm{~g}) \tag{61}
\end{equation*}
$$

In the following we will also consider the degree-variances $\sigma_{\mathrm{k}, \ell}(\Delta \mathrm{g}, \mathrm{T})$ of the covariance function $\operatorname{cov}_{k}\left(\Delta \mathrm{~g}_{\mathrm{P}}, \mathrm{T}_{\mathrm{Q}}\right)$ which is related to the covariance (36) by:

$$
\operatorname{cov}_{k}\left(\Delta g_{\rho}, T_{Q}\right)=\operatorname{cov}\left(\Delta g_{\boldsymbol{p}}, \zeta_{Q}\right) \cdot G^{\prime}
$$

Using (36) and (59) we get:

$$
\begin{align*}
\operatorname{cov}_{k}\left(\Delta g_{P}, T_{Q}\right) & =-D_{r}\left(\sum_{l=0}^{\infty} \sigma_{k}, \ell(T, T) s^{\ell+1} P_{l}(t)\right)-\frac{2}{r}\left(\sum_{l=0}^{\infty} \sigma_{k, \ell}(T, T) s^{\ell+1} P_{l}(t)\right) \\
& =\sum_{l=0}^{\infty} \sigma_{k}, \ell(T, T) \frac{(l-1)}{r} s^{\ell+1} P_{\ell}(t)  \tag{62}\\
& =\frac{R}{r} \sum_{l=0}^{\infty} \sigma_{\mathrm{k}, \ell}(T, T) \frac{(l-1)}{R} s^{\ell+1} P_{l}(\mathrm{t}) .
\end{align*}
$$

Hence, using (58) we see that $\mathrm{I}=1$ and $\mathrm{J}=0$ and that

$$
\begin{equation*}
\sigma_{\mathrm{k}, \ell}(\Delta \mathrm{~g}, \mathrm{~T})=\sigma_{\mathrm{k}, \ell}(\mathrm{~T}, \mathrm{~T}) \cdot \frac{(\ell-1)}{\mathrm{R}} \quad \text { and } \tag{63}
\end{equation*}
$$

$$
\begin{equation*}
\operatorname{cov}_{\mathrm{k}}\left(\Delta \mathrm{~g}_{\mathrm{p}}, \mathrm{~T}_{\mathrm{Q}}\right)=\frac{\mathrm{R}}{\mathrm{r}} \sum_{\ell=0}^{\infty} \mathrm{\sigma}_{\mathrm{k}, \ell}(\Delta \mathrm{~g}, \mathrm{~T}) \mathrm{s}^{\ell+1} \mathrm{P}_{\ell}(\mathrm{t}) . \tag{64}
\end{equation*}
$$

(Note, that the introduced notation can't be used for covariance-functions involving deflections. These covariance functions can be expressed as the sums of series in $P_{\ell}^{\prime}(t)$ and $P_{\ell}^{\prime \prime}(t)$ (apostrophe mean differentiation with respect to $t$ ), and not on the form (58) as a series in $P_{\ell}(t)$ and $s^{\ell+1}$.)

Five different models of the anomaly degree variances will be discussed below, i.e., k will take on values $1,2, \ldots 5$.

In Tscherning (1972), analytic models have been described for covariance function having anomaly degree-variances equal to:

$$
\begin{align*}
& \sigma_{1}, \ell(\Delta \mathrm{~g}, \Delta \mathrm{~g})=\mathrm{A}_{1}(\ell-1)^{2}, \quad \ell>1  \tag{65}\\
& \sigma_{2, \ell}(\Delta \mathrm{~g}, \Delta \mathrm{~g})=\mathrm{A}_{2}(\ell-1) / \ell, \quad \ell>1 \text { and }  \tag{66}\\
& \sigma_{3, \ell}(\Delta \mathrm{~g}, \Delta \mathrm{~g})=\mathrm{A}_{3}(\ell-1) /(\ell-2), \quad \ell>2 \tag{67}
\end{align*}
$$

where $A_{1}, A_{2}$, and $A_{3}$ (and below $A_{4}$ and $A_{5}$ ) are positive constants of dimension mgal ${ }^{2}$. These types of models have been further cons idered by Rapp (1972a).

$$
\begin{align*}
& \sigma_{4} \ell(\Delta \mathrm{~g}, \Delta \mathrm{~g})=\mathrm{A}_{4} \frac{(\ell-1)}{(\ell-2)(\ell+B)} \quad \text { and }  \tag{68}\\
& \sigma_{5} \ell(\Delta \mathrm{~g}, \Delta \mathrm{~g})=\mathrm{A}_{5} \frac{(\ell-1)}{(\ell-2)\left(\ell+B+\beta \ell^{2}\right)}, \ell>2 \tag{69}
\end{align*}
$$

For $\mathrm{i}+\mathrm{j}=\frac{1}{\beta}$ and $\mathrm{i} \cdot \mathrm{j}=\mathrm{B} / \beta$ we can write (69):

$$
\begin{equation*}
\sigma_{5, \ell}(\Delta \mathrm{~g}, \Delta \mathrm{~g})=\frac{\mathrm{A}_{5}}{\beta} \cdot \frac{(\ell-1)}{(\ell-2)(\ell+\mathrm{i})(\ell+\mathrm{j})} \tag{70}
\end{equation*}
$$

As indicated above, the covariance functions corresponding to models 1,2 and 3 can be represented by closed expressions. (By closed expression we mean expressions which only contain a finite number of terms). This is also true for model 4 and 5 , provided we place some restrictions on $B$ or $i$ and $j$. First of all the resulting degreevariances have to be greater than or equal to zero for $\ell$ greater than 2. Hence, B and i , j will have to be greater than -2 . And the technique used below for the derivation will imply that we have to restrict $B$ and $i, j$ to integer values and that we also will have to require that i is unequal to j and that all three quantities are greater than -1 .

We will not consider the covariance functions derived using the model (65) because the anomaly degree-variances are unrealistic. Thus, the model leads to very simple closed expressions for the covariance functions, which can be found, e.g. in Tscherning (1972).

The technique we will use for the derivation of the closed covariance expressions is very simple. The covariance functions can be split into components which, upon multipl ication by appropriate constants will yield the covariance function. These components can be expressed as:

$$
\begin{align*}
& F=\sum_{l=0}^{\infty} s^{\ell+1} P_{\ell}(t) \text { and }  \tag{71}\\
& F_{1}=\sum_{l=0}^{\infty} \frac{1}{\ell+i} s^{\ell+1} P_{\ell}(t) \text { for } i>0  \tag{72}\\
& F_{i}=\sum_{l=-1+1}^{\infty} \frac{1}{\ell+i} s^{\ell+1} P_{\ell}(t) \text { for } i \leq 0, \text { and } \tag{73}
\end{align*}
$$

as the first and second derivatives of $F$ or $F_{1}$ with respect to $t$,

$$
F^{\prime}, F^{\prime \prime}, F_{i}^{\prime}, F_{1}^{\prime \prime}
$$

We have for example (using (59) (61) and (67)):

$$
\begin{aligned}
\operatorname{cov}_{3}\left(T_{P}, T_{Q}\right) & =A_{3} \sum_{\ell=3}^{\infty} \frac{R^{2}}{(\ell-1)(\ell-2)} s^{\ell+1} P_{\ell}(t) \\
& =A_{3} \cdot R^{2} \sum_{\ell=3}^{\infty}\left(\frac{1}{\ell-2}-\frac{1}{\ell-1}\right) s^{\ell+1} P_{\ell}(t) \\
& =A_{3} \cdot R^{2}\left(F_{-2}-\left(F_{-1}-s^{3} P_{2}(t)\right)\right)
\end{aligned}
$$

The closed expression for the function $F$ can be derived using the well known formula (Heiskanen and Moritz, 1967, eq. 1-80):

$$
\begin{equation*}
F=s \cdot \sum_{\ell=0}^{\infty} s^{\ell} P_{\ell}(t)=\frac{s}{\sqrt{1-2 s t+s^{a}}} \tag{74}
\end{equation*}
$$

The denominator will be one of the bas ic quantities in the following derivations, so we will use:

$$
\begin{align*}
& \mathrm{L}=\sqrt{1-2 \mathrm{st}+\mathrm{s}^{\rho}}, \\
& \mathrm{M}=1-\mathrm{L}-\mathrm{s} \cdot \mathrm{t} \text { and }  \tag{75}\\
& \mathrm{N}=1+\mathrm{L}-\mathrm{st} .
\end{align*}
$$

We then have:

$$
F=\frac{s}{I}
$$

The functions $F_{1}$ can be derived by multiplying $F$ or $\frac{1}{L}$ by an appropriate power of $s$ and integrating the expression with respect to $s$.

Using:

$$
\int_{0} \mathrm{~s}^{\ell+\mathrm{i}-1} \mathrm{ds}=\frac{\mathrm{s}^{\ell+\mathrm{i}}}{\ell+\mathrm{i}}, \ell+\mathrm{i}>0
$$

we see, that by integrating

$$
\begin{equation*}
\frac{s^{i-1}}{L}=\sum_{l=0}^{\infty} s^{\ell+i-1} P_{\ell}(t) \tag{76}
\end{equation*}
$$

we should be able to find $F_{1}$. We have by (72)

$$
\begin{equation*}
s^{i-1} F_{1}=\sum_{l=0}^{\infty} \frac{s^{\ell+i}}{l+i} P_{l}(t) \text { for } i>0 \text { and by (73): } \tag{77}
\end{equation*}
$$

$$
\begin{equation*}
s^{i-1} \mathrm{~F}_{1}=\sum_{l=0, \ell \neq-1}^{\infty} \frac{s^{\ell+j}}{\ell+i} \mathrm{P}_{\ell}(\mathrm{t})-\sum_{\ell=0}^{1-1} \frac{\mathrm{~s}^{\ell+1}}{\ell+\mathrm{i}} \mathrm{P}_{\ell}(\mathrm{t}), \mathrm{i} \leq 0 \tag{78}
\end{equation*}
$$

-he integrals:

$$
\begin{equation*}
\int \frac{\mathrm{s}^{1}}{\mathrm{~L}} \mathrm{ds}, \mathrm{i}=-2,-1,0,1,2 \tag{79}
\end{equation*}
$$

$\therefore$ an be found in integral tables as Gradshteyn-Ryzhik (abbreviated below to G. R.), (1965).
From these basic integrals, $\mathrm{F}_{1}$ can be computed using recursion formulae. We will Eirst consider negative powers of $s$.

Using G.R. 2. 268 we get:

$$
\begin{equation*}
\int \frac{d s}{s^{1} L}=-\frac{L}{(i-1) s^{1-1}}+\frac{(2 i-3) \cdot t}{(i-1)} \int \frac{d s}{s^{1-1} L}-\frac{i-2}{(i-1)} \int \frac{d s}{s^{i-2} L}, i>0 \tag{80}
\end{equation*}
$$

and hence

$$
\begin{align*}
\int \frac{d s}{s^{2} L} & =-\frac{L}{s}+t \cdot \int \frac{d s}{s \cdot L}+a_{-2}  \tag{81}\\
\int \frac{d s}{s^{3} L} & =-\frac{L}{2 s^{2}}+\frac{3}{2} t\left(-\frac{L}{s}+t \int \frac{d s}{s \cdot L}\right)-\frac{1}{2} \int \frac{d s}{s \cdot L}+a_{-3}  \tag{82}\\
& =-\frac{3 t s+1}{2 s^{2}} \cdot L+P_{2}(t) \int \frac{d s}{s \cdot L}+a_{-3}
\end{align*}
$$

where $a_{-2}, a_{-3}$ are integration constants. From G.R. 2.266 we have:

$$
\begin{align*}
\int \frac{d s}{s \cdot L} & =-\ln \frac{2-2 t s+2 \cdot L}{s}+a_{-1}=\ln \frac{2}{1-t s+L}+\ln (s)+\ln (4)+a_{-1}  \tag{83}\\
& =\ln \frac{2}{N}+\ln (s)+\ln (4)+a_{-1}
\end{align*}
$$

The constant $a_{-1}$ is determined requiring (78) to be zero for $s$ equal to zero.

$$
s^{-1} F_{0}=\int \frac{d s}{s \cdot L}-\int \frac{d s}{s}=\ln \frac{2}{N}+\ln (4)+a_{-1},
$$

hence $\mathrm{a}_{-1}=-2 n(4)$ and then:

$$
\begin{equation*}
\mathrm{s}^{-1} \mathrm{~F}_{\mathrm{O}}=\ln \frac{2}{\mathrm{~N}} \tag{84}
\end{equation*}
$$

Then we can compute $\mathrm{F}_{-1}$ and $\mathrm{F}_{-2}$.

$$
\begin{align*}
& s^{-2} F_{-1}=\sum_{\ell=2}^{\infty} \frac{s^{\ell-1}}{\ell-1} P_{\ell}(t)=-\frac{L}{s}+t \cdot \ell \frac{2}{N}-\int \frac{d s}{s^{2}}+a_{-a} \\
& =\frac{1-L}{s}+t \cdot \ln \frac{2}{N}+a_{-2}=\frac{1-t s-L}{s}+t \cdot \ln \frac{2}{N} \text { or }  \tag{85}\\
& F_{-1}=s \cdot\left(M+t s \cdot \ln \frac{2}{N}\right),  \tag{86}\\
& s^{-3} F_{-2}=\sum_{\ell=3}^{\infty} \frac{s^{\ell-2}}{\ell-2} P_{\ell}(t)=\int\left(\frac{1}{s^{3}}+\frac{t}{s^{2}}+\frac{P_{2}(t)}{s}\right) d s+\int \frac{d s}{s^{3} L}+a_{-3} \\
& =\frac{1}{2 s^{2}}+\frac{t}{s}-P_{2}(t) \cdot \ln (s)+\left(P_{1}(t) \cdot\left(\ln \frac{2}{N}+\ln (s)\right)-\frac{3 t s+1}{2 s^{2}} L+a-3\right) \\
& =(1+2 t s-(3 t s+1) \cdot L) /\left(2 s^{2}\right)+P_{2}(t) \cdot \ln \frac{2}{N}+a_{-3} .
\end{align*}
$$

The constant $a_{-3}$ can now be determined. Because $\ell \frac{2}{\mathrm{~N}}$ is zero for $s$ equal to zero, we must have:

$$
\lim _{s \rightarrow 0} \frac{1+2 \mathrm{ts}-(3 \mathrm{ts}+1) \cdot \mathrm{L}}{2 \mathrm{~s}^{2}}=-\mathrm{a}_{-3}
$$

The limit can be determined using the rule of l'Hospital two times. Note first, that $D_{s} L=(s-t) / L$. We then get

$$
\begin{aligned}
& \lim _{s \rightarrow 0} \frac{1+2 s t-(3 t s+1) \cdot L}{2 s^{2}}=\lim _{s \rightarrow 0} \frac{-3 t \cdot L-(3 t s+1)(s-t) / L+2 t}{4 s} \\
& =\lim _{s \rightarrow 0}\left(-3 t(s-t) / L-\left(6 t s+1-3 t^{2}\right) / L+\left(3 t s^{2}+s-3 t^{2} s-t\right)(s-t) / L^{3}\right) / 4
\end{aligned}
$$

and hence:

$$
a_{-2}=-\frac{7 t^{2}-1}{4}=-\frac{3}{2} t^{2}+\frac{1-t^{2}}{4}
$$

We then get:

$$
\begin{align*}
F_{-2} & =s\left(\left(-3 t^{2} s^{2}+2 t s+1-(3 t s+1) \cdot L\right) / 2+\left(P_{2}(t) \cdot \ln \frac{2}{N}+\frac{1-t^{2}}{4}\right) \cdot s^{2}\right) \\
& =s\left((1-t s-L)(3 t s+1) / 2+s^{2}\left(P_{2}(t) \cdot \ln \frac{2}{N}+\frac{1-t^{2}}{4}\right)\right)  \tag{87}\\
& =s\left(M \cdot(3 t s+1) / 2+s^{2}\left(P_{2}(t) \cdot \ln \frac{2}{N}+\left(1-t^{2}\right) / 4\right)\right) .
\end{align*}
$$

For the evaluation of covariance functions involving deflections, we have to compute $F_{0}^{\prime}, F_{0}^{\prime \prime}, F_{-1}^{\prime}, F_{-1}^{\prime \prime}, F_{-2}^{\prime}$, and $F_{-2}^{\prime \prime}$ (where again the apostrophe means differentiation with respect to $t=\cos \psi$ ).

We will first compute some auxiliary quantities:

$$
\begin{aligned}
& D_{t} L=-\frac{s}{L}, D_{t}\left(\frac{1}{L}\right)=\frac{s}{L^{3}}, \\
& D_{t} M=-s+\frac{S}{L}=s \frac{(1-L)}{L} \\
& D_{t} N=-s-\frac{S}{L}=-s(1+L) / L
\end{aligned}
$$

Hence from (84) we get by differentiation:

$$
\begin{align*}
F_{0}^{\prime} & =s \cdot D_{t} \ln \frac{2}{N}=\frac{s^{2}(1+L)}{L \cdot N}=s^{2}\left(\frac{1}{L \cdot N}+\frac{1}{N}\right)  \tag{88}\\
F_{0}^{\prime \prime} & =s^{2}\left(\frac{-N(-s / L)-L(-s(1+L) / L)}{(L \cdot N)^{2}}+\frac{s(1+L) / L}{N^{2}}\right)  \tag{89}\\
& =s^{3}\left(\frac{N+L(1+L)}{L^{3} \cdot N^{2}}+\frac{1+L}{L \cdot N^{2}}\right)=s^{3}\left(\frac{N+L}{L^{3} \cdot N^{2}}+\frac{2+L}{L \cdot N^{2}}\right)
\end{align*}
$$

For $\mathrm{F}_{-1}$ we get using (86)

$$
\begin{align*}
F_{-1}^{\prime} & =s\left(D_{t} M+F_{O}+t \cdot F_{0}^{\prime}\right) \\
& =s\left(s(1-L) / L+F_{0}+t \cdot s^{3}\left(\frac{1}{L \cdot N}+\frac{1}{N}\right)\right)  \tag{90}\\
& =s^{2}\left((1-L) / L+\ln \frac{2}{N}+t \cdot s\left(\frac{1}{L \cdot N}+\frac{1}{N}\right)\right) \\
F_{-1}^{\prime \prime} & =s\left(D_{t}^{3} M+2 F_{0}^{\prime}+t \cdot F_{0}^{\prime \prime}\right) \\
& =s\left(\frac{s^{3}}{L^{3}}+2 s^{2}\left(\frac{1}{L \cdot N}+\frac{1}{N}\right)+t \cdot s^{3}\left(\frac{N+L}{L^{3} \cdot N^{3}}+\frac{2+L}{L \cdot N^{2}}\right)\right)  \tag{91}\\
& =s^{3}\left(\frac{1}{L^{3}}+\frac{2(1+L)}{L \cdot N}+t s\left(\frac{1}{L^{3} \cdot N}+\frac{(1+L)^{2}}{(L \cdot N)^{2}}\right)\right)
\end{align*}
$$

and for $\mathrm{F}_{-2}$ we get by (87)

$$
\begin{align*}
F_{-2}^{\prime} & =s\left[D_{t} M \cdot(3 t s+1) / 2+M \cdot \frac{3 s}{2}+s^{2}\left(3 t \cdot \ln \frac{2}{N}+P_{2}(t) \cdot D_{t} \ln \frac{2}{N}-t / 2\right)\right] \\
& =s^{2}\left[(3 t s+1)(1-L) /(2 L)+\frac{3}{2} M+s\left(3 t \cdot \ln \frac{2}{N}+P_{2}(t) \cdot s\left(\frac{1}{L N}+\frac{1}{N}\right)-t / 2\right)\right]  \tag{92}\\
& =s^{2}\left[\frac{1}{2}((3 t s+1) / L+2-7 t s-3 L)+s\left(3 t \cdot \ln \frac{2}{N}+P_{2}(t) \cdot s\left(\frac{1+L}{L \cdot N}\right)\right)\right]
\end{align*}
$$

$$
\begin{align*}
F_{-a}^{\prime \prime}= & s^{2}\left[\frac{1}{2}\left(\frac{L \cdot 3 s-(3 t s+1)(-s / L)}{L^{2}}-7 s+\frac{3 s}{L}\right)+s\left(32 n \frac{2}{N}+6 t s \frac{1+L}{L \cdot N}\right.\right. \\
& \left.\left.+P_{2}(t) \cdot s\left(-\frac{-s(1 \cdot L)}{N^{2} \cdot L}-\frac{(-s / L) \cdot N-L(s(1+L) / L)}{(L \cdot N)^{2}}\right)\right)\right]  \tag{93}\\
= & \left.s^{3}\left[\left(\frac{6}{L}+\frac{3 t s+1}{L^{3}}-7\right) \frac{1}{2}+3 e n \frac{2}{N}+6 t s \frac{1+L}{L \cdot N}+P_{2}(t) \cdot s^{2}\left(\left(\frac{1+L}{L \cdot N}\right)^{2}+\frac{1}{L^{3} N}\right)\right)\right]
\end{align*}
$$

I:E closed expressions for $\mathrm{F}_{1}$, i> 0 can be found using another recursion formula,
$\approx \ldots .2 .263$. We will treat this case in a more general way, because in this case we mza: to derive expressions not only for $i=1,2$ and 3 but for $i=1$ to $\infty$. It is also $=\doteq=$ essary to have a recursion formula well suited for actual computations.

- \# ¿ have (using G. R. 2.263):

$$
\begin{align*}
& \int \frac{s^{1} d s}{L}=\frac{s^{i-1} \cdot L}{i}+\frac{(2 i-1) t}{i} \int \frac{s^{i-1}}{L} d s-\frac{(i-1)}{i} \int \frac{s^{i-2}}{L} d s \text { or }  \tag{94}\\
& \frac{1}{s^{i-1}} \int \frac{s^{1} d s}{L}=\left(L+(2 i-1) t \cdot \frac{1}{s^{i-1}} \int \frac{s^{i-1} d s}{L}-\frac{(i-1)}{s} \cdot \frac{1}{s^{i-2}} \int \frac{s^{i-2} d s}{L}\right) \cdot \frac{1}{i} \tag{95}
\end{align*}
$$

Fニalizing that:

$$
F_{1}=\frac{1}{s^{i}-1} \int_{0} \frac{s^{i-1}}{L} d s \quad \text { for } i>0
$$

a get

$$
\begin{equation*}
\mathrm{F}_{1+1}=\left(\mathrm{L}+(2 \mathrm{i}-1) \mathrm{t} \cdot \mathrm{~F}_{1}-\frac{(\mathrm{i}-1)}{\mathrm{s}} \cdot \mathrm{~F}_{1-1}\right) \cdot \frac{1}{\mathrm{~s} \cdot \mathrm{i}} . \tag{96}
\end{equation*}
$$

Eortunately we can use the recursion formula for the computation of $D_{t} F_{1}=F_{1}^{\prime}$ and
$\mathcal{F}_{1}^{2}=F_{1}^{\prime \prime}$ as well.
-ifferentiating (96) we have:

$$
\begin{equation*}
F_{1+1}^{\prime}=\left(D_{t} L+(2 i-1)\left(F_{1}+t \cdot F_{1}^{\prime}\right)-\frac{(i-1)}{s} \cdot F_{1-1}^{\prime}\right) \cdot \frac{1}{i \cdot s} \text { and } \tag{97}
\end{equation*}
$$

Sifferentiating one time more:

$$
\begin{equation*}
\mathrm{F}_{1+1}^{\prime \prime}=\left(\mathrm{D}_{\mathrm{t}}^{2} \mathrm{~L}+(2 \mathrm{i}-1)\left(2 \mathrm{~F}_{1}^{\prime}+\mathrm{t} \cdot \mathrm{~F}_{1}^{\prime \prime}\right)-\frac{(\mathrm{i}-1)}{\mathrm{s}} \cdot \mathrm{~F}_{1-1}^{\prime \prime}\right) \cdot \frac{1}{\mathrm{i} \cdot \mathrm{~s}} \tag{98}
\end{equation*}
$$

with $D_{t} L=-\frac{S}{L}$ and $D_{t}^{2} L=-\frac{S^{2}}{L^{3}}$.

As in the case where i was less than or equal to zero, we must now compute the first two terms in the recursion formula, i.e.

$$
\mathrm{F}_{1}, \mathrm{~F}_{1}^{\prime}, \mathrm{F}_{1}^{\prime \prime}, \mathrm{F}_{2}, \mathrm{~F}_{2}^{\prime} \text { and } \mathrm{F}_{2}^{\prime \prime}
$$

Us ing G. R. 2. 2641 we get

$$
F_{1}=\int \frac{d s}{L}+a_{1}=\ln (2 \cdot L+2 \cdot s-2 t)+a_{1}
$$

and hence, by (96)

$$
F_{2}=\frac{1}{s}\left[\int \frac{s d s}{L}+a_{2}\right]=\frac{1}{s}\left[L+t \int \frac{d s}{L}+a_{2}\right]
$$

Computation of the limites of the integrals for $s \rightarrow 0$ give us the integration constants:

$$
\begin{aligned}
& a_{1}=-\ln (2-2 t) \quad \text { and } \\
& a_{8}=-t \ln (2-2 t)-1
\end{aligned}
$$

Hence

$$
\begin{equation*}
F_{1}=\ln \left(\frac{L+S-t}{1-t}\right)=\ln \left(1+\frac{2 s}{1-s+L}\right), \tag{99}
\end{equation*}
$$

which can be verified by multiplying the numerator and the denominator by ( $1-\mathrm{s}+\mathrm{L}$ ). The last expression for $F_{1}$ is the best suited for numerical use, because it avoids dividing by zero for $\psi=0$.

For $F_{2}$ we get:

$$
\begin{equation*}
F_{2}=\frac{1}{S}\left(L+t \cdot F_{1}+a_{e}\right)=\frac{1}{S}\left(L-1+t \cdot F_{1}\right) . \tag{100}
\end{equation*}
$$

The first and second derivatives of $\mathrm{F}_{1}$ and $\mathrm{F}_{2}$ becomes:

$$
\begin{align*}
\mathrm{F}_{1}^{\prime} & =\frac{(1-\mathrm{s}+\mathrm{L})}{(1+\mathrm{s}+\mathrm{L})} \cdot \frac{(-2 \mathrm{~s})(-\mathrm{s} / \mathrm{L})}{(1-\mathrm{s}+\mathrm{L})^{2}}=\frac{2 \mathrm{~s}^{2}}{(1+\mathrm{S}+\mathrm{L})(1-\mathrm{s}+\mathrm{L}) \mathrm{L}}  \tag{101}\\
& =\frac{2 \mathrm{~s}^{2}}{(1+\mathrm{L}-\mathrm{ts}) \cdot \mathrm{L} \cdot 2}=\frac{\mathrm{s}^{2}}{(1+\mathrm{L}-\mathrm{ts}) \cdot \mathrm{L}}=\frac{\mathrm{s}^{2}}{\mathrm{~L} \cdot \mathrm{~N}} . \\
{F_{1}^{\prime \prime}}^{\prime \prime} & \mathrm{s}^{2}\left[\frac{-\mathrm{N}(-\mathrm{s} / \mathrm{L})-\mathrm{L}(-\mathrm{s} / \mathrm{L}-\mathrm{s})}{\mathrm{L}^{2} \mathrm{~N}^{2}}\right]=\mathrm{s}^{2}\left[\frac{1+\mathrm{L}-\mathrm{ts}+\mathrm{L}+\mathrm{L}^{2}}{\mathrm{~L}^{3} \mathrm{~N}^{2}}\right]  \tag{102}\\
& =\mathrm{s}^{3}\left[\frac{1+\mathrm{L}}{\mathrm{~L}^{2} \mathrm{~N}^{2}}+\frac{1}{\mathrm{~N} \cdot \mathrm{~L}^{3}}\right] \\
F_{9}^{\prime} & =\frac{1}{\mathrm{~s}}\left(-\mathrm{s} / \mathrm{L}+\mathrm{t} \cdot \mathrm{~F}_{1}^{\prime}+\mathrm{F}_{1}\right)=-\frac{1}{\mathrm{~L}}+\frac{\mathrm{ts}}{\mathrm{~L} \cdot \mathrm{~N}}+\mathrm{F}_{1} / \mathrm{s}  \tag{103}\\
\mathrm{~F}_{2}^{\prime \prime} & =\frac{1}{\mathrm{~s}}\left(-\mathrm{s}^{2} / \mathrm{L}^{3}+2 \mathrm{~F}_{1}^{\prime}+\mathrm{t} \cdot \mathrm{~F}_{1}^{\prime \prime}\right) \tag{104}
\end{align*}
$$

We will now derive the relations between the functions $\mathrm{F}_{1}$ and the covariance models $2,3,4$ and 5 .

## Model 2.

Using (59), (61), (66), we get:

$$
\begin{aligned}
\operatorname{cov}_{z}\left(\mathrm{~T}_{P}, \mathrm{~T}_{Q}\right) & =\mathrm{K}_{z}(\mathrm{P}, \mathrm{Q})=\sum_{\ell=0}^{\infty} \sigma_{z, \ell(\mathrm{~T}, \mathrm{~T}) \cdot \mathrm{s}^{\ell+1} \mathrm{P}_{\ell}(\mathrm{t})} \\
& =\sum_{\ell=2}^{\infty} \frac{\mathrm{R}^{2}(\ell-1)^{z}}{} \sigma_{z \ell \ell}(\Delta \mathrm{~g}, \Delta \mathrm{~g}) \mathrm{s}^{\ell+1} \mathrm{P}_{\ell}(\mathrm{t}) \\
& =\mathrm{A}_{z} \cdot \mathrm{R}^{2} \sum_{\ell=2}^{\infty} \frac{1}{\ell \cdot(\ell-1)} \mathrm{s}^{\ell+1} \mathrm{P}_{\ell}(\mathrm{t}) \\
& =\mathrm{A}_{z} \cdot \mathrm{R}^{2}\left(\sum_{\ell=2}^{\infty}\left(\frac{1}{\ell-1}-\frac{1}{\ell}\right) \mathrm{s}^{\ell+1} \mathrm{P}_{\ell}(\mathrm{t})\right)
\end{aligned}
$$

and by (73), (84) and (86)

$$
\begin{align*}
\left.\operatorname{cov}_{2}^{\prime} T_{P}, T_{Q}\right) & =A_{8} R^{3}\left[\left(F_{-1}+t s^{2}-F_{0}\right]=A_{2} R^{2}\left[s\left(M+t s \cdot l n \frac{2}{N}+t s\right)-s \cdot l n \frac{2}{N}\right]\right.  \tag{105}\\
& =A_{2} R^{3} \cdot s\left[1-L+(t s-1) \ell n \frac{2}{N}\right]
\end{align*}
$$

In the same way, we get using (64), (63), (66) and (84)

$$
\begin{align*}
\operatorname{cov}_{2}\left(T_{P}, \Delta g\right) & =\frac{R}{r} \sum_{\ell=0}^{\infty} \sigma_{2, \ell}(\Delta g, T) S^{\ell+1} P_{\ell}(t)=A_{2} \cdot R \cdot\left(\frac{R}{r}\right) \cdot \sum_{\ell=2}^{\infty} \frac{s^{\ell+1}}{\ell} P_{\ell}(t)  \tag{106}\\
& =A_{2} \cdot \frac{R^{2}}{r} \cdot\left(F_{0}-t s^{2}\right)=A_{2} \cdot \frac{R^{2}}{r} \cdot s\left(l n \frac{2}{N}-t s\right)
\end{align*}
$$

and by $(60),(66),(73),(74)$ and (84):

$$
\begin{align*}
\operatorname{cov}_{2}\left(\Delta g_{\rho}, \Delta g_{Q}\right) & =A_{2} \sum_{l=2}^{\infty} \frac{\ell-1}{\ell} s^{\ell+2} P_{l}(t) \\
& =A_{2} \cdot s\left(\sum_{l=2}^{\infty} s^{\ell+1} P_{l}(t)-\sum_{l=2}^{\infty} \frac{s^{\ell+1}}{\ell} P_{l}(t)\right)  \tag{107}\\
& =A_{2} s\left(F-s-t s^{2}-\left(F_{0}-s^{2} t\right)\right)=A_{a} s^{2}\left(\frac{1}{L}-1-l n \frac{2}{N}\right)
\end{align*}
$$

The covariance functions invglving deflections of the vertical will, as mentioned above contain $\mathrm{K}_{2}^{\prime}, \mathrm{K}_{\mathfrak{g}}^{\prime \prime}$ and $-\mathrm{D}_{\mathrm{r}} \mathrm{K}_{z}^{\prime}-\frac{2}{\mathrm{r}} \mathrm{K}_{2}^{\prime}$.

Differentiating (105) gives:

$$
\begin{align*}
& K_{2}^{\prime}=A_{2} R^{2}\left(F_{-1}^{\prime}+s^{2}-F_{0}^{\prime}\right) \quad \text { and }  \tag{108}\\
& K_{2}^{\prime \prime}=A_{2} R^{2}\left(F_{-1}^{\prime \prime}-F_{0}^{\prime \prime}\right) \tag{109}
\end{align*}
$$

Because $-\mathrm{D}_{\mathrm{r}} \mathrm{K}_{2}-\frac{2}{\mathrm{r}} \mathrm{K}_{2}^{\prime}=\operatorname{cov}_{2}\left(\Delta \mathrm{~g}_{\mathrm{P}}, \mathrm{T}_{\mathrm{Q}}\right)$ we get by differentiating (106):

$$
-\mathrm{D}_{\mathrm{r}} \mathrm{~K}_{\mathrm{a}}^{\prime}-\frac{2}{\mathrm{r}} \mathrm{~K}_{\mathrm{a}}^{\prime}=\mathrm{A}_{2} \frac{\mathrm{R}^{2}}{\mathrm{r}}\left(\mathrm{~F}_{\mathrm{O}}^{\prime}-\mathrm{s}^{2}\right)
$$

Combining the three last equations with (55), (57), (51), and (53) we get the following equations, which can be evaluated using equations (88) - (91).

$$
\begin{align*}
\operatorname{cov}_{2}\left(l_{\rho}, l_{Q}\right) & =\left(\mathrm{t} \cdot \mathrm{~K}_{2}^{\prime}-\sin ^{2} \psi \cdot \mathrm{~K}_{2}^{\prime \prime}\right) /\left(\mathrm{G} \cdot \mathrm{G}^{\prime} \cdot \mathrm{r} \cdot \mathrm{r}^{\prime}\right)  \tag{111}\\
= & \mathrm{A}_{2} \cdot \frac{\mathrm{R}^{2}}{\mathrm{r} \cdot \mathrm{r}^{\prime}}\left(\mathrm{t}\left(\mathrm{~F}_{-1}^{\prime}-\mathrm{s}^{2}-\mathrm{F}_{0}^{\prime}\right)-\sin ^{2} \psi \cdot\left(\mathrm{~F}_{-1}^{\prime \prime}-\mathrm{F}_{0}^{\prime \prime}\right)\right) /\left(\mathrm{G} \cdot \mathrm{G}^{\prime}\right) \\
\operatorname{cov}_{2}\left(\mathrm{~m}_{\mathrm{P}}, \mathrm{~m}_{\mathrm{Q}}\right) & =\mathrm{K}_{\mathrm{R}}^{\prime} /\left(\mathrm{G} \cdot \mathrm{G}^{\prime} \cdot \mathrm{r} \cdot \mathrm{r}^{\prime}\right)=\mathrm{A}_{2} \frac{\mathrm{R}^{2}}{\mathrm{r} \cdot \mathrm{r}^{\prime}}\left(\mathrm{F}_{-1}^{\prime}-\mathrm{s}^{2}-\mathrm{F}_{0}^{\prime}\right) /\left(\mathrm{G} \cdot \mathrm{G}^{\prime}\right)  \tag{112}\\
\operatorname{cov}_{2}\left(\ell_{\rho}, \zeta_{Q}\right) & =\sin \psi \cdot \mathrm{K}_{\mathrm{Q}}^{\prime} /\left(\mathrm{G} \cdot \mathrm{G}^{\prime} \cdot \mathrm{r}\right)=\mathrm{A}_{2} \frac{\mathrm{R}^{2}}{\mathrm{r}}\left(\mathrm{~F}_{-1}^{\prime}-\mathrm{s}^{2}-\mathrm{F}_{0}^{\prime}\right) \cdot \sin \psi /\left(\mathrm{G} \cdot \mathrm{G}^{\prime}\right)  \tag{113}\\
\operatorname{cov}_{a}\left(\ell_{\rho}, \Delta \mathrm{g}_{Q}\right) & =\mathrm{A}_{2} \cdot \frac{\mathrm{R}^{2}}{\mathrm{r} \cdot \mathrm{r}^{\prime}} \cdot\left(\mathrm{F}_{0}^{\prime}-\mathrm{s}^{2}\right) \cdot \sin \psi \cdot \frac{1}{\mathrm{G}}=\frac{\mathrm{A}_{2} \cdot \mathrm{~s}}{\mathrm{G}} \cdot \sin \psi\left(\mathrm{~F}_{0}^{\prime}-\mathrm{s}^{2}\right) \tag{114}
\end{align*}
$$

## Model 3.

From (59), (61) and (67) we get:

$$
\begin{aligned}
\operatorname{cov}_{3}\left(\mathrm{~T}_{\mathrm{P}}, \mathrm{~T}_{\mathrm{Q}}\right) & \left.=\mathrm{K}_{3}(\mathrm{P}, \mathrm{Q})=\sum_{\ell=3}^{\infty} \frac{\mathrm{R}^{2}}{(\ell-1)^{2}} \cdot \sigma_{3, \ell}(\Delta \mathrm{~g}, \Delta \mathrm{~g}) \mathrm{s}^{\ell+1} \mathrm{P}_{\ell}^{\prime} \mathrm{t}\right) \\
& =\mathrm{A}_{3} \mathrm{R}^{2} \sum_{\ell=3}^{\infty} \frac{1}{(\ell-1)(\ell-2)} \mathrm{s}^{\ell+1} \mathrm{P}_{\ell}(\mathrm{t})=\mathrm{A}_{3} \mathrm{R}^{2} \sum_{\ell=3}^{\infty}\left(\frac{1}{\ell-2}-\frac{1}{\ell-1}\right) \mathrm{s}^{\ell+1} \mathrm{P}_{\ell}(\mathrm{t})
\end{aligned}
$$

and then using (73):

$$
\begin{equation*}
\operatorname{cov}_{3}\left(\mathrm{~T}_{\mathrm{P}}, \mathrm{~T}_{Q}\right)=\mathrm{A}_{3} R^{2} \cdot\left[\mathrm{~F}_{-2}-\left(\mathrm{F}_{-1}-\mathrm{s}^{3} \mathrm{P}_{2}(\mathrm{t})\right)\right] \tag{115}
\end{equation*}
$$

For the covariances between the gravity anomaly and the anomalous potential we get using (62), (63), (67) and (73):

$$
\begin{align*}
\operatorname{cov}_{3}\left(\mathrm{~T}_{\mathrm{P}}, \Delta \mathrm{~g}_{\mathrm{Q}}\right) & \left.=\sum_{l=3}^{\infty}\left(\frac{\mathrm{R}}{\mathrm{r}}\right)^{\prime}\right) \cdot \frac{\mathrm{R}}{\ell-1} \sigma_{\mathbf{a}}, \ell(\Delta \mathrm{g}, \Delta \mathrm{~g}) \mathrm{s}^{\ell+1} \mathrm{P}_{\ell}(\mathrm{t}) \\
& =\mathrm{A}_{3} \frac{\mathrm{R}^{2}}{\mathrm{r}^{\prime}} \sum_{l=3}^{\infty} \frac{1}{\ell-2} \mathrm{~s}^{\ell+1} \mathrm{P}_{\ell}(\mathrm{t})  \tag{116}\\
& =\mathrm{A}_{3} \frac{\mathrm{R}^{2}}{\mathrm{r}^{\prime}} \cdot \mathrm{F}_{-2} .
\end{align*}
$$

And for $\operatorname{cov}_{3}\left(\Delta \mathrm{~g}_{\mathrm{p}}, \Delta \mathrm{g}_{\mathrm{Q}}\right)$ we get using (60), (67), (71), (73) and (74):

$$
\begin{align*}
\operatorname{cov}_{3}\left(\Delta \mathrm{~g}_{\mathrm{p}}, \Delta \mathrm{~g}_{\ell}\right)= & \mathrm{A}_{3} \sum_{\ell=3}^{\infty} \frac{\ell-1}{\ell-2} \mathrm{~s}^{\ell+2} \mathrm{P}_{\ell}(\mathrm{t})=\mathrm{A}_{3} \cdot \mathrm{~s}\left[\sum_{\ell=3}^{\infty} \mathrm{s}^{\ell+1} \mathrm{P}_{\ell}(\mathrm{t})\right. \\
& \left.+\sum_{\ell=3}^{\infty} \frac{1}{\ell-2} \mathrm{~s}^{\ell+1} \mathrm{P}_{\ell}(\mathrm{t})\right]  \tag{117}\\
= & A_{3} \cdot \mathrm{~s}\left[\frac{\mathrm{~s}}{\mathrm{~L}}-\mathrm{s}-\mathrm{s}^{2} \mathrm{t}-\mathrm{s}^{3} \mathrm{P}_{\mathrm{a}}(\mathrm{t})+\mathrm{F}_{-\Omega}\right]
\end{align*}
$$

The formula (115) becomes using (86) and (87):

$$
\begin{align*}
\operatorname{cov}_{3}\left(T_{P}, \mathrm{~T}_{\mathrm{Q}}\right) & =\mathrm{A}_{3} \mathrm{R}^{2}\left[-\mathrm{s}\left(\mathrm{M}+\mathrm{ts} \cdot \ln \frac{2}{\mathrm{M}}\right)+\mathrm{s}^{3} \mathrm{P}_{\mathrm{l}}(\mathrm{t})+\mathrm{s}(\mathrm{M}(3 \mathrm{ts}+1) / 2\right. \\
& \left.\left.+\mathrm{s}^{2}\left(\mathrm{P}_{2}(\mathrm{t}) \cdot \ln \frac{2}{\mathrm{~N}}+\left(1-\mathrm{t}^{2}\right) / 4\right)\right)\right]  \tag{118}\\
& =\mathrm{A}_{3} \mathrm{R}^{2} \cdot\left[\mathrm{~s}^{3}\left(\mathrm{P}_{2}(\mathrm{t})\left(1+\ln \frac{2}{\mathrm{~N}}\right)+\frac{\sin ^{2} \psi}{4}\right)-\mathrm{s}^{2} \cdot \operatorname{ten} \frac{2}{\mathrm{M}}+\mathrm{s}(3 \mathrm{ts}-1) \cdot \frac{\mathrm{M}}{2}\right]
\end{align*}
$$

This is the correct version of the formula given by Lauritzen (1973, p. 82), in which the quantities here called $M$ and $N$ have been interchanged and the $R^{2}$ factor is missing.

Explicit expressions can be written down for (116) and (117) as well, using (86) and (87). But generally it is easier to compute the values of (86) and (87) separately and then evaluate the covariances using (115)-(117).

The derivatives necessary for the evaluation of the covariances involving deflections ((38) - (44) and (51) - (57)) becomes by differentiating (115) and (116):

$$
\begin{align*}
& K_{3}^{\prime}=A_{3} R^{8}\left[F_{-2}^{\prime}+3 s^{3} t-F_{-1}^{\prime}\right]  \tag{119}\\
& K_{3}^{\prime \prime}=A_{3} R^{2}\left[F_{-2}^{\prime \prime}+3 s^{3}-F_{-1}^{\prime \prime}\right] \quad \text { and }  \tag{120}\\
& -D_{r} K_{3}^{\prime}-\frac{2}{r} K_{3}^{\prime}=D_{t} \operatorname{cov}_{3}\left(\Delta g_{p}, T_{Q}\right)=\frac{A_{3} \cdot R^{8}}{r} \cdot F_{-2}^{\prime}, \tag{121}
\end{align*}
$$

which then can be evaluated using the formula for $F_{-1}^{\prime}, F_{-2}^{\prime}, F_{-1}^{\prime \prime}$, and $F_{-a}^{\prime \prime},(90)-(93)$. Combining the three last equations with (55), (57), (51), and (53) we get:

$$
\begin{align*}
& \operatorname{cov}_{3}\left(l_{F}, l_{Q}\right)=\left(t \cdot K_{3}^{\prime}-\sin ^{2} \psi \cdot K_{3}^{\prime \prime}\right) /\left(G \cdot G^{\prime} \cdot r \cdot r^{\prime}\right)  \tag{122}\\
& =A_{3} \cdot \mathrm{~S} \cdot\left[\mathrm{t}\left(\mathrm{~F}_{-2}^{\prime}+3 t s^{3}-\mathrm{F}_{-1}^{\prime}\right)-\sin ^{2} \psi\left(\mathrm{~F}_{-2}^{\prime \prime}+3 \mathrm{~s}^{3}-\mathrm{F}_{-1}^{\prime \prime}\right)\right] \cdot \frac{1}{\mathrm{G} \cdot \mathrm{G}^{\prime}} \\
& \operatorname{cov}_{3}\left(\mathrm{~m}_{\mathrm{p}}, \mathrm{~m}_{\mathrm{Q}}\right)=\mathrm{K}_{3}^{\prime} /\left(\mathrm{G} \cdot \mathrm{G}^{\prime} \cdot \mathrm{r} \cdot \mathrm{r}^{\prime}\right)=\mathrm{A}_{3}{ }^{2} \mathrm{~S}\left(\mathrm{~F}_{-2}^{\prime}+3 t \mathrm{~s}^{3}-\mathrm{F}_{-1}^{\prime}\right),  \tag{123}\\
& \operatorname{cov}_{3}\left(\ell_{\mathrm{p}}, \zeta_{Q}\right)=\sin \psi \cdot \mathrm{K}_{3}^{\prime} /\left(\mathrm{G}^{\prime} \mathrm{G}^{\prime} \cdot \mathrm{r}\right)=\mathrm{A}_{3} \frac{\mathrm{R}^{2}}{\mathrm{r} \cdot \mathrm{G}^{2} \mathrm{G}^{\prime}}\left(\mathrm{F}_{-2}^{\prime}+3 t s^{3}-\mathrm{F}_{-1}^{\prime}\right) \cdot \sin \psi  \tag{124}\\
& \operatorname{cov}_{3}\left(\ell_{\rho}, \Delta \mathrm{g}_{Q}\right)=\sin \psi\left(-\mathrm{D}_{\mathrm{r}}^{\prime} \mathrm{K}^{\prime}-\frac{2}{\mathrm{r}^{\prime}} \mathrm{K}^{\prime}\right) /(\mathrm{G} \cdot \mathrm{r})=\mathrm{A}_{3} \cdot \mathrm{~S} \cdot \sin \psi \cdot \mathrm{~F}_{-\mathrm{G}}^{\prime} \cdot \frac{1}{\mathrm{G}} \tag{125}
\end{align*}
$$

Model 4. Using again (59) and (61) and now (68) we get

$$
\begin{align*}
\operatorname{cov}_{4}\left(\mathrm{~T}_{\mathrm{P}}, \mathrm{~T}_{\mathrm{Q}}\right)=\mathrm{K}_{\underline{L}}(\mathrm{P}, \mathrm{Q}) & =\sum_{\ell=3}^{\infty} \frac{\mathrm{R}^{2}}{(\ell-1)^{2}} \sigma_{4}, \ell(\Delta \mathrm{~g}, \Delta \mathrm{~g}) \cdot \mathrm{s}^{\ell+1} \mathrm{P}_{\ell}(\mathrm{t})  \tag{126}\\
& =\mathrm{A}_{4} \cdot \mathrm{R}^{2} \cdot \sum_{\ell=3}^{\infty} \cdot \frac{1}{(\ell-1)(\ell-2)(\ell+\mathrm{B})} \cdot \mathrm{s}^{\ell+1} \cdot \mathrm{P}_{\ell}(\mathrm{t})
\end{align*}
$$

Unfortunately we will now have to introduce one more notation related to the degreevariances. We will define:

$$
\begin{equation*}
\tau_{k}, \ell(\mathrm{~T}, \mathrm{~T})=\sigma_{\mathrm{k}, \ell}(\mathrm{~T}, \mathrm{~T}) \cdot \frac{1}{\mathrm{~A}_{\mathrm{k}} \cdot \mathrm{R}^{2}} \tag{127}
\end{equation*}
$$

$$
\begin{align*}
& \tau_{k}, \ell(\Delta g, T)=\sigma_{k}, \ell(\Delta g, T) \cdot \frac{1}{A_{k} \cdot R} \quad \text { and }  \tag{128}\\
& \tau_{k, \ell}(\Delta g, \Delta g)=\sigma_{k}, \ell(\Delta g, \Delta g) \cdot \frac{1}{A_{k}} . \tag{129}
\end{align*}
$$

All the quantities (127)-(129) are unitless quantities, and we have e.g. using (127), (61) and (68):

$$
\mathrm{T}_{4}, \ell(\mathrm{~T}, \mathrm{~T})=\frac{\mathrm{R}^{2}}{(\ell-1)^{2}} \sigma_{4}, \ell(\Delta \mathrm{~g}, \Delta \mathrm{~g}) \cdot \frac{1}{\mathrm{~A}_{4} \cdot \mathrm{R}^{2}}=\frac{1}{(\ell-1)(\ell-2)(\ell+\mathrm{B})}
$$

This quantity can be partitioned as follows:

$$
\begin{aligned}
\tau_{4, \ell}(T, T) & =\frac{1}{\ell+B}\left[\frac{1}{\ell-2}-\frac{1}{\ell-1}\right]=\frac{1}{B+2}\left(\frac{1}{\ell-2}-\frac{1}{\ell+B}\right)-\frac{1}{B+1}\left(\frac{1}{\ell-1}-\frac{1}{\ell+B}\right) \\
& =\frac{1}{(B+2)(B+1)}\left[\frac{B+1}{\ell-2}-\frac{B+2}{\ell-1}+\frac{1}{\ell+B}\right] .
\end{aligned}
$$

hence using (126), (127), (72) and (73) we get:

$$
\begin{align*}
\operatorname{cov}_{4}\left(\mathrm{~T}_{P}, \mathrm{~T}_{Q}\right)=\mathrm{K}_{4}(\mathrm{P}, \mathrm{Q})= & \frac{\mathrm{A}_{4} \cdot \mathrm{R}^{2}}{(\mathrm{~B}+2)(\mathrm{B}+1)}\left[\sum_{\ell=3}^{\infty} \frac{\mathrm{B}+1}{\ell-2} \mathrm{~s}^{\ell+1} \mathrm{P}_{\ell}(\mathrm{t})-\sum_{\ell=3}^{\infty} \frac{\mathrm{B}+2}{\ell-1} \mathrm{~s}^{\ell+1} \mathrm{P}_{\ell}(\mathrm{t})\right. \\
& \left.+\sum_{\ell=3}^{\infty} \frac{1}{\ell+\mathrm{B}} \mathrm{~s}^{\ell+1} \mathrm{P}_{\ell}(\mathrm{t})\right] \tag{130}
\end{align*}
$$

$$
=\frac{A_{4} \cdot R^{2}}{(B+2)(B+1)}\left[(B+1) \cdot F_{-2}-(B+2)\left(F_{-1}-S^{3} P_{2}(t)\right)\right.
$$

$$
\left.+\mathrm{F}_{\mathrm{B}}-\frac{\mathrm{s}}{\mathrm{~B}}-\frac{\mathrm{s}^{2} \mathrm{t}}{\mathrm{~B}+1}-\frac{\mathrm{s}^{3} \mathrm{P}_{\mathrm{z}}(\mathrm{t})}{\mathrm{B}+2}\right]
$$

Correspondingly we get using (128), (68) and (63):

$$
\tau_{4, \ell}(\Delta \mathrm{~g}, \mathrm{~T})=\frac{1}{(\ell-2)(\ell+B)}=\frac{1}{\mathrm{~B}+2}\left[\frac{1}{\ell-2}-\frac{1}{\ell+B}\right] \text { and hence using }(64),(72)
$$

and (73):

$$
\begin{align*}
\operatorname{cov}_{4}\left(\Delta g_{p}, T_{Q}\right) & =A_{4} \frac{R^{2}}{r}\left[\sum_{l=3}^{\infty} \frac{1}{\ell-2} s^{\ell+1} P_{l}(t)-\sum_{\ell=3}^{\infty} \frac{1}{\ell+B} s^{\ell+1} P_{l}(t)\right] \cdot \frac{1}{B+2}  \tag{131}\\
& =\frac{A_{4} R^{2}}{r \cdot(B+2)}\left[F_{-2}-\left(F_{E}-\frac{s}{B}-\frac{s^{2} t}{B+1}-\frac{s^{3} P_{2}(t)}{B+2}\right)\right] .
\end{align*}
$$

For $\tau_{4}, \ell(\Delta g, \Delta g)$ we get in a similar way us ing (129) and (68):

$$
\begin{aligned}
\tau_{4} \ell(\Delta \mathrm{~g}, \Delta \mathrm{~g}) & =\frac{\ell-1}{(\ell-2)(\ell+B)}=\frac{\ell-2+1}{(\ell-2)(\ell+\mathrm{B})}=\frac{1}{\ell+\mathrm{B}}+\left(\frac{1}{\ell-2}-\frac{1}{\ell+\mathrm{B}}\right) \frac{1}{\mathrm{~B}+2} \\
& =\frac{\mathrm{B}+1}{(\mathrm{~B}+2)(\ell+\mathrm{B})}+\frac{1}{(\mathrm{~B}+2)(\ell-2)}=\frac{1}{(\mathrm{~B}+2)}\left(\frac{\mathrm{B}+1}{\ell+\mathrm{B}}+\frac{1}{\ell-2}\right)
\end{aligned}
$$

and hence using (60), (72) and (73)

$$
\begin{align*}
\operatorname{cov}_{4}\left(\Delta \mathrm{~g}_{\rho}, \Delta \mathrm{g}_{\ell}\right) & =\frac{\mathrm{A}_{4}}{(\mathrm{~B}+2)}\left(\sum_{\ell=3}^{\infty} \frac{\mathrm{B}+1}{\ell+\mathrm{B}} \mathrm{~s}^{\ell+2} \mathrm{P}_{\ell}(\mathrm{t})+\sum_{\ell=3}^{\infty} \frac{1}{\ell-2} \mathrm{~s}^{\ell+2} \mathrm{P}_{\ell}(\mathrm{t})\right)  \tag{132}\\
& =\frac{\mathrm{A}_{4} \cdot \mathrm{~s}}{(\mathrm{~B}+2)}\left[(\mathrm{B}+1)\left(\mathrm{F}_{\mathrm{B}}-\frac{\mathrm{s}}{\mathrm{~B}}-\frac{\mathrm{s}^{2} \mathrm{t}}{\mathrm{~B}+1}-\frac{\mathrm{s}^{3} \mathrm{P}_{2}(\mathrm{t})}{\mathrm{B}+2}\right)+\mathrm{E}_{2}\right]
\end{align*}
$$

We will now differentiate (130) and (131) getting the formula necessary for the computation of the covariances involving deflections;

$$
\begin{align*}
& K_{4}^{\prime}=\frac{A_{4} \cdot R^{2}}{(B+2)(B+1)}\left[(B+1) F_{-2}^{\prime}-(B+2)\left(F_{-1}^{\prime}-3 t s^{3}\right)+F_{B}^{\prime}-\frac{s^{2}}{B+1}-\frac{3 s^{3} t}{B+2}\right]  \tag{133}\\
& K_{4}^{\prime \prime}=\frac{A_{4} \cdot R^{2}}{(B+2)(B+1)}\left[(B+1) F_{-2}^{\prime \prime}-(B+2)\left(F_{-1}^{\prime \prime}-3 S^{3}\right)+F_{B}^{\prime \prime}-\frac{3 s^{3}}{B+2}\right]  \tag{134}\\
& -D_{r} K^{\prime}-\frac{2}{r} K_{4}^{\prime}=D_{t}\left(\operatorname{cov}\left(\Delta g_{P}, T_{Q}\right)\right)=\frac{A_{4} R^{2}}{r \cdot(B+2)}\left[F_{-2}^{\prime}-\left(F_{B}^{\prime}-\frac{s^{2}}{B+1}-\frac{3 s^{3} t}{B+2}\right)\right] \tag{135}
\end{align*}
$$

The formula (133)-(135) can be evaluated using (90)-(93) and the recursion formula (97) and (98) with the "initial values" given by (101)-(104).

By using (133)-(135) we can write down the covariance functions (55), (57), (51), and (53). We get:

$$
\begin{align*}
& \operatorname{cov}_{4}\left(l_{\rho}, l_{Q}\right)=\left(t \cdot K_{4}^{\prime}-\sin ^{2} \psi K_{4}^{\prime \prime}\right) /\left(G \cdot G^{\prime} \cdot r \cdot r^{\prime}\right) \\
& =\frac{A_{4} \cdot s}{(B+2)(B+1)}\left[t \cdot\left((B+1) F_{-2}^{\prime}-(B+2)\left(F_{-1}^{\prime}-3 t s^{3}\right)+F_{B}^{\prime}-\frac{s^{2}}{B+1}-\frac{3 s^{3} t}{B+2}\right)\right. \\
& \left.\div \sin ^{2} \psi\left((B+1) F_{-2}^{\prime \prime}-(B+2)\left(F_{-1}^{\prime \prime}-3 s^{3}\right)+F_{B}^{\prime \prime}-\frac{3 s^{3}}{B+2}\right)\right] \cdot \frac{1}{G \cdot G^{\prime}}, \\
& \operatorname{cov}_{4}\left(m_{P}, m_{Q}\right)=K_{4}^{\prime} /\left(G \cdot G^{\prime} \cdot r \cdot r^{\prime}\right) \\
& =\frac{A_{4} \cdot s}{(B+2)(B+1) \cdot G \cdot G^{\prime}}\left[(B+1) F_{-2}^{\prime}-(B+2)\left(F_{-1}^{\prime}-3 t s^{3}\right)+F_{B}^{\prime}-\frac{s^{2}}{B+1}-\frac{3 s^{3} t}{B+2}\right] \\
& \operatorname{cov}_{4}\left(\ell_{p}, \zeta_{Q}\right)=\sin \psi \cdot K_{4}^{\prime} /\left(G \cdot G^{\prime} \cdot r\right)  \tag{138}\\
& =\frac{A_{4} \cdot R^{2}}{(B+2)(B+1) r \cdot G \cdot G^{\prime}} \cdot\left[(B+1) F_{-2}^{\prime}-(B+2)\left(F_{-1}^{\prime}-3 t s^{3}\right)+F_{B}^{\prime}\right. \\
& \left.-\frac{s^{2}}{B+1}-\frac{3 s^{3} t}{B+2}\right]
\end{align*}
$$

and finally:

$$
\begin{align*}
\operatorname{cov}_{4}\left(\ell_{P}, \Delta G_{Q}\right) & =\sin \psi \cdot\left(-D_{\mathrm{r}}^{\prime} \cdot \mathrm{K}_{4}^{\prime}-\frac{2}{\mathrm{r}} \mathrm{~K}_{4}^{\prime}\right) \cdot \frac{1}{\mathrm{G} \cdot \mathrm{r}}  \tag{139}\\
& =\frac{A_{4} \cdot \sin \psi \cdot \mathrm{~s}}{(\mathrm{~B}+2) \cdot \mathrm{G}}\left[\mathrm{~F}_{-2}^{\prime}-\left(\mathrm{F}_{\mathrm{B}}^{\prime}-\frac{\mathrm{s}^{2}}{\mathrm{~B}+1}-\frac{3 \mathrm{~s}^{3} \mathrm{t}}{\mathrm{~B}+2}\right)\right]
\end{align*}
$$

Model 5. Using (127), (61) and (69) we get:

$$
\begin{aligned}
T_{5, \ell}(\mathrm{~T}, \mathrm{~T}) & =\frac{\mathrm{R}^{2}}{(\ell-1)} \sigma_{5, \ell(\Delta \mathrm{~g}, \Delta \mathrm{~g})} \frac{1}{\mathrm{~A}_{5} \cdot \mathrm{R}^{2}}=\frac{1}{(\ell-1)(\ell-2)(\ell+\mathrm{i})(\ell+\mathrm{j})} \\
& =\frac{1}{\mathrm{j}-\mathrm{i}}\left[\frac{1}{(\ell-1)(\ell-2)(\ell+\mathrm{i})}-\frac{1}{(\ell-1)(\ell-2)(\ell+\mathrm{j})}\right]
\end{aligned}
$$

$$
\begin{aligned}
& =\frac{1}{\mathrm{j}-\mathrm{i}}\left[\left(\frac{1}{(\ell-2)(\mathrm{i}+2)}-\frac{1}{(\ell-1)(\mathrm{i}+1)}+\frac{1}{(\ell+\mathrm{i})(\mathrm{i}+1)(\mathrm{i}+2)}\right)\right. \\
& \left.\quad-\left(\frac{1}{(\ell-2)(\mathrm{j}+2)}-\frac{1}{(\ell-1)(\mathrm{j}+1)}+\frac{1}{(\ell+\mathrm{j})(\mathrm{j}+1)(\mathrm{j}+2)}\right)\right] \\
& =\frac{1}{\mathrm{j}-\mathrm{i}}\left[\frac{\mathrm{j}+2-\mathrm{i}-2}{(\ell-2)(\mathrm{i}+2)(\mathrm{j}+2)}+\frac{\mathrm{i}+1-\mathrm{j}-1}{(\ell-1)(\mathrm{j}+1)(\mathrm{i}+1)}+\frac{1}{(\ell+\mathrm{i})(\ell+1)(\mathrm{i}+2)}\right. \\
& \left.\quad-\frac{1}{(\ell+\mathrm{j})(\mathrm{j}+1)(\mathrm{j}+2)}\right] \\
& =\frac{1}{(\ell-2)(\mathrm{i}+2)(\mathrm{j}+2)}-\frac{1}{(\ell-1)(\mathrm{i}+1)(\mathrm{j}+1)}+\left[\frac{1}{(\ell+\mathrm{i})(\mathrm{i}+1)(\mathrm{i}+2)}\right. \\
& \left.\quad-\frac{1}{(\ell+\mathrm{j})(\mathrm{j}+1)(\mathrm{j}+2)}\right] \frac{1}{\mathrm{j}-\mathrm{i}},
\end{aligned}
$$

and by (128), (63) and (69)

$$
\begin{gathered}
T_{5}, \ell(\Delta \mathrm{~g}, \mathrm{~T})=\frac{1}{(\ell-2)(\ell+\mathrm{i})(\ell+\mathrm{j})}=\frac{1}{\mathrm{j}-\mathrm{i}}\left[\frac{1}{\ell-2} \cdot \frac{1}{\ell+\mathrm{i}}-\frac{1}{\ell-2} \cdot \frac{1}{\ell+\mathrm{j}}\right] \\
=\frac{1}{\mathrm{j}-\mathrm{i}}\left[\frac{1}{\mathrm{i}+2}\left(\frac{1}{\ell-2}-\frac{1}{\ell+\mathrm{i}}\right)-\frac{1}{\mathrm{j}+2}\left(\frac{1}{\ell-2}-\frac{1}{\ell+\mathrm{j}}\right)\right] \\
=\frac{1}{\ell-2}+\frac{1}{\mathrm{j}-\mathrm{i}}\left[\frac{1}{(\mathrm{j}+2)(\ell+\mathrm{j})}-\frac{1}{(\ell+\mathrm{i})(\mathrm{i}+2)}\right]
\end{gathered}
$$

and finally by (129) and (69)

$$
\begin{aligned}
\tau_{5} \ell(\Delta \mathrm{~g}, \Delta \mathrm{~g}) & =\frac{\ell-1}{(\ell-2)(\ell+\mathrm{i})(\ell+\mathrm{j})}=\frac{1}{\mathrm{j}-\mathrm{i}}\left[\frac{\mathrm{i}+1}{\mathrm{i}+2} \cdot \frac{1}{\ell+\mathrm{i}}+\frac{1}{(\ell-2)(\mathrm{i}+2)}\right. \\
& \left.-\frac{\mathrm{j}+1}{\mathrm{j}+2} \cdot \frac{1}{\ell+\mathrm{j}}-\frac{1}{(\ell-2)(\mathrm{j}+2)}\right]
\end{aligned}
$$

$$
=\frac{1}{(\ell-2)(i+2)(j \cdot 2)}+\frac{1}{j-i}\left[\frac{i+1}{i+2} \cdot \frac{1}{\ell+i}-\frac{j+1}{j+2} \cdot \frac{1}{\ell+j}\right]
$$

and hence using (59), (72) and (73) we get:

$$
\begin{align*}
& \operatorname{cov}_{5}\left(T_{P}, T_{Q}\right)= K_{5}(P, Q)=A_{5} \cdot R^{2}\left[\frac{1}{(i+2)(j+2)} \sum_{\ell=3}^{\infty} \frac{1}{\ell-2} s^{\ell+1} P_{\ell}(t)\right. \\
&-\frac{1}{(i+1)(j+1)} \sum_{\ell=3}^{\infty} \frac{1}{\ell-1} s^{\ell+1} P_{\ell}(t) \\
&+\frac{1}{j-i}\left(\frac{1}{(i+1)(i+2)} \sum_{\ell=3}^{\infty} \frac{1}{\ell+i} s^{\ell+1} P_{\ell}(t)\right.  \tag{140}\\
&\left.\left.-\frac{1}{(j+1)(j+2)} \sum_{\ell=3}^{\infty} \frac{1}{\ell+j} s^{\ell+1} P_{\ell}(t)\right)\right] \\
&=A_{5} R^{2}\left[\frac{1}{(i+2)(j+2)} F_{-2}-\frac{1}{(i+1)(j+1)}\left(F_{-1}-s^{3} P_{2}(t)\right)\right. \\
&+\frac{1}{j-i}\left(\frac{1}{(i+1)(i+2)}\left(F_{1}-\frac{s}{i}-\frac{s^{2} t}{i+1} \frac{s^{3} P_{2}(t)}{i+2}\right)-\frac{1}{(j+1)(j+2)}\left(F_{1}-\frac{s}{j}\right.\right. \\
&\left.\left.\left.\quad-\frac{s^{2} t}{j+1}-\frac{s^{3} P_{2}(t)}{j+2}\right)\right)\right],
\end{align*}
$$

by (62)

$$
\begin{align*}
\operatorname{cov}_{5}\left(\Delta g_{p}, T_{Q}\right) & =A_{5} \frac{R^{2}}{r}\left[\sum_{\ell=3}^{\infty} \frac{1}{\ell-2} s^{\ell+1} P_{\ell}(t)+\frac{1}{j-i}\left(\frac{1}{j+2} \sum_{\ell=3}^{\infty} \frac{1}{\ell+j} s^{\ell+1} P_{\ell}(t)\right.\right.  \tag{141}\\
& \left.\left.-\frac{1}{i+2} \sum_{\ell=3}^{\infty} \frac{1}{\ell+i} s^{\ell+1} P_{\ell}(t)\right)\right]
\end{align*}
$$

$$
\begin{align*}
=A_{5} & \frac{R^{2}}{r}\left[F_{-2}+\frac{1}{j-i}\left(\frac{1}{j+2}\left(F_{j}-\frac{s}{j}-\frac{s^{2} t}{j+1}-\frac{s^{2} P_{B}(t)}{j+2}\right)\right.\right.  \tag{141}\\
& \left.\left.-\frac{1}{i+2}\left(F_{q}-\frac{s}{i}-\frac{s^{2} t}{i+1}-\frac{s^{3} P_{3}(t)}{i+2}\right)\right)\right]
\end{align*}
$$

and by (60)

$$
\begin{align*}
\operatorname{cov}_{5}\left(\Delta g_{P}, \Delta g_{Q}\right)= & A_{5}\left[\frac{1}{(i+2)(j+2)} \sum_{\ell=3}^{\infty} \frac{1}{\ell-2} s^{\ell+2} P_{\ell}(t)\right.  \tag{142}\\
+ & \left.\frac{1}{j-i}\left(\frac{i+1}{i+2} \sum_{\ell=3}^{\infty} \frac{1}{\ell+i} s^{\ell+2} P_{\ell}(t)-\frac{j+1}{j+2} \sum_{\ell=3}^{\infty} \frac{1}{\ell+j} s^{\ell+2} P_{\ell}(t)\right)\right] \\
= & A_{5} \cdot S\left[\frac{1}{(i+2)(j+2)} F_{-2}+\frac{1}{j-i}\left(\frac{i+1}{i+2}\left(F_{q}-\frac{s}{i}-\frac{s^{2}}{i+1}-\frac{s^{3} P_{2}(t)}{i+2}\right)\right.\right. \\
& \left.\left.-\frac{j+1}{j+2}\left(F_{g}-\frac{s}{j}-\frac{s^{2} t}{j+1}-\frac{s^{3} P_{2}(t)}{j+2}\right)\right)\right]
\end{align*}
$$

The covariances (140)-(142) can then be evaluated using (86), (87) and the recursion formula (96) with "initial values" (99) and (100). As in the other models it is necessary to compute $K_{5}^{\prime}, \mathrm{K}_{5}^{\prime \prime}$ and $-\mathrm{D}_{\mathrm{r}} \mathrm{K}_{5}^{\prime}-\frac{2}{\mathrm{r}} \mathrm{K}_{5}^{\prime}$ to find the expressions for the covariance functions involving deflections of the vertical. The formulae can be derived by differentiating (140) and (141) and later evaluated using the proper recursion formula exactly as explained in model 4.

Note in the equations (140), (141) and (142) the denominators are equal to $j-i$, $i+2$, $j+2, i+1, j+1$. The occurence of these and similar quantities are the reason for the above mentioned restrictions on $i$ and $j$ (and B).

The above described expressions for the closed covariance functions can also be used in cases, where a set of empirical degree-variances are used in connection with degree-variances defined through one of the models (65)-(69). In this case, the basic covariance function $\operatorname{cov}\left(\mathrm{T}_{\mathrm{P}}, \mathrm{T}_{\mathrm{Q}}\right)$ is represented by, e.g.

$$
\begin{equation*}
\sum_{\ell=0}^{n} \hat{\sigma}_{\ell}(T, T) s^{\ell+1} P_{\ell}(t)+\sum_{\ell=n+1}^{\infty} \sigma_{k, l}(T, T) s^{\ell+1} P_{\ell}(t) \tag{143}
\end{equation*}
$$

where $\hat{\sigma}_{l}(T, T)$ are the empirically determined degree-variances as would be computed from equation (15). We will distinguish between the above mentioned covariance functions $\operatorname{cov}_{k}(A, B)$ and this new type of covariance function by a subscript E, i. e., $\operatorname{cov}_{E}\left(T_{p}, T_{Q}\right)$ is equal to expression (143). We rewrite (143):

$$
\begin{align*}
\operatorname{cov}_{\mathrm{E}}\left(\mathrm{~T}_{P}, \mathrm{~T}_{Q}\right) & =\sum_{\ell=0}^{n}\left(\hat{\sigma}_{\ell}(\mathrm{T}, \mathrm{~T})-\sigma_{\mathrm{k}, \ell}(\mathrm{~T}, \mathrm{~T})\right) \mathrm{s}^{\ell+1} \mathrm{P}_{\ell}(\mathrm{t})+\sum_{\ell=0}^{\infty} \sigma_{\mathrm{k}, \ell}(\mathrm{~T}, \mathrm{~T}) \mathrm{s}^{\ell+1} \mathrm{P}_{\ell}(\mathrm{t})  \tag{144}\\
& =\sum_{\ell=0}^{n}\left(\hat{\sigma}_{\ell}(\mathrm{T}, \mathrm{~T})-\sigma_{\mathrm{k}, \ell}(\mathrm{~T}, \mathrm{~T})\right) \mathrm{s}^{\ell+1} \mathrm{P}_{\ell}(\mathrm{t})+\operatorname{cov}_{\mathrm{k}}\left(\mathrm{~T}_{p}, \mathrm{~T}_{Q}\right)
\end{align*}
$$

Noting, that the relations (61) and (63) are valid for empirical degree-variances as well, we find using $(60),(63)$, and the relations (34) - (37), (51), (53), (55) and (57)

$$
\begin{align*}
& \operatorname{cov}_{E}\left(\Delta \mathrm{~g}_{\mathrm{P}}, \mathrm{~T}_{\mathrm{Q}}\right)=\sum_{\ell=0}^{\mathrm{n}}\left(\hat{\mathrm{O}}_{\ell}(\Delta \mathrm{g}, \mathrm{~T})-\sigma_{\mathrm{k}, \ell}(\Delta \mathrm{~g}, \mathrm{~T})\right) \cdot \frac{\mathrm{R}}{\mathrm{r}} \mathrm{~s}^{\ell+1} \mathrm{P}_{\ell}(\mathrm{t})+\operatorname{cov}_{\mathrm{k}}\left(\Delta \mathrm{~g}_{\mathrm{P}}, \mathrm{~T}_{\mathrm{Q}}\right),  \tag{145}\\
& \operatorname{cov}_{\mathrm{E}}\left(\Delta \mathrm{~g}_{\mathrm{p}}, \Delta \mathrm{~g}_{\mathrm{q}}\right)=\sum_{\ell=0}^{n}\left(\hat{\sigma}_{\ell}(\Delta \mathrm{g}, \Delta \mathrm{~g})-\sigma_{\mathrm{k}, \ell}(\Delta \mathrm{~g}, \Delta \mathrm{~g})\right) \mathrm{S}^{\ell+2} \mathrm{P}_{\ell}(\mathrm{t})+\operatorname{cov}_{\mathrm{k}}\left(\Delta \mathrm{~g}_{\mathrm{P}}, \Delta \mathrm{~g}_{\mathrm{q}}\right) \\
& \operatorname{cov}_{E}\left(l_{P}, l_{Q}\right)=\left(\sum_{l=0}^{n}\left(\hat{\sigma}_{l}(T, T)-\sigma_{k}, \ell(T, T)\right) s^{\ell+1}\left(t \cdot P_{l}^{\prime}(\mathrm{t})-\sin ^{2} \psi \cdot \mathrm{P}_{l}^{\prime \prime}(\mathrm{t})\right) /\left(\mathrm{G} \cdot \mathrm{G}^{\prime} \cdot \mathrm{r} \cdot \mathrm{r}^{\prime}\right)\right. \\
& +\operatorname{cov}_{k}\left(l_{p}, l_{Q}\right)  \tag{147}\\
& =\left(t \cdot \sum_{l=0}^{n}\left(\hat{\sigma}_{l}(T, T)-\sigma_{k, \ell}(T, T)\right) S^{\ell+1} P_{l}^{\prime}(t)-\sin ^{i} \psi \sum_{l=0}^{n}\left(\hat{\sigma}_{l}(T, T)-\sigma_{k_{l}, \ell}(T, T)\right)\right. \\
& \left.\cdot s^{\ell+1} P_{\ell}^{\prime \prime}(t)\right) /\left(G \cdot G^{\prime} \cdot r \cdot r^{\prime}\right)+\operatorname{cov}_{k}\left(l_{P}, l_{Q}\right)
\end{align*}
$$

$$
\begin{align*}
\operatorname{cov}_{E}\left(\mathrm{~m}_{P}, \mathrm{~m}_{Q}\right)= & \left(\sum_{l=0}^{n}\left(\hat{\sigma}_{l}(\mathrm{~T}, \mathrm{~T})-\sigma_{k}, \ell(\mathrm{~T}, \mathrm{~T})\right) \mathrm{s}^{\ell+1} \mathrm{P}_{\ell}^{\prime}(\mathrm{t})\right) /\left(\mathrm{G} \cdot \mathrm{G}^{\prime} \cdot \mathrm{r} \cdot \mathrm{r}^{\prime}\right)  \tag{148}\\
& +\operatorname{cov}_{k}\left(\mathrm{~m}_{P}, \mathrm{~m}_{Q}\right) \\
\operatorname{cov}_{E}\left(\ell_{P}, \ell_{Q}\right)= & \sin \psi\left(\sum_{l=0}^{n}\left(\hat{\sigma}_{l}(\mathrm{~T}, \mathrm{~T})-\sigma_{\mathrm{k}, l}(\mathrm{~T}, \mathrm{~T})\right) \mathrm{s}^{\ell+1} \mathrm{P}_{\ell}(\mathrm{t})\right) /\left(\mathrm{G} \cdot \mathrm{G}^{\prime} \cdot \mathrm{r}\right)+\operatorname{cov}\left(\ell_{P}, \ell_{Q}\right) \tag{149}
\end{align*}
$$

and finally

$$
\begin{align*}
\operatorname{cov}_{\varepsilon}\left(\ell_{p}, \Delta \mathrm{~g}_{Q}\right)= & \sin \psi\left(\sum_{\ell=0}^{\mathrm{R}}\left(\hat{\sigma}_{l}(\Delta \mathrm{~g}, \mathrm{~T})-\sigma_{\mathrm{k}, \ell( }(\Delta \mathrm{g}, \mathrm{~T})\right) \mathrm{s}^{\ell+1} \cdot \frac{\mathrm{R}}{\mathrm{r}^{\prime}} \cdot \mathrm{P}_{\ell}^{\prime}(\mathrm{t})\right) /(\mathrm{G} \cdot \mathrm{r})  \tag{150}\\
& +\operatorname{cov}_{\mathrm{k}}\left(\ell_{\rho}, \Delta \mathrm{g}_{\mathrm{q}}\right)
\end{align*}
$$

where $\mathrm{P}_{\ell}^{\prime}(\mathrm{t})$ and $\mathrm{P}_{\ell}^{\prime \prime}(\mathrm{t})$ are the $\ell^{\prime}$ th order Legendre polynomial differentiated with respect to $t$ one and two times respectively.

We now define $\epsilon_{l}$ through the following equations:

$$
\begin{align*}
& \epsilon_{l}(\mathrm{~T}, \mathrm{~T})=\hat{\sigma}_{\ell}(\mathrm{T}, \mathrm{~T})-\sigma_{\mathrm{k}, \ell}(\mathrm{~T}, \mathrm{~T})  \tag{151}\\
& \epsilon_{\ell}(\Delta \mathrm{g}, \mathrm{~T})=\hat{\sigma}_{\ell}(\Delta \mathrm{g}, \mathrm{~T})-\sigma_{\mathrm{k}, \ell}(\Delta \mathrm{~g}, \mathrm{~T}) \text { and }  \tag{152}\\
& \epsilon_{\ell}(\Delta \mathrm{g}, \Delta \mathrm{~g})=\hat{\sigma}_{\ell}(\Delta \mathrm{g}, \Delta \mathrm{~g})-\dot{\sigma}_{\mathrm{k}, \ell}(\Delta \mathrm{~g}, \Delta \mathrm{~g}) \tag{153}
\end{align*}
$$

We can then see, that the covariance functions (144)-(150) involves the summation of finite series.

$$
\begin{align*}
& \sum_{l=0}^{n} \epsilon_{l}(T, T) s^{\ell+1} P_{l}(t), \sum_{l=0}^{n} \epsilon_{l}(T, \Delta \mathrm{~g}) \mathrm{s}^{\ell+1} \mathrm{P}_{\ell}(\mathrm{t}), \sum_{l=0}^{n} \epsilon_{l}(\Delta \mathrm{~g}, \Delta \mathrm{~g}) \mathrm{s}^{\ell+1} \mathrm{P}_{\ell}(\mathrm{t})  \tag{154}\\
& \sum_{l=0}^{n} \epsilon_{\ell}(\mathrm{T}, \mathrm{~T}) \mathrm{s}^{\ell+1} \mathrm{P}_{\ell}^{\prime}(\mathrm{t}), \sum_{l=0}^{n} \epsilon_{\ell}(\mathrm{T}, \Delta \mathrm{~g}) \mathrm{s}^{\ell+1} \mathrm{P}_{\ell}^{\prime}(\mathrm{t}) \tag{155}
\end{align*}
$$

and

$$
\begin{equation*}
\sum_{\ell=0}^{n} \epsilon_{\ell}(T, T) s^{\ell+1} P_{l}^{\prime \prime}(\mathrm{t}) \tag{156}
\end{equation*}
$$

Recurs ion algorithms for the summation of those three types of series will be given in the next section.

Using the above developed expressions (144)-(150) it is possible to compute covariance functions of and between height anomalies, gravity anomalies and deflections of the vertical corresponding to the recommended model for the anomaly degree-variances. This is possible because we have selected the value B in Table Seven (p. 22 ) equal to the integer 24.

Using the empirical determined value of $\hat{\sigma}_{2}(\Delta \mathrm{~g}, \Delta \mathrm{~g})=7.5 \mathrm{mgal}^{2}$ (cf. Table Two) we can then, for example, write down an expression for the covariance functions of the anomalous potential:

$$
\begin{aligned}
\operatorname{cov}_{E}\left(T_{P}, T_{Q}\right) & =7.5 \cdot 10^{-10} \cdot R^{2} \cdot s^{3} \cdot P_{2}(t)+A \cdot 10^{-10} \cdot R^{2} \sum_{\ell=3}^{\infty} \frac{1}{(\ell-1)(\ell-2)(\ell+24)} \mathrm{s}^{\ell+1} \mathrm{P}_{\ell}(\mathrm{t}) \\
& =7.5 \cdot 10^{-10} \cdot \mathrm{R}^{2} \cdot \mathrm{~s}^{2} \cdot \mathrm{P}_{2}(\mathrm{t})+\operatorname{cov}_{4}\left(\mathrm{~T}_{P}, \mathrm{~T}_{P}\right),
\end{aligned}
$$

where the factor $10^{-10}$ is used to convert the covariance into units of $\mathrm{m}^{4} / \mathrm{sec}^{4}$, supposing $R$ in units of meters.

In a similar way we can write down the expressions for the covariance functions, $\operatorname{cov}_{E}\left(\Delta \mathrm{~g}_{\mathrm{P}}, \zeta_{Q}\right), \operatorname{cov}_{E}\left(\Delta \mathrm{~g}_{\mathrm{P}}, \Delta \mathrm{g}_{Q}\right), \operatorname{cov}_{E}\left(\Delta \mathrm{~g}_{\mathrm{P}}, \ell_{Q}\right), \operatorname{cov}_{E}\left(\zeta_{P}, l_{Q}\right), \operatorname{cov}_{E}\left(\ell_{P}, \ell_{Q}\right)$ and $\operatorname{cov}_{E}\left(\mathrm{~m}_{\mathrm{P}}, \mathrm{m}_{\mathrm{Q}}\right)$.

We have computed values of the covariances for varying spherical distance $\psi$ and for $P$ and Q lying on the surface of the Earth and 500 km above the surface of the Earth respectively. See tables 9 and 10 and figures 3-9.

The radius of the Bjerhammar sphere, R , has been determined as:

$$
\mathrm{R}=\sqrt{\mathrm{s}_{\text {table } 7}} \cdot \mathrm{R}_{\mathrm{g}}=(0.999617)^{\frac{1}{2}} \cdot 6371.0 \mathrm{~km}=6369.8 \mathrm{~km}
$$

The quantities $r$ and $r^{\prime}$ are computed by adding the actual height above the reference ellipsoid (here 0 and 500 km ) to the adopted mean Earth radius, $R_{0}$.

The subroutine presented in the appendix has been used for the computation of the given values.

Table 9
Covariance between various quantities computed using the anomaly degree variances of model 4 and $\sigma_{8}(\Delta \mathrm{~g}, \Delta \mathrm{~g})=7.5 \mathrm{mgal}^{2}$ at the surface of the sphere
approximating the earth $\left(\mathrm{R}_{\mathrm{e}}=6371 \mathrm{~km}\right)$.

|  |  |  |  | Covarian | Between |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | $\psi$ | $\begin{gathered} \Delta \mathrm{g}_{\mathrm{P}}, \Delta \mathrm{~g}_{\mathrm{Q}} \\ \mathrm{mgal}^{2} \end{gathered}$ | $\begin{gathered} \Delta g_{P}, l_{Q} \\ \mathrm{mgal} \cdot \operatorname{arcsec} \end{gathered}$ | $\begin{aligned} & \Delta g_{P}, \zeta_{Q} \\ & \mathrm{mgal} \cdot \mathrm{~m} \end{aligned}$ | $\begin{gathered} \ell_{P}, \ell_{Q} \\ \operatorname{arcsec}^{2} \end{gathered}$ | $m_{p}, m_{Q}$ <br> $\operatorname{arcsec}{ }^{2}$ | $\begin{gathered} \ell_{P}, \zeta_{Q} \\ \operatorname{arcsec} \cdot \mathrm{~m} \end{gathered}$ | $\begin{gathered} \zeta_{P}, \zeta_{Q} \\ \mathrm{~m}^{2} \end{gathered}$ |
| $0^{\circ}$ | 0.01 | 1795.0 | 0.0 | 452.3 | 45.3 | 45.3 | 0.0 | 926.1 |
| 0 | 30.0 | 801.8 | 67.3 | 434.8 | 19.2 | 27.1 | 7.3 | 925.0 |
| 1 | 0.0 | 572.7 | 59.9 | 417.7 | 14.1 | 21.7 | 11.7 | 922.4 |
| 1 | 30.0 | 452.6 | 54.2 | 402.3 | 11.5 | 18.7 | 15.1 | 918.8 |
| 2 | 0.0 | 375.5 | 49.8 | 388.3 | 9.8 | 16.7 | 18.0 | 914.3 |
| 2 | 30.0 | 320.9 | 46.2 | 375.4 | 8.6 | 15.2 | 20.4 | 909.1 |
| 3 | 0.0 | 279.9 | 43.3 | 363.3 | 7.7 | 14.0 | 22.6 | 903.3 |
| 3 | 30.0 | 247.6 | 40.9 | 352.0 | 7.0 | 13.1 | 24.6 | 896.9 |
| 4 | 0.0 | 221.6 | 38.8 | 341.3 | 6.4 | 12.3 | 26.4 | 890.0 |
| 5 | 0.0 | 181.9 | 35.4 | 321.3 | 5.5 | 11.0 | 29.6 | 874.9 |
| 6 | 0.0 | 152.8 | 32.8 | 303.0 | 4.8 | 10.0 | 32.4 | 858.2 |
| 8 | 0.0 | 112.8 | 28.9 | 269.8 | 3.7 | 8.6 | 36.9 | 820.7 |
| 10 | 0.0 | 86.1 | 26.2 | 240.2 | 2.9 | 7.5 | 40.5 | 778.9 |
| 12 | 0.0 | 66.9 | 24.0 | 213.2 | 2.3 | 6.7 | 43.3 | 733.7 |
| 14 | 0.0 | 52.2 | 22.3 | 188.3 | 1.7 | 0.1 | 45.4 | 685.9 |
| 16 | 0.0 | 40.6 | 20.8 | 165.1 | 1.2 | 5.5 | 47.0 | 636.0 |
| 18 | 0.0 | 31.1 | 19.4 | 143.4 | 0.7 | 5.0 | 48.0 | 584.8 |
| 20 | 0.0 | 23.2 | 18.2 | 123.2 | 0.3 | 4.6 | 48.6 | 532.7 |
| 22 | 0.0 | 16.5 | 17.0 | 104.2 | -0.0 | 4.2 | 48.7 | 480.2 |
| 24 | 0.0 | 10.9 | 15.9 | 86.5 | -0.4 | 3.9 | 48.5 | 427.7 |
| 26 | 0.0 | 6.0 | 14.8 | 69.9 | -0.7 | 3.5 | 47.9 | 375.7 |
| 28 | 0.0 | 1.8 | 13.8 | 54.5 | $-1.0$ | 3.2 | 47.0 | 324.5 |
| 30 | 0.0 | -1.7 | 12.7 | 40.2 | -1.2 | 3.0 | 45.8 | 274.4 |
| 35 | 0.0 | -8.6 | 10.3 | 9.2 | -1.8 | 2.4 | 41.7 | 156.2 |
| 40 | 0.0 | -12.9 | 7.9 | -15.? | -2.2 | 1.8 | 36.3 | 50.9 |
| 45 | 0.0 | -15.4 | 5.6 | -33.3 | -2.4 | 1.4 | 30.1 | -38.7 |
| 50 | 0.0 | -16.3 | 3.5 | -45.6 | -2. 5 | 1.0 | 23.4 | -110.8 |
| 55 | 0.0 | -15.9 | 1.7 | -52.6 | -2.5 | 0.7 | 16.6 | $-164.7$ |
| 60 | 0.0 | -14.5 | 0.0 | -54.8 | -2.4 | 0.4 | 10.0 | -200.5 |
| 65 | 0.0 | -12.4 | -1.3 | -53.0 | -2.2 | 0.1 | 3.9 | -219.2 |
| 70 | 0.0 | -9.8 | -2.4 | -47.9 | -1.9 | -0.1 | -1.5 | -222.2 |
| 75 | 0.0 | -6.9 | -3.2 | -40.3 | -1.5 | -0.2 | -6. 1 | -211.8 |
| 80 | 0.0 | -4.0 | -3.6 | -31.0 | -1.2 | -0.3 | -9.7 | -190.3 |
| 85 | 0.0 | -1.1 | -3.8 | -20.9 | -0.8 | -0.4 | $-12.3$ | -160.4 |
| 90 | 0.0 | 1.6 | -3.8 | -10.6 | -0.4 | -0.4 | -13.9 | -124.9 |
| 95 | 0.0 | 3.9 | -3.5 | -0.8 | -0.1 | -0.5 | -14.5 | -86.5 |
| 00 | 0.0 | 5.7 | -3.0 | 7.9 | 0.2 | -0.5 | -14.3 | -47.5 |
| 05 | 0.0 | 6.9 | -2.4 | 15.2 | 0.5 | -0.4 | -13.3 | $-10.2$ |
| 10 | 0.0 | 7.6 | -1.7 | 20.7 | 0.7 | -0.4 | -11.7 | 23.7 |
| 15 | 0.0 | 7.7 | -0.9 | 24.1 | 0.8 | -0.3 | -9.7 | 52.7 |
| 20 | 0.0 | 7.2 | -0.1 | 25.5 | 0.9 | -0.3 | -7.5 | 76.0 |
| 25 | 0.0 | 6.2 | 0.6 | 24.9 | 0.9 | -0.? | -5.2 | 93.0 |
| 30 | 0.0 | 4.7 | 1.2 | 22.5 | 0.8 | -0.1 | -2.9 | 103.9 |
| 35 | 0.0 | 2.9 | 1.7 | 18.5 | 0.7 | -0.0 | -0.9 | 108.9 |
| 40 | 0.0 | 0.8 | 2.1 | 13.4 | 0.6 | 0.0 | 0.9 | 108.9 |
| 45 | 0.0 | $-1.4$ | 2.3 | 7.4 | 0.4 | 0.1 | 2.2 | 104.7 |
| 50 | 0.0 | -3.7 | 2.4 | 1.1 | 0.2 | 0.2 | 3.0 | 97.5 |
| 55 | 0.0 | -5.8 | 2.3 | -5.2 | 0.1 | 0.3 | 3.4 | 88.7 |
| 60 | 0.0 | -7.7 | 2.0 | -11.0 | -0.1 | 0.3 | 3.4 | 79.3 |
| 65 | 0.0 | -9.3 | 1.6 | -15.9 | -0.2 | 0.4 | 2.9 | 70.7 |
| 70 | 0.0 | -10.5 | 1.1 | -19.7 | -0.3 | 0.4 | 2.1 | 63.8 |
| 75 | 0.0 | -11.3 | 0.6 | -22.1 | -0.4 | 0.4 | 1.1 | 59.4 |
| 80 | 0.0 | -11.5 | 0.0 | -22.9 | -0.4 | 0.4 | 0.0 | 57.8 |

Table 10
Covariances between various quantities computed using the anomaly degree variances of model 4 and $\hat{\sigma}_{2}(\Delta \mathrm{~g}, \Delta \mathrm{~g})=7.5 \mathrm{mgal}$ at a height 500 km above the earth ( $r=R_{\theta}+500 \mathrm{~km}$ ).
Covariances Between

|  | $\psi$ | $\begin{gathered} \Delta \mathrm{g}_{\mathrm{P}}, \Delta \mathrm{~g}_{\mathrm{a}} \\ \mathrm{mgal}^{2} \end{gathered}$ | $\begin{gathered} \Delta g_{P}, l_{Q} \\ \mathrm{mgal} \cdot \operatorname{arcsec} \end{gathered}$ | $\begin{gathered} \Delta \mathrm{g}_{\rho} \zeta_{\mathrm{Q}} \\ \mathrm{mgal} \cdot \mathrm{~m} \end{gathered}$ | $\begin{gathered} \ell_{p}, \ell_{Q} \\ \operatorname{arcsec}^{2} \end{gathered}$ | $\begin{gathered} m_{p}, m_{Q} \\ \operatorname{arcsec} \sec ^{2} \end{gathered}$ | $\begin{gathered} \ell_{P}, \zeta_{Q} \\ \operatorname{arcsec} \cdot \mathrm{~m} \end{gathered}$ | $\begin{gathered} \zeta_{P}, \zeta_{Q} \\ \mathrm{~m}^{2} \end{gathered}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $0^{\circ}$ | $0.0{ }^{1}$ | 64.1 | 0.0 | 162.2 | 4.3 | 4.3 | 0.0 | 648.4 |
| 0 | 30.0 | 64.0 | 0.6 | 162.1 | 4.3 | 4.3 | 1.2 | 648.3 |
| 1 | 0.0 | 63.7 | 1.3 | 161.8 | 4.3 | 4.3 | 2.5 | $64 \% .7$ |
| 1 | 30.0 | 63.2 | 1.9 | 161.4 | 4.2 | 4.3 | 3.7 | 64.5.\% |
| 2 | 0.0 | 62.5 | 2.5 | 160.7 | 4.2 | 4.3 | 4.9 | 645.5 |
| 2 | 30.0 | 61.7 | 3.1 | 159.9 | 4.1 | 4.2 | $6 . ?$ | 643.9 |
| 3 | 0.0 | 60.7 | 3.7 | 158.9 | 4.1 | $4 . ?$ | 7.4 | 642.0 |
| 3 | 30.0 | 59.5 | 4.2 | 157.7 | 4.0 | 4.2 | 8.5 | 639.7 |
| 4 | 0.0 | 54.3 | 4.7 | 156.4 | 3.9 | 4.? | 9.7 | 637.11 |
| 5 | 0.0 | 55.5 | 5.7 | 153.4 | 3.7 | 4.1 | 1.1 .9 | 630.7 |
| 6 | 0.0 | 52.4 | 6.5 | 149.8 | 3.6 | 4.1) | 14.0 | 623.2 |
| 9 | 0.0 | 46.0 | 7.7 | 141.5 | 3.1 | 3.4 | 17.4 | 604.5 |
| 10 | 0.0 | 39.7 | 8.5 | 132.1 | 2.7 | 3.7 | 21.3 | 581.7 |
| 12 | 0.0 | 33.7 | 8.9 | 121.9 | 2.3 | 3.5 | 24.2 | 5ら5.? |
| 14 | 0.0 | 28.3 | 9.2 | 111.3 | 1.4 | 3.3 | 26.6 | 525.5 |
| 16 | 0.0 | 23.5 | 9.2 | 100.5 | 1.5 | 3.1 | 28.6 | 493.4 |
| 18 | 10.0 | 19.1 | 9.1 | 89.8 | 1.2 | 7.9 | 30.2 | $459 . ?$ |
| 20 | 0.0 | 15.3 | 8.9 | 79.2 | 0.8 | 2.4 | 31.3 | 423.4 |
| 22 | 0.0 | 11.9 | 8.7 | 68.9 | 0.5 | 2.6 | 32.1 | 346.4 |
| 24 | 0.0 | 8.9 | 8.4 | 58.8 | 0.3 | 2.4 | 32.6 | 34\%.8 |
| 26 | 0.0 | 6.3 | 8.0 | 49.2 | 0.0 | 2.? | 32.8 | 310.7 |
| 28 | 0.0 | 3.9 | 7.6 | 40.1 | -0.? | 2.1 | 32.6 | 272.7 |
| 30 | 0.0 | 1.9 | 7.2 | 31.4 | -0.4 | 1.9 | 32.2 | 235.0 |
| 35 | 0.0 | -2.2 | 6.9 | 11.9 | -0.9 | 1.6 | 30.2 | 143.4 |
| 40 | 0.0 | -5.0 | 4.8 | -4.1 | -1.2 | 1.3 | 27.1 | 60.4 |
| 45 | 0.0 | -6. 7 | 3.5 | -16.5 | -1.5 | 1.0 | 23.1 | -17.6 |
| 50 | 0.0 | $-7.6$ | 2.3 | -25.4 | -1.6 | 0.7 | 18.5 | $-73.2$ |
| 55 | 0.0 | -7.8 | 1.3 | -31.0 | -1.7 | 0.5 | 13.8 | -120.? |
| 60 | 0.0 | -7.3 | 0.3 | -33.5 | -1. $n$ | 0.3 | 4.0 | $-153.3$ |
| 65 | 0.0 | -6.5 | -0.5 | -33.4 | -1.5 | 0.1 | 4.5 | $-172.8$ |
| 70 | 0.0 | -5.3 | -1.2 | -31.0 | -1.3 | 0.0 | 0.4 | -179.7 |
| 75 | 0.0 | -3.9 | -1.7 | -27.0 | -1.1 | -0.1 | -3.2 | -175.4 |
| 80 | 0.0 | -2.5 | $-2.0$ | -21.7 | $-0.9$ | -0.? | -6.1 | -16].7 |
| 85 | 0.0 | -1.0 | -2. 2 | -15.6 | $-0.6$ | -0. 3 | -r. 3 | -140.5 |
| 90 | 0.0 | 0.3 | -2.2 | -9.3 | -0.4 | -0.3 | -9.8 | -113.9 |
| 95 | 0.0 | 1.5 | -2.0 | -3.1 | $-0.2$ | -0.3 | $-10.6$ | $-84.1$ |
| 100 | 0.0 | 2.5 | -1.8 | 2.6 | 0.1 | $-0.3$ | $-10.8$ | -52.9 |
| 105 | 0.0 | 3.2 | -1.4 | 7.5 | 0.2 | -0.3 | $-10.3$ | - ? 2. 1 |
| 110 | 0.0 | 3.6 | -1.0 | 11.3 | 0.4 | $-0.3$ | -9.4 | 6.7 |
| 115 | 0.0 | 3.7 | -0.6 | 13.9 | 0.5 | -0.3 | -8.? | 32.4 |
| 120 | 0.0 | 3.6 | -0.2 | 15.3 | 0.5 | -0.2 | $-6.7$ | 54.1 |
| 125 | 0.0 | 3.1 | 0.3 | 15.5 | 0.6 | -0.? | $-5.1$ | 71.3 |
| 130 | 0.0 | 2.5 | 0.6 | 14.6 | 0.5 | -0. 1 | -3.5 | 83.9 |
| 135 | 0.0 | 1.6 | 0.9 | 12.7 | 0.5 | -0.1 | -2.1 | 92.1 |
| 140 | 0.0 | 0.6 | $1 . ?$ | 10.0 | 0.4 | -0.0 | -0.8 | 46.1 |
| 145 | 0.0 | -0.4 | 1.3 | 6.7 | 0.3 | 0.0 | 0.3 | 96. |
| 1.50 | 0.0 | - 1.5 | 1. 3 | 3.2 | 0.2 | 0.1 | 1.0 | 44.9 |
| 155 | 0.0 | -2.5 | 1.3 | -0.4 | 0.1 | 0.1 | ]. 4 | 91.3 |
| 160 | 0.0 | -3.5 | 1.1 | -3.7 | -0.0 | 0.1 | 1.6 | 86. 8 |
| 165 | 0.0 | -4.2 | 0.9 | -6.5 | -0.1 | $0 . ?$ | 1.4 | 82.4 |
| 170 | 0.0 | -4.8 | 0.7 | $-8.7$ | -0. 2 | $0 . ?$ | 1.1 | 78.6 |
| 175 | 0.0 | -5.2 | 0.3 | -10.0 | -0. 2 | 0.7 | 0.6 | 76.1 |
| 180 | 0.0 | $-5.3$ | 0.0 | -1.0.5 | -0. 2 | 0.7 | 0.0 | 75.3 |

## Figure 3








9. Application of the Covariance Models for the Representation of Local Covariance Functions.

Local covariance functions of point or mean gravity anomalies may be estimated by formulas similar to (3) and (4) applied on the gravity data in a certain limited area. Thus, the anomalies must be centered, i.e. the mean value over the considered area will have to be subtracted.

Disregarding gravity information outs ide the cons idered area and subtraction of the local mean value correspond heuristically to disregarding the information contained in the low order harmonics.

We will here define a n'th order local (isotropic) covariance function as a covariance function, which can be derived from the covariance function of the anomalous potential (158) using the law of propagation of covariances:

$$
\begin{equation*}
\operatorname{cov}_{k}^{n}\left(T_{p}, T_{Q}\right)=K_{k}^{n}(P, Q)=\sum_{\ell={ }_{n}+1}^{\infty} \sigma_{k, \ell}(T, T) s^{\ell+1} P_{\ell}(t) \tag{158}
\end{equation*}
$$

where the superscript $n$ is the order of the local covariance function and the subscript $k$ is an integer used (as before) to distinguish between the different degree-variance models. Thus $K_{k}^{n}(P, Q)$ is in fact a special case of the models $\operatorname{cov}_{E}\left(T_{P}, T_{Q}\right)$ cons idered above, having ail degree-variances up to and inclusive of degree $n$ equal to zero. We can then rewrite (158):

$$
\begin{align*}
\mathrm{K}_{\mathrm{k}}^{\mathrm{n}}(\mathrm{P}, \mathrm{Q}) & =\sum_{\ell=0}^{\infty} \sigma_{\mathrm{x}, \ell}(\mathrm{~T}, \mathrm{~T}) \mathrm{s}^{\ell+1} \mathrm{P}_{\ell}(\mathrm{t})-\sum_{\ell=0}^{\mathrm{n}} \sigma_{\mathrm{k}, \ell}(\mathrm{~T}, \mathrm{~T}) \mathrm{s}^{\ell+1} \mathrm{P}_{\ell}(\mathrm{t})  \tag{158~A}\\
& =\sum_{\ell=0}^{\mathrm{n}}\left(-\sigma_{\mathrm{x}, \ell}(\mathrm{~T}, \mathrm{~T})\right) \mathrm{s}^{\ell+1} \mathrm{P}_{\ell}(\mathrm{t})+\mathrm{K}_{\mathrm{k}}(\mathrm{P}, \mathrm{Q})
\end{align*}
$$

For the quantity $\epsilon(T, T)$ defined in (151) we have:
and hence:

$$
\epsilon_{\ell}(T, T)=-\sigma_{k_{1}} \ell(T, T)
$$

$$
\begin{aligned}
& \epsilon_{l}(\Delta \mathrm{~g}, \mathrm{~T})=-\sigma_{\mathrm{k}, \ell}(\Delta \mathrm{~g}, \mathrm{~T}) \quad \text { and } \\
& \epsilon_{\ell}(\Delta \mathrm{g}, \Delta \mathrm{~g})=-\sigma_{\mathrm{k}, \ell}(\Delta \mathrm{~g}, \Delta \mathrm{~g})
\end{aligned}
$$

Then we can use the expressions (145)-(150) to write down the different covariance functions derived from (158):

$$
\begin{align*}
& \operatorname{cov}_{k}^{\mathrm{n}}\left(\Delta \mathrm{~g}_{\rho}, \mathrm{T}_{\mathrm{Q}}\right)=\operatorname{cov}_{\mathrm{k}}\left(\Delta \mathrm{~g}_{\mathrm{P}}, \mathrm{~T}_{\mathrm{q}}\right)-\frac{\mathrm{R}}{\mathrm{r}} \sum_{\ell=0}^{\mathrm{n}} \sigma_{\mathrm{k}, \ell}(\Delta \mathrm{~g}, \mathrm{~T}) \mathrm{S}^{\ell+1} \mathrm{P}_{\ell}(\mathrm{t}),  \tag{159}\\
& \operatorname{cov}_{k}^{n}\left(\Delta g_{g}, \Delta \mathrm{~g}_{\mathrm{Q}}\right)=\operatorname{cov}_{\mathrm{k}}\left(\Delta \mathrm{~g}_{\mathrm{p}}, \Delta \mathrm{~g}_{\mathrm{q}}\right)-\sum_{\ell=0}^{n} \sigma_{\mathrm{k}, \ell}(\Delta \mathrm{~g}, \Delta \mathrm{~g}) \mathrm{s}{ }^{\ell+2} \mathrm{P}_{\ell}(\mathrm{t}),  \tag{160}\\
& \operatorname{cov}_{k}^{n}\left(\ell_{P}, \ell_{Q}\right)=\operatorname{cov}_{k}\left(\ell_{P}, \ell_{Q}\right)-\left(t \sum_{\ell=0}^{n} \sigma_{k, \ell}(T, T) S^{\ell+1} P_{l}^{\prime}(t)\right.  \tag{161}\\
& \left.-\sin ^{i s} \psi \sum_{\ell=0}^{n} \sigma_{\mathrm{k}, \ell}(\mathrm{~T}, \mathrm{~T}) \mathrm{s}^{\ell+1} \mathrm{P}_{\ell}^{\prime \prime}(\mathrm{t})\right) /\left(\mathrm{G} \cdot \mathrm{G}^{\prime} \cdot \mathrm{r} \cdot \mathrm{r}^{\prime}\right), \\
& \operatorname{cov}_{k}^{n}\left(m_{P}, m_{Q}\right)=\operatorname{cov}_{k}\left(m_{p}, m_{Q}\right)-\left(\sum_{\ell=0}^{n} \sigma_{\mathrm{K}, \ell}(\mathrm{~T}, \mathrm{~T}) \mathrm{s}^{\ell+1} \mathrm{P}_{\ell}^{\prime}(\mathrm{t})\right) /\left(\mathrm{G} \cdot \mathrm{G}^{4} \cdot \mathrm{r} \cdot \mathrm{r}^{\prime}\right)  \tag{162}\\
& \operatorname{cov}_{k}^{\mathrm{R}}\left(\ell_{\rho}, \zeta_{Q}\right)=\operatorname{cov}_{k}\left(\ell_{\rho}, \zeta_{Q}\right)-\sin \psi\left(\sum_{\ell=0}^{\natural} \sigma_{\mathrm{k}, \ell}(\mathrm{~T}, \mathrm{~T}) \mathrm{s}^{\ell+1} \mathrm{P}_{\ell}^{\prime}(\mathrm{t})\right) /\left(\mathrm{G}^{\circ} \cdot \mathrm{G}^{\prime} \cdot \mathrm{r}\right)  \tag{163}\\
& \text { and } \\
& \left.\operatorname{cov}_{\mathrm{k}}^{\mathrm{n}}\left(\ell_{\mathrm{p}}, \Delta \mathrm{~g}_{\mathrm{q}}\right)=\operatorname{cov}_{\mathrm{k}}\left(\ell_{\mathrm{p}}, \Delta \mathrm{~g}_{\mathrm{q}}\right)-\left(\sum_{\ell=0}^{\mathrm{n}} \sigma_{\mathrm{k}, \ell}(\Delta \mathrm{~g}, \mathrm{~T}) \mathrm{s}^{\ell+1} \mathrm{P}_{\ell}^{\prime}(\mathrm{t})\right) \sin \psi \cdot \mathrm{R} / \cdot \mathrm{G} \cdot \mathrm{r} \cdot \mathrm{r}^{\prime}\right) \tag{164}
\end{align*}
$$

The evaluation of the terms derived from the "global" covariance function $\operatorname{cov}_{k}\left(T_{P}, T_{Q}\right)$ have been explained in the preceding section. We will then have to evaluate the sums of the series (154), (155) and (156), which are series in the Legendre polynomials $\mathrm{P}_{\ell}(\mathrm{t})$, their first derivatives $\mathrm{P}_{\ell}^{\prime}(\mathrm{t})$ and the ir second derivatives $\mathrm{P}_{\ell}^{\prime \prime}(\mathrm{t})$, respectively.

This kind of series can be evaluated easily without explicity evaluating the functions $\mathrm{P}_{l}(\mathrm{t}), \mathrm{P}_{\mathrm{l}}^{\prime}(\mathrm{t})$ and $\mathrm{P}_{l}^{\prime \prime}(\mathrm{t})$. The technique is similar to the so called Horner-procedure for the evaluation of a usual polynomial:

$$
\begin{align*}
\operatorname{Pol}(t) & =a_{n} t^{n}+a_{n-1} t^{n-1}+a_{n-2} t^{n-2}+\ldots+a_{1} t+a_{0}  \tag{165}\\
& =\left(\ldots\left(\left(a_{n} t+a_{n-1}\right) t+a_{n-2}\right) t+\ldots+a_{1}\right) t+a_{0}
\end{align*}
$$

We can express this procedure through a recursion algorithm with terms:

$$
\begin{equation*}
b_{\ell}=b_{\ell+1} \cdot t+a_{\ell}, \tag{166}
\end{equation*}
$$

where the recursion starts with $b_{n+1}=0$ and where the value of Pol(t) is equal to the final recursion term $b_{0}$. The first, second (and higher order) derivatives of Pol( t ) can be evaluated using recursion as well. The recursion formulas are found by differentiating (166)

$$
\begin{align*}
& b_{l}^{\prime}=b_{l+1}^{\prime} \cdot t+b_{l+1}  \tag{167}\\
& b_{l}^{n}=b_{l+1}^{\prime \prime} \cdot t+2 b_{l+1}^{\prime} \tag{168}
\end{align*}
$$

and the derivatives will be $\operatorname{Pol}^{\prime}(t)=b_{0}^{\prime}$ and $\operatorname{Pol}^{\prime \prime}(t)=b_{0}^{\prime \prime}$
This type of algorithm, which starts by accumulating the high order terms are especially useful when $t$ is less than one, i.e. when a usual evaiuation of $t^{l}$ and multiplication with $\mathrm{a}_{\ell}$ contingently would add a small number to already accumulated terms. The essential point in the procedure is the simple fact that,

$$
t^{\ell+1}-t \cdot t^{\ell}=0
$$

i. e., that there exists a recursion formula for the function $t^{2}$.

It is well known, that we have a simple recursion formula for the Legendre polynomials $\mathrm{P}_{\ell}(\mathrm{t})$. By inspecting the formula for the covariance functions, we also note the term s $\mathrm{s}^{\ell+1}$ or $\mathrm{s}^{\ell+2}$, which becomes smaller and smaller for $\ell$ increasing, because $s$ is less than 1 . So we can hope to find simple recursion formulas for the sums (154), (155) and (156), which furthermore should behave well numerically.

A general treatment of this type of summation problem is given in Clenshaw (1955, p. 118) (also valid for many other well known series as e.g. Chebyshev series or Neumann series of Bessel functions). He regards the sum of̂ a series:

$$
\begin{equation*}
S_{n}=\sum_{\ell=0}^{n} \mathrm{a}_{\ell} \mathrm{p}_{\ell}(\mathrm{t}) \tag{169}
\end{equation*}
$$

for which there exists a three-term recursion formula between the functions $p_{l}(t)$ :

$$
\begin{equation*}
\mathrm{p}_{\ell+1}(\mathrm{t})+\mathrm{e}_{\ell} \mathrm{p}_{\ell}(\mathrm{t})+\mathrm{f}_{\ell} \mathrm{p}_{\ell-1}(\mathrm{t})=0 \tag{170}
\end{equation*}
$$

The coefficients $e_{\ell}$ and $f_{\ell}$ may be dependent of $t$ as well as on $\ell$.
He proves, that the recurs ion algorithm:

$$
\begin{equation*}
\mathrm{b}_{\ell}-\mathrm{e}_{\ell} \mathrm{b}_{\ell+1}-\mathrm{f}_{\ell+1} \mathrm{~b}_{\ell+1}+\mathrm{a}_{\ell} \tag{171}
\end{equation*}
$$

with $b_{n+a}=b_{n+1}=0$, will furnish us with the sum (169), so that

$$
\begin{equation*}
S_{n}=b_{0} p_{0}(t)+b_{1}\left(p_{1}(t)+e_{0} p_{0}(t)\right) \tag{172}
\end{equation*}
$$

after $\mathrm{n}+1$ recursion steps.
In the example above (165), we have $f_{\ell}=0, e_{\ell}=-t$ and then we get from '171)

$$
\mathrm{b}_{\ell}=\mathrm{t} \cdot \mathrm{~b}_{\ell+1}+\mathrm{a}_{\ell} \text { and } \mathrm{S}_{\mathrm{n}}=\mathrm{b}_{0}+\mathrm{b}_{1}(\mathrm{t}+(-\mathrm{t}))=\mathrm{b}_{0}
$$

as stated above in (166).
By differentiation of the recursion formula (171) and the formula (172) we get

$$
\begin{align*}
& b_{l}^{\prime}=-e_{l}^{\prime} b_{l+1}-e_{l} b_{l+1}^{\prime}-f_{l+1}^{\prime} b_{l+2}-f_{l+1} b_{l+2}^{\prime}  \tag{173a}\\
& b_{l}^{\prime \prime}=-e_{l}^{\prime \prime} b_{l+1}-e_{l} b_{l+1}^{\prime \prime}-2 e_{l}^{\prime} b_{l+1}^{\prime}-f_{l+1}^{\prime \prime} b_{l+2}-f_{l+1} b_{l+2}^{\prime \prime}-2 f_{l+1}^{\prime} b_{l+1}^{\prime} \tag{173b}
\end{align*}
$$

and

$$
\begin{align*}
S_{n}^{\prime}= & b_{0}^{\prime} p_{0}(t) \\
& +b_{\circ} p_{0}^{\prime}(t)+b_{1}^{\prime}\left(p_{1}(t)+e_{\circ} p_{0}(t)\right)  \tag{173c}\\
& +b_{1}\left(p_{1}^{\prime}(t)+e_{\circ}^{\prime} p_{0}(t)+e_{\circ} p_{o}^{\prime}(t)\right) \text { and }  \tag{173d}\\
S_{n}^{\prime \prime}=b_{0}^{\prime \prime} p_{0}(t) & +b_{0} p_{0}^{\prime \prime}(t)+2 b_{0}^{\prime} p_{0}^{\prime}(t)+b_{1}^{\prime \prime}\left(p_{1}(t) \cdot e_{\circ} p_{0}(t)\right) \\
& +2 b_{0}^{\prime}\left(p_{1}^{\prime}(t)+e_{\circ}^{\prime} p_{\circ}(t)+e_{\circ} p_{0}^{\prime}(t)\right)+b_{1}\left(p_{1}^{\prime \prime}(t)+e_{\circ}^{\prime \prime} p_{\circ}(t)+e_{\circ} p_{\circ}^{\prime \prime}(t)+2 e_{\circ}^{\prime} p_{0}^{\prime}(t)\right) .
\end{align*}
$$

For the Legendre polynomials we have the well known recursion formula:

$$
\begin{equation*}
P_{\ell+1}(t)-\frac{2 \ell+1}{\ell+1} \cdot t \cdot P_{\ell}(t)+\frac{\ell}{\ell+1} P_{\ell-1}(t)=0 \tag{174}
\end{equation*}
$$

Thus, by multiplying (174) with $\mathrm{s}^{\ell+2}$ we get:

$$
\begin{equation*}
\mathrm{s}^{\ell+2} \mathrm{P}_{\ell+1}(\mathrm{t})-\frac{2 \ell+1}{\ell+1} \mathrm{t} \cdot \mathrm{~s}\left(\mathrm{~s}^{\ell+1} \cdot \mathrm{P}_{\ell}(\mathrm{t})\right)+\frac{\ell \cdot \mathrm{s}^{2}}{\ell+1}\left(\mathrm{~s}^{\ell} \mathrm{P}_{\ell}(\mathrm{t})\right)=0, \tag{175}
\end{equation*}
$$

and thereby in fact a recursion formula for the functions

$$
\mathrm{p}_{\ell}(\mathrm{t})=\mathrm{s}^{\ell+1} \mathrm{P}_{\ell}(\mathrm{t}),
$$

which then directly can be applied on the series (154)-(156).
The quantities $e_{\ell}$ and $f_{\ell}$ in (170) becomes:

$$
\begin{align*}
& \mathrm{e}_{\ell}=-\frac{2 \ell+1}{\ell+1} \cdot \mathrm{t} \cdot \mathrm{~s} \text { and }  \tag{176a}\\
& \mathrm{f}_{\ell}=\frac{\ell \cdot \mathrm{s}^{2}}{\ell+1} \tag{176b}
\end{align*}
$$

Using (176) and that $p_{0}(t)=s \cdot P_{0}(t)=s$ and $p_{1}(t)=s^{2} t$, we get:

$$
\begin{align*}
& \mathrm{e}_{l}^{\prime}=-\frac{2 l+1}{l+1} \cdot \mathrm{~s},  \tag{177}\\
& \mathrm{e}_{l}^{\prime \prime}=0  \tag{178}\\
& \mathrm{f}_{\ell}^{\prime}=\mathrm{f}_{l}^{\prime \prime}=0  \tag{179}\\
& \mathrm{p}_{\mathrm{o}}^{\prime}(\mathrm{t})=\mathrm{p}_{0}^{\prime \prime}(\mathrm{t})=0  \tag{180}\\
& \mathrm{p}_{1}^{\prime}(\mathrm{t})=\mathrm{s}^{2} \text { and }  \tag{181}\\
& \mathrm{p}_{1}^{\prime \prime}(\mathrm{t})=0 \tag{182}
\end{align*}
$$

Then by (177)-(182) and (171)-(173) we get the following recursion formula for the quantities (154)-(156)(with a ${ }_{\ell}$ equal to $\sigma_{\mathrm{k}, \ell(\mathrm{T}, \mathrm{T}), \sigma_{\mathrm{k}, \ell}(\Delta \mathrm{g}, \mathrm{T}) \text { or } \sigma_{\mathrm{k}, \ell}(\Delta \mathrm{g}, \Delta \mathrm{g}) \cdot \mathrm{s}, ~}^{\text {s }}$ respectively):

$$
\begin{align*}
b_{\ell} & =-e_{\ell} b_{\ell+1}-f_{l+1} b_{\ell+3}+a_{\ell}  \tag{183}\\
& =\frac{2 \ell \ell+1}{\ell+1} \cdot t \cdot s \cdot b_{\ell+1}-\frac{(\ell+1) \cdot s^{2}}{(\ell+2)} b_{\ell: 2}+a_{\ell}, \\
S_{n} & =b_{0} \cdot s+b_{1}\left(s^{2} t-(s t) s\right)=b_{0} \cdot s,  \tag{184}\\
b_{\ell}^{\prime} & =\frac{2 \ell+1}{\ell+1} \cdot s^{\cdot} \cdot b_{\ell+1}+\frac{2 \ell+1}{\ell+1} \cdot t \cdot s \cdot b_{\ell+1}^{\prime}-\frac{(\ell+1)}{\ell+2} s^{2} \cdot b_{\ell+2}^{\prime}  \tag{185}\\
& =\frac{2 \ell+1}{\ell+1} s\left(b_{\ell+1}+t \cdot b_{\ell+1}^{\prime}\right)-\frac{(\ell+1) s^{2}}{(\ell+2)} b_{\ell+2}^{\prime} \\
S_{n}^{\prime} & =b_{0}^{\prime} \cdot s \tag{186}
\end{align*}
$$

and finally:

$$
\begin{align*}
& b_{\ell}^{\prime \prime}=\frac{2 \ell+1}{\ell+1} s \cdot\left(2 b_{l+1}^{\prime}+t \cdot b_{l+1}^{\prime \prime}\right)-\frac{(\ell+1) s^{2}}{(\ell+2)} b_{l+2}^{\prime \prime}, \text { with }  \tag{187}\\
& S_{n}^{\prime \prime}=b_{0}^{\prime \prime} \cdot s \tag{188}
\end{align*}
$$

We would like to point out, that the recursion formulas (183)-(188) are valid for the computation of sums of a usual Legendre-series. The formulas can be obtained from equations (183)-(188) simply by putting s equal to one.

The subroutine presented in the appendix has been used to compute $\operatorname{cov}_{4}^{20}\left(\Delta g_{p}, l_{q}\right)$, $\operatorname{cov}_{4}^{20}\left(\Delta \mathrm{~g}_{\mathrm{p}}, 5_{\mathrm{q}}\right), \operatorname{cov}_{4}^{20}\left(\ell_{p}, \ell_{\mathrm{Q}}\right), \operatorname{cov}_{4}^{d \rho}\left(\mathrm{~m}_{\mathrm{p}}, \mathrm{m}_{\mathrm{q}}\right), \operatorname{cov}_{4}^{20}\left(\ell_{p}, 5_{q}\right)$ and $\operatorname{cov}_{4}^{20}\left(5_{p}, 5_{q}\right)$ for spherical distance $\psi$ varying with $\frac{1^{\circ}}{2}$ increments from $0^{\circ}$ to $25^{\circ}$. The values are shown in table 11. (The degree-variance model defined by the constants given in Table 7 has aga in been used.)

The analytic local covariance functions model discussed above can be used to find approximations for the empirical determined local covariance functions. Such a

Table 11
Covariances between various quantities computed from the local 20 th order covariance functions using the anomaly degree variances of model 4.

## Covariances Between

|  | $\psi$ | $\begin{gathered} \Delta \mathrm{g}_{\rho}, \Delta \mathrm{g}_{\mathrm{Q}} \\ \mathrm{mgal}^{\text {e }} \end{gathered}$ | $\begin{gathered} \Delta \mathrm{g}_{P}, l_{Q} \\ \text { mgal } \cdot \operatorname{arc} \mathrm{sec} \end{gathered}$ | $\begin{gathered} \Delta \mathrm{g}_{\mathrm{p}}, \zeta_{\mathrm{Q}} \\ \mathrm{mgal} \cdot \mathrm{~m} \end{gathered}$ | $\ell_{P}, \ell_{Q_{n}}$ <br> $\operatorname{arcsec}{ }^{2}$ | $\begin{gathered} m_{p}, m_{Q} \\ \operatorname{arcsec}^{3} \end{gathered}$ | $\ell_{P}, \zeta_{Q}$ <br> $\operatorname{arcsec} \cdot m$ | $\begin{gathered} \zeta_{P}, \zeta_{Q} \\ \mathrm{~m}^{2} \end{gathered}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 0 | $0.0{ }^{\circ}$ | 1519.6 | 0.0 | 88.3 | 34.5 | 34.5 | 0.0 | 13.3 |
| 0 | 30.0 | 527.1 | 64.2 | 71.2 | 8.5 | 16.3 | 4.4 | 12.6 |
| 1 | 0.0 | 300.1 | 53.8 | 55.3 | 3.5 | 11.0 | 5.9 | 11.2 |
| 1 | 30.0 | 183.5 | 45.2 | 42.0 | 1.0 | 8.0 | 6.5 | 9.5 |
| 2 | 0.0 | 111.2 | 37.9 | 30.9 | -0.5 | 6.1 | 6.5 | 7.7 |
| 2 | 30.0 | 62.6 | 31.5 | 21.5 | -1.4 | 4.6 | 6.3 | 6.0 |
| 3 | 0.0 | 28.6 | 26.0 | 13.8 | -2.0 | 3.6 | 5.8 | 4.4 |
| 3 | 30.0 | 4.7 | 21.1 | 7.5 | -2.4 | 2.8 | 5.2 | 2.9 |
| 4 | 0.0 | -12.2 | 16.8 | 2.3 | -2.6 | 2.1 | 4.5 | 1.6 |
| 4 | 30.0 | -23.7 | 13.0 | -1.7 | -2.7 | 1.6 | 3.8 | 0.4 |
| 5 | 0.0 | -31.2 | 9.6 | -4.7 | $-2.6$ | 1.? | 3.1 | -0.5 |
| 5 | 30.0 | -35.4 | 6.7 | -6.9 | -2.5 | 0.8 | 2.4 | -1.2 |
| 6 | 0.0 | -37.2 | 4.1 | -8.3 | -2.4 | 0.5 | 1.7 | -1.8 |
| 6 | 30.0 | -37.0 | 1.9 | -9.1 | -2. 2 | 0.3 | 1.1 | -2.2 |
| 7 | 0.0 | -35.4 | 0.1 | -9.4 | -1.9 | 0.2 | 0.6 | -2.4 |
| 7 | 30.0 | -32.6 | -1.4 | -9.2 | -1.7 | 0.0 | 0.1 | -2.5 |
| 8 | 0.0 | -29.0 | -2.6 | -8.7 | -1.4 | -0.1 | -0.3 | -2.5 |
| 8 | 30.0 | -24.8 | -3.5 | -7.9 | -1.1 | -0.1 | -0.6 | -2.3 |
| 9 | 0.0 | -20.4 | -4.1 | -6.8 | -0.9 | -0.2 | -0.9 | -2.1 |
| 9 | 30.0 | -15.8 | -4.5 | -5.7 | -0.6 | -0.2 | -1.1 | -1.9 |
| 10 | 0.0 | -11.2 | -4.7 | -4.4 | -0.4 | -0.2 | -1.2 | -1.5 |
| 10 | 30.0 | -6.9 | -4.7 | -3.2 | -0.1 | -0.2 | -1.3 | -1.2 |
| 11 | 0.0 | -2.8 | -4.5 | -1.9 | 0.1 | -0.2 | -1.3 | -0.8 |
| 11 | 30.0 | 0.8 | -4.2 | -0.8 | 0.2 | -0.2 | -1.3 | -0.5 |
| 12 | 0.0 | 4.1 | -3.8 | 0.3 | 0.4 | -0. 2 | -1.2 | -0.2 |
| 12 | 30.0 | 6.8 | -3.3 | 1.3 | 0.5 | -0.2 | -1.1 | 0.2 |
| 13 | 0.0 | 8.9 | -2.7 | 2.1 | 0.6 | -0.1 | -0.9 | 0.4 |
| 13 | 30.0 | 10.6 | -2.1 | 2.7 | 0.6 | -0.1 | -0.8 | 0.7 |
| 14 | 0.0 | 11.6 | -1.5 | 3.2 | 0.7 | -0.1 | -0.6 | 0.8 |
| 14 | 30.0 | 12.2 | -0.9 | 3.5 | 0.7 | -0.1 | -0.4 | 1.0 |
| 15 | 0.0 | 12.3 | -0.4 | 3.7 | 0.7 | -0.0 | -0.2 | 1.1 |
| 15 | 30.0 | 12.0 | 0.2 | 3.7 | 0.6 | -0.0 | -0.1 | 1.1 |
| 16 | 0.0 | 11.3 | 0.7 | 3.6 | 0.6 | 0.0 | 0.1 | 1.1 |
| 16 | 30.0 | 10.2 | 1.1 | 3.4 | 0.5 | 0.0 | 0.2 | 1.1 |
| 17 | 0.0 | 8.9 | 1.4 | 3.0 | 0.4 | 0.0 | 0.4 | 1.0 |
| 17 | 30.0 | 7.4 | 1.7 | 2.6 | 0.3 | 0.0 | 0.5 | 0.9 |
| 18 | 0.0 | 5.8 | 1.9 | 2.1 | 0.2 | 0.1 | 0.5 | 0.8 |
| 18 | 30.0 | 4.0 | 2.0 | 1.6 | 0.2 | 0.1 | 0.6 | 0.6 |
| 19 | 0.0 | 2.3 | 2.0 | 1.1 | 0.1 | 0.1 | 0.6 | 0.4 |
| 19 | 30.0 | 0.6 | 2.0 | 0.6 | -0.0 | 0.1 | 0.6 | 0.3 |
| 20 | 0.0 | -0.9 | 1.9 | 0.0 | -0.1 | 0.1 | 0.6 | 0.1 |
| 20 | 30.0 | -2.4 | 1.7 | -0.5 | -0.2 | 0.1 | 0.6 | -0.1 |
| 21 | 0.0 | -3.6 | 1.5 | -0.9 | -0.2 | 0.0 | 0.5 | -0.2 |
| 21 | 30.0 | -4.7 | 1.3 | -1.3 | -0.3 | 0.0 | 0.5 | -0.3 |
| 22 | 0.0 | -5.5 | 1.0 | -1.6 | -0.3 | 0.0 | 0.4 | -0.4 |
| 22 | 30.0 | -6.1 | 0.7 | -1.8 | -0.3 | 0.0 | 0.3 | -0.5 |
| 23 | 0.0 | -6.5 | 0.5 | -2.0 | -0.3 | 0.0 | 0.2 | -0.6 |
| 23 | 30.0 | -6.6 | 0.2 | -2.1 | -0.3 | 0.0 | 0.1 | -0.6 |
| 24 | 0.0 | -6.4 | -0.1 | -2.1 | -0.3 | 0.0 | 0.0 | -0.7 |
| 24 | 30.0 | -6.1 | -0.4 | -2.0 | -0.3 | -0.0 | -0.1 | -0.6 |
| 25 | 0.0 | $-5.6$ | -0.6 | -1.9 | -0.3 | -0.0 | -0.1 | -0.6 |

covariance function (of e.g. point gravity anomalies) differ from a global covariance function by having another (generally smaller) value for spherical distance $\psi$ equal to zero and by having its first zero point occurring for a much smaller spherical distance. We will denote this distance by $\psi_{1}$, i. e. $\operatorname{cov}_{n}^{\pi}\left(\Delta g_{\rho}, \Delta g_{Q}\right)=0$ for the spherical distance between $P$ and $Q$ equal to $\psi_{1}$ and all points $P$ and $Q$ with smaller spherical distance will have a positive covariance.

Note in Table 11, that the $\psi_{1}$ value is equal to $3^{\circ} 37^{\prime}$. The first zero point for $\operatorname{cov}_{\mathrm{E}}\left(\Delta \mathrm{g}_{p}, \Delta \mathrm{~g}_{\mathrm{Q}}\right)$ was (cf. Table 9 ) equal to $29^{\circ}$. It is a general trend (which can be verified for the here discussed degree-covariance models by computational experiments), that the first zero point $\psi_{2}$ occurs at decreasing $\psi$ values for increasing order of the local covariance function. Table 12 shows the value of $\psi_{1}$ for $\operatorname{cov}_{4}^{\mathrm{n}}\left(\Delta \mathrm{g}_{p}, \Delta \mathrm{~g}_{q}\right)$ for various n values. Note in the table, that the first zero point will occur between $\psi=0$ and $\psi=90^{\circ} / \mathrm{n}$.

Table 12

The spherical distance of the first zero point $\left(\psi_{1}\right)$ for some n'th order local covariance functions of gravity anomalies. The degree-variance model used is given by the constants of Table 7.

| Order (n) | $\psi_{1}$ | Order (n) |
| ---: | ---: | ---: |
|  | $\psi_{1}$ |  |
| 20 | $3^{\circ} 37^{\prime}$ | 140 |
| 40 | $2^{\circ} 55^{\prime}$ | 160 |
| $35^{\prime}$ |  |  |
| 60 | $1^{\circ} 18^{\prime}$ | 180 |
| 80 | $59^{\prime}$ | $27^{\prime}$ |
| 100 | $48^{\prime}$ | 220 |
| 120 | $40^{\prime}$ | 240 |

By inspecting the graph of an empirically estimates local covariance function it is generally possible to find it's first zero point. The corresponding order of the local covariance function can hence be estimated by determining a $n$ greater that $90^{\circ} / \psi_{1}$ for which the two zero points are as near to each other as possible. The local covariance function $\operatorname{cov}_{k}^{n}\left(\Delta g_{p}, \Delta g_{Q}\right)$ can then be fitted to the estimated covariance function by multiplying the degree-variances of the adopted model by the ratio between the empirical determined variance and the value of cov $\mathrm{v}_{\mathrm{k}}^{\mathrm{n}}$ ( $\Delta \mathrm{g}_{\mathrm{p}}, \Delta \mathrm{g}_{\mathrm{Q}}$ ) (i.e. the value for $\psi=0^{\circ}$ ).

## 10. Representation of covariance functions of mean gravity anomalies.

We will in this section regard the covariance function of mean gravity anomalies and discuss a representation of these by a certain related point gravity covariance function.

In section 2 above we described how covariance functions of different kinds of mean gravity anomalies can be represented by a covariance function of mean gravity anomalies, meaned over a spherical cap.

The relation between the degree-variances of this spherical cap mean gravity covariance function and the degree-variances of the point anomaly covariance is (cf. equation (11)):

$$
\begin{equation*}
g_{( }(\overline{\Delta \mathrm{g}}, \overline{\Delta \mathrm{~g}})=\beta_{\ell}^{2} c_{\ell}=\beta_{\ell}^{2} \cdot \sigma_{\ell}(\Delta \mathrm{g}, \Delta \mathrm{~g}) \tag{189}
\end{equation*}
$$

where the quantities $\beta_{l}$ are given by equation (12). From this equation we have that

$$
\begin{aligned}
\beta_{\ell} & =\frac{1}{1-\cos \psi_{0}} \cdot \frac{1}{2 \ell+1}\left[P_{\ell-1}\left(\cos \psi_{0}\right)-P_{\ell+1}\left(\cos \psi_{0}\right)\right] \\
& \leq \frac{1}{1-\cos \psi_{0}} \cdot \frac{1}{2 \ell+1} \cdot 2
\end{aligned}
$$

because $\mathrm{P}_{\ell}\left(\cos \psi_{0}\right)$ is less than or equal to one for all $\psi_{0}$.
Hence (for $\psi_{0} \neq 0$ ):

$$
\lim _{\ell \rightarrow \infty} B_{\ell}=0
$$

Therefore it is not necessary to carry out the summation of the series representing $\bar{C}(P, Q)$ to the same height degree as for the series representing $C(P, Q)$. The recursion formula (172) may in this case, be well suited for computation of mean anomaly covariance values.

Unfortunately, none of the degree-variance models (65)-(69) result in closed expressions for $\bar{C}(P, Q)$. But we may get an intuitive feeling of how a possible representation can be obtained by regarding the graphs of the two point anomaly covariance functions in Figure 3 and compare these with the graph of the mean anomaly covariance function in Figure one. The graphs of the mean anomaly covariance function will either lie in between or near the graphs of the two point
anomaly covariance functions. In fact, by varying the height of the points $P$ and $Q_{\text {, points }} Q_{1}$ and $Q_{z}$ can be found for which the anomaly covariance function $C\left(Q_{1}, Q_{2}\right)$ gives a good approximation to e.g. the $1^{\circ} \times 1^{0}$ mean anomaly covariance function. Table 13 gives the mean square variation of the point anomalies for some values of the height of $Q_{1}\left(h_{Q_{1}}\right)$ and $Q_{2}\left(h_{Q_{2}}\right)$ above the surface of the Earth. (The values have been computed using the subroutine presented in the appendix).

Table 13
Table of the point anomaly variance $C\left(Q_{1} Q_{1}\right)$ for different heights $h_{Q_{1}}$ equal to $h_{Q_{2}}$

| $\begin{aligned} & \frac{h_{Q_{1}}}{\mathrm{~km}} \end{aligned}$ | $\begin{gathered} \mathrm{C}\left(\mathrm{Q}_{1}, \mathrm{Q}_{1}\right) \\ \mathrm{mgal}^{2} \end{gathered}$ | $\begin{gathered} \mathrm{h}_{\mathrm{p}_{1}} \end{gathered}$ | $\begin{gathered} C\left(Q_{1}, Q_{1}\right) \\ \text { mgal }^{8} \end{gathered}$ |
| :---: | :---: | :---: | :---: |
| 0 | 1795 | 80.0 | 343 |
| 2.5 | 1346 | 160.0 | 207 |
| 5.0 | 1148 | 320.0 | 108 |
| 10.0 | 931 | 640.0 | 46 |
| 20.0 | 715 | 1280.0 | 14 |
| 40.0 | 515 |  |  |

The height, $\mathrm{h}_{\mathrm{p}}$, corresponding to the value $\overline{\mathrm{C}}(\mathrm{P}, \mathrm{P})=919.66$ for the $1^{U} \times 1^{\circ}$ mean anomaly covariance (Table One) has been estimated to be 10.4 km .

For the point anomaly covariance functions for points $Q_{1}$ and $Q_{z}$ in this height we have:

$$
\begin{equation*}
C\left(Q_{1}, Q_{2}\right)=\sum_{l=2}^{\infty} c_{l}\left(\frac{R}{R_{e}+h_{Q}}\right)^{2 \ell+4} P_{l}(t)=\sum_{l=2}^{\infty} c_{l}\left(\frac{R}{R_{\theta}}\right)^{2 \ell+4}\left(\frac{R_{\theta}}{R_{\theta}+h_{Q}}\right)^{2 \ell+4} P_{l}(t) \tag{190}
\end{equation*}
$$

In using $C\left(Q_{1}, Q_{B}\right)$ as a representation for $\bar{C}(P, Q)$, where $P$ and $Q$ are on the surface of the Earth, we are approximating

$$
\begin{equation*}
\mathrm{C}\left(\overline{\Delta \mathrm{~g}_{\mathrm{p}}},{\left.\overline{\Delta g_{\mathrm{Q}}}\right)}=\sum_{\ell==}^{\infty} c_{l} \beta_{l}^{2}\left(\frac{\mathrm{R}}{\mathrm{R}_{\mathrm{e}}}\right)^{2 \ell+4} \mathrm{P}_{\ell}(\mathrm{t})\right. \tag{191}
\end{equation*}
$$

by

$$
\sum_{\ell=2}^{\infty} c_{l}\left(\frac{\mathrm{R}_{e}}{\mathrm{R}_{e}+\mathrm{h}_{\mathrm{Q}}}\right)^{2 \ell+4}\left(\frac{\mathrm{R}}{\mathrm{R}_{e}}\right)^{2^{\ell+4}} \mathrm{P}_{\ell}(\mathrm{t}),
$$

## i.e. we are approximating

$$
\beta_{l}^{*} \text { by }\left(\frac{R_{0}}{R_{\theta}+h_{Q}}\right)^{3 l+4}
$$

In Table 14 values of $\beta_{\ell}^{2}$ and $\left(\frac{R_{\theta}}{R_{\theta}+h_{Q}}\right)^{a \ell+4}$ are presented corresponding to the $1^{\theta} \times 1^{\nu}$ mean anomaly covariance function. The values of $R_{l}^{2}$ has been obtained by squaring the values given in Table $B$ of the appendix.

## Table 14

| Values of $B_{\ell}^{\dot{e}}$ and $\left(\frac{\mathrm{R}_{9}}{\mathrm{R}_{\mathrm{e}}+\mathrm{h}_{Q}}\right)^{2 \ell+4}$ for $\mathrm{h}_{Q}=10.4 \mathrm{~km}$ and $\psi_{0}=0 \breve{ }{ }^{\circ} 564$ |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: |
| $\ell$ | $\beta_{\ell}^{\text {Z }}$ | $\left(\frac{R_{e}}{R_{\theta}+h_{Q}}\right)^{2}$ | $\ell$ | $8_{l}^{2}$ | $\left(\frac{\mathrm{R}_{\mathrm{e}}}{\mathrm{R}_{\mathrm{e}}+\mathrm{h}_{Q}}\right)^{\cdot a \ell+4}$ |
| 2 | 0.999 | 0.987 | 60 | 0.915 | 0.817 |
| 10 | 0.997 | 0.961 | 70 | 0.885 | 0.791 |
| 20 | 0.990 | 0.931 | 80 | 0.853 | 0.765 |
| 30 | 0.978 | 0.901 | 90 | 0.817 | 0.741 |
| 40 | 0.961 | 0.872 | 100 | 0.779 | 0.717 |
| 50 | 0.940 | 0.844 | 110 | 0.738 | 0.694 |

Table 14 shows the similarity between the $\beta_{l}^{\ell}$ terms and the $\left(R_{0} /\left(R_{\theta}+h_{Q}\right)\right)$ terms for the specific $\psi_{0}$ and $h_{Q}$ chosen.

Table 15 gives values of (1) the empirical $1^{\wedge}$ equal area mean gravity anomaly covariance function as taken from Table one, and designated as cov ( $\overline{\Delta g}_{p}, \overline{\Delta g}_{q}$ ), (2) the point gravity and point height anomaly covariance functions $\operatorname{cov}_{M}\left(\Delta g_{Q_{1}}\right.$, $\left.\Delta \mathrm{g}_{Q_{B}}\right), \operatorname{cov}_{M}\left(\zeta_{Q_{1}}, \zeta_{Q_{2}}\right)$ for $h_{Q_{1}}=h_{Q_{Q 2}}=10.4 \mathrm{~km}$ and (3) the (circular cap, $\psi_{0}=0.564$ ) mean gravity and height anomaly covariance functions $\operatorname{cov}_{M}\left(\Delta \mathrm{~g}_{\mathrm{P}}, \Delta \mathrm{g}_{\mathrm{Q}}\right), \operatorname{cov}_{M}\left(\bar{\zeta}_{P}\right.$, $\left.\bar{\zeta}_{Q}\right)$. The subscript $M$ indicates, that we have used the anomaly degree variance model of table seven, with $\sigma_{2}(\Delta g, \Delta g)=7.5 \mathrm{mgal}^{\dot{\varepsilon}}$. The table shows a reasonable good argument between the empirical determined covariance function and the two functions $\operatorname{cov}_{M}\left(\overline{\Delta g}_{P},{\overline{\Delta g_{Q}}}\right)$ and $\operatorname{cov}_{M}\left(\Delta \mathrm{~g}_{Q_{1}}, \Delta \mathrm{~g}_{Q_{2}}\right)$. We also see, that it is reasonable to use the point height anomaly covariance function $\operatorname{cov}_{M}\left(\zeta_{Q_{1}}, \zeta_{Q_{B}}\right)$ for the representation of the mean height anomaly.

Values of the empirical $1^{\circ} \times 1^{\circ}$ mean gravity anomaly covariance function and related point and (spherical cap) mean gravity and height anomaly covariance functions.

| $\psi$ | $\begin{gathered} \operatorname{cov}\left(\overline{\Delta g_{P}}, \overline{\Delta g_{q}}\right) \\ \mathrm{mgal}^{2} \end{gathered}$ | $\begin{gathered} \operatorname{cov}_{M}\left(\Delta \mathrm{~g}_{\mathrm{q}_{1}}, \Delta \mathrm{~g}_{\mathrm{qa}_{\mathrm{a}}}\right) \\ \mathrm{mgal}^{\mathrm{a}} \end{gathered}$ | $\begin{gathered} \operatorname{cov}_{\mathrm{M}}\left(\overline{\Delta g_{\mathrm{P}},}, \overline{\Delta g_{\mathrm{q}}}\right) \\ \mathrm{mgal}^{2} \end{gathered}$ | $\begin{gathered} \operatorname{cov}_{M}\left(\zeta_{Q_{1}}, \zeta_{Q_{2}}\right) \\ \mathrm{m}^{2} \end{gathered}$ | $\begin{gathered} \operatorname{cov}_{M}\left(\bar{\zeta}_{P}, \bar{\zeta}_{Q}\right) \\ \mathrm{m}^{2} \end{gathered}$ |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 0.0 | 919.7 | 919.7 | 848.0 | 916.8 | 926.2 |
| 0.5 | 671.6 | 698.2 | 749.5 | 915.9 | 925.3 |
| 1.0 | 493.4 | 530.9 | 577.8 | 913.7 | 923.0 |
| 1.5 | 368.2 | 429.2 | 455.7 | 910.3 | 919.5 |
| 2.0 | 285.4 | 360.6 | 377.4 | 906.2 | 915.1 |
| 2.5 | 236.1 | 310.4 | 322.2 | 901.2 | 909.9 |
| 3.0 | 211.4 | 272.0 | 280.7 | 895.7 | 904.2 |
| 3.5 | 200.7 | 241.5 | 248.3 | 889.5 | 897.8 |
| 4.0 | 193.4 | 216.7 | 222.1 | 882.9 | 891.0 |
| 5.0 | 155.9 | 178.5 | 182.2 | 868.2 | 875.9 |
| 6.0 | 141.4 | 150.3 | 153.0 | 851.8 | 859.2 |
| 8.0 | 117.4 | 111.3 | 112.9 | 815.0 | 821.7 |
| 10.0 | 96.5 | 85.1 | 86.2 | 773.9 | 779.8 |
| 12.0 | 74.6 | 66.2 | 66.9 | 729.2 | 734.6 |
| 14.0 | 59.8 | 51.7 | 52.2 | 681.9 | 686.7 |
| 16.0 | 46.0 | 40.2 | 40.6 | 632.5 | 636.8 |
| 18.0 | 37.0 | 30.9 | 31.1 | 581.7 | 585.5 |
| 20.0 | 29.3 | 23.1 | 23.2 | 530.1 | 533.4 |
| 22.0 | 21.6 | 16.5 | 16.5 | 478.0 | 480.8 |
| 24.0 | 11.8 | 10.9 | 10.9 | 425.9 | 428.3 |
| 26.0 | 6.6 | 6.1 | 6.0 | 374.2 | 376.2 |
| 28.0 | 0.5 | 2.0 | 1.9 | 323.3 | 324.9 |
| 30.0 | -3.3 | -1.6 | 1. 7 | 273.6 | 274.8 |
| 35.0 | -12.7 | -8.3 | -8.6 | 156.0 | 156.4 |
| 40.0 | -15.4 | -12.7 | -12.9 | 51.2 | 50.9 |
| 45.0 | -11.9 | -15.1 | -15.4 | -38.0 | -38.8 |
| 50.0 | -17.9 | -16.0 | -16.3 | -109.9 | -111.0 |
| 55.0 | -17.4 | -15.7 | -15.9 | -163.6 | -164.9 |
| 60.0 | -12.5 | -14.3 | -14.5 | -199.4 | -200.8 |
| 65.0 | -9.1 | -12.3 | -12.4 | -218.0 | -219.4 |
| 70.0 | -8.8 | -9.7 | -9.8 | -221.2 | -222.5 |
| 75.0 | -6.1 | -6.9 | -6.9 | -210.9 | -212. 0 |
| 80.0 | -5.9 | -3.9 | -3.9 | -189.6 | -190.5 |
| 85.0 | -6.0 | -1.1 | -1.1 | -160.0 | -160.6 |
| 90.0 | -1.8 | 1.6 | 1.6 | -124. 7 | -125.1 |
| 95.0 | 1.5 | 3.8 | 3.9 | -86. 5 | -86.6 |
| 100.0 | 8.0 | 5.6 | 5.7 | -47. 7 | -47.6 |
| 105.0 | 9.4 | 6.8 | 6.9 | -10.5 | -10.2 |
| 110.0 | 9.2 | 7.5 | 7.6 | 23.2 | 23.7 |
| 115.0 | 10.5 | 7.6 | 7.7 | 54.8 | 52.8 |
| 120.0 | 7.0 | 7.1 | 7.2 | 75.4 | 76.1 |
| 125.0 | 5.6 | 4.6 | 6.2 | 93.9 | 93.2 |
| 130.0 | 10.8 | 0.8 | 4.7 | 103.4 | 104.0 |
| 135.0 | 8.8 | -3.6 | 2.9 | 108.5 | 109.1 |
| 140.0 | 1.8 | -5.9 | 0.8 | 108.6 | 109.1 |
| 150.0 | -6. 7 | -7.6 | -3.7 | 97.5 | 97.6 |
| 160.0 | $-6.1$ | -9.2 | -7. 7 | 79.6 | 79.5 |
| 170.0 | -17.2 | -10.4 | -10.5 | 64.2 | 63.9 |
| 180.0 | -72.8 | -11.3 | -11.5 | 58.3 | 57.9 |

Using the $5^{\circ}$ equal area mean gravity anomalies estimated from the $1^{\circ} \times 1^{\circ}$ anomalies used for the empirical covariance functions given in section 3 , and with the procedures described by Rapp (1972) we have computed empirical covariance values using equation (4). The values are shown as plusses in Figure 10. This covariance function can be represented by a spherical cap mean anomaly covariance function with $\psi_{0}=2 .^{\circ} 821$ (cf. section 2). Values are shown in Figure 10 as small circles as computed from equation (11) with the anomaly degree variance model 4 with $\sigma_{2}^{2}(\Delta \mathrm{~g}, \Delta \mathrm{~g})=7.5$ and the summation taken to $\mathrm{n}=144$. For a height of 98.45 km the point anomaly variances becomes equal to the variance of the $5^{\circ} \times 5^{\circ}$ equal area mean gravity anomalies, $298.3 \mathrm{mgal}^{2}$. The graph of this covariance function is shown as a solid line in Figure 10 as well. Again, we can observe a good agreement between the different covariance function.


Least squares collocation is a method of estimating various gravimetric dependent quantities through knowledge of the covariances between such quantities. This report has developed a new model for anomaly degree variances from which covariances for various quantities can be derived with closed formulas. Thus, these covariances between anomalies, height anomalies or (geoid undulations), deflections, etc., are all self-consistent since they are derived from a single starting point, an anomaly degree variance model.

The covariances implied by the results of this report are basically global in nature. This arises from the manner in which the anomaly degree variance model was developed where consideration was given to low degree information concerning the earth's gravitational field, and the global variances of point $1^{\circ}$ and $5^{\circ}$ gravity anomalies. It is shown, however, in Section 7 how the global covariance functions can be easily modified to obtain local covariance functions. In addition, mean covariance functions can reasonably be approximated by the point covariance functions evaluated for certain heights above the surface of the earth as explained in Section 9.

Although several anomaly degree variance models and their corresponding covariance functions are discussed, the model recommended was Model 4, defined by equation ( 25 A ) and the constants of Table Seven. Numerical results from this model are reported in the text as computed from a computer program utilizing subroutine COVA given as a Fortran program in the appendix. This latter program may be used to evaluate needed covariances to be used in any applications of least squares collocation involving anomalies, height anomalies, and deflections of the vertical.

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## Appendices

Appendix A - Table A - Original $1^{\circ}$ Covariance Results
Appendix B - Table B - Anomaly Degree Variances from the Modified $1^{\circ}$ Covariance Function

Appendix C - Computer Program for subroutine COVA.

Table A
Original $1^{\circ}$ Covariance Results


#### Abstract

Number of Product Pairs


 21828.67757.0 109192.0 156505.0 231698.0 255476.0 316123.0 352844.0 410614.0 462226.0 488882.0 541557.0 579519.0 630360.0 664455.0 702812.0 755997.0 787650.0 826509.0 860109.0 903486.0 939201.0 975421.0 1015701.0 1042704.0 1092474.0 1114590.0 1156002.0 1179137.0 1218563.0 1250747.0 1273942.0 1310993.0 1348712.0 1367757.0 1398827.0 1418730.0 1450340.0 1477700.0 1504202.0 1523277.0 1558690.0 1574082.0 1596818.0 1628569.0 1653239.0 1680734.0 1686870.0 1726841.0 1741001.0

| $\begin{gathered} \overline{\mathrm{C}}(\psi) \\ (\mathrm{mgal})^{2} \end{gathered}$ | Number of Product Pairs |
| :---: | :---: |
| 996.66 | 1764605.0 |
| 523.91 | 1786245.0 |
| 349.25 | 1806095.0 |
| 285.34 | 1826864.0 |
| 266.68 | 1844149.0 |
| 227.34 | 1868090.0 |
| 212.10 | 1887521.0 |
| 193.81 | 1914884.0 |
| 184.37 | 1916381.0 |
| 179.38 | 1946611.0 |
| 157.95 | 1961243.0 |
| 149.53 | 1975934.0 |
| 130.63 | 1990242.0 |
| 124.89 | 2009601.0 |
| 109.62 | 2026129.0 |
| 96.16 | 2027477.0 |
| 89.68 | 2050882.0 |
| 82.13 | 2051483.0 |
| 74.36 | 2071575.0 |
| 67.34 | 2079142.0 |
| 60.76 | 2077609.0 |
| 54.93 | 2094994.0 |
| 47.43 | 2105922.0 |
| 41.32 | 2101952.0 |
| 32.62 | 2105738.0 |
| 29.11 | 2102632.0 |
| 23.12 | 2116443.0 |
| 17.58 | 2110570.0 |
| 13.35 | 2115762.0 |
| 9.12 | 2117756.0 |
| 6.69 | 2113185.0 |
| 2.96 | 2117032.0 |
| 1.64 | 2105401.0 |
| -0.51 | 2103280.0 |
| -4.50 | 2107743.0 |
| -7.22 | 2102737.0 |
| -9.18 | 2087672.0 |
| -9.94 | 2084487.0 |
| -11.54 | 2085244.0 |
| -10.81 | 2071837.0 |
| -11.78 | 2045875.0 |
| -10.19 | 2055334.0 |
| -10.18 | 2047715.0 |
| -8.15 | 2033495.0 |
| -8.61 | 2015615.0 |
| -9.50 | 2013437.0 |
| -11.61 | 2004502.0 |
| -12.45 | 1980606.0 |
| -13.82 | 1968387.0 |
| -13.93 | 1961354.0 |


| $\psi^{\circ}$ | $\begin{gathered} \overline{\mathrm{C}}(\psi) \\ (\mathrm{mgal})^{2} \end{gathered}$ |
| :---: | :---: |
| 50.006 | -17.40 |
| 51.005 | -18.07 |
| 52.005 | -19.60 |
| 53.005 | -20.37 |
| 54.004 | -19.37 |
| 55.003 | $-19.95$ |
| 56.003 | -20.25 |
| 57.005 | -20.02 |
| 58.005 | -18.34 |
| 59.004 | -17.93 |
| 60.005 | -18.71 |
| 61.005 | -19.37 |
| 62.003 | -18.59 |
| 63.003 | -17.64 |
| 64.004 | -18.51 |
| 65.003 | -18.87 |
| 66.002 | -19.63 |
| 67.001 | -19.39 |
| 68.001 | -20.23 |
| 69.002 | -21.42 |
| 70.001 | -21.16 |
| 70.999 | -21.45 |
| 72.001 | -22.20 |
| 73.003 | $-21.10$ |
| 74.002 | $-19.75$ |
| 75.001 | -20.06 |
| 76.000 | -21.50 |
| 77.000 | -21.97 |
| 77.999 | -22.09 |
| 78.999 | -20.67 |
| 79.999 | -20.75 |
| 81.000 | -20.98 |
| 82.000 | -22.20 |
| 82.998 | -21.60 |
| 83.997 | -21.02 |
| 85.000 | -21.26 |
| 85.994 | -19.37 |
| 86.998 | -18.26 |
| 87.999 | -17.64 |
| 89.001. | -17.47 |
| 89.998 | -16.56 |
| 90.995 | -16.23 |
| 91.998 | -14.73 |
| 92.998 | -14.07 |
| 93.997 | -12.31 |
| 94.996 | -11.32 |
| 95.998 | -9.30 |
| 96.999 | -7.56 |
| 97.997 | $-5.72$ |
| 98.996 | -3.04 |


| 1948486.0 | 99.997 | -0.60 | 852904.0 | 149.987 | -7.80 |
| ---: | ---: | ---: | ---: | ---: | ---: |
| 1934609.0 | 100.998 | 0.99 | 822165.0 | 150.989 | -4.83 |
| 1920081.0 | 101.998 | 4.39 | 789927.0 | 151.987 | -4.22 |
| 1900164.0 | 102.997 | 4.85 | 769423.0 | 152.985 | -4.53 |
| 1893305.0 | 103.997 | 6.68 | 736964.0 | 153.986 | -4.96 |
| 1868003.0 | 104.996 | 7.21 | 717976.0 | 154.987 | -7.27 |
| 1855030.0 | 105.995 | 6.92 | 681178.0 | 155.988 | -5.17 |
| 1840927.0 | 106.994 | 8.65 | 657584.0 | 156.987 | -5.01 |
| 1829592.0 | 107.996 | 10.10 | 627832.0 | 157.987 | -3.86 |
| 1812532.0 | 108.998 | 12.34 | 600252.0 | 158.987 | -7.10 |
| 1785540.0 | 109.997 | 14.40 | 571041.0 | 159.986 | -6.74 |
| 1777238.0 | 110.995 | 16.23 | 538097.0 | 160.983 | -8.25 |
| 1763249.0 | 111.997 | 18.94 | 513528.0 | 161.980 | -17.24 |
| 1733448.0 | 112.997 | 20.09 | 485066.0 | 162.978 | -12.73 |
| 1725795.0 | 113.996 | 20.36 | 463013.0 | 163.984 | -17.85 |
| 1694670.0 | 114.994 | 22.09 | 427031.0 | 164.984 | -17.72 |
| 1689370.0 | 115.993 | 21.49 | 400307.0 | 165.982 | -17.77 |
| 1667306.0 | 116.995 | 22.06 | 376038.0 | 166.983 | -19.51 |
| 1643495.0 | 117.995 | 21.43 | 342830.0 | 167.983 | -17.32 |
| 1625432.0 | 118.993 | 20.82 | 316084.0 | 168.980 | -18.81 |
| 1611598.0 | 119.993 | 21.89 | 282284.0 | 169.969 | -23.79 |
| 1592001.0 | 120.994 | 23.02 | 263998.0 | 170.969 | -27.46 |
| 1562632.0 | 121.993 | 22.46 | 234373.0 | 171.980 | -28.13 |
| 1556133.0 | 122.993 | 20.39 | 197286.0 | 172.976 | -36.25 |
| 1529283.0 | 123.995 | 19.34 | 174514.0 | 173.965 | -40.11 |
| 1507415.0 | 124.995 | 20.00 | 138568.0 | 174.944 | -40.99 |
| 1481503.0 | 125.994 | 18.47 | 122845.0 | 175.952 | -44.19 |
| 1461365.0 | 126.993 | 18.46 | 84807.0 | 176.959 | -51.04 |
| 1437710.0 | 127.993 | 18.96 | 55162.0 | 177.912 | -54.50 |
| 1416123.0 | 128.993 | 20.20 | 31636.0 | 178.836 | -67.54 |
| 1390757.0 | 129.992 | 21.32 | 4922.0 | 179.854 | -66.82 |
| 1365957.0 | 130.991 | 19.60 |  |  |  |
| 1348160.0 | 131.991 | 16.58 |  |  |  |
| 1312392.0 | 132.990 | 14.76 |  |  |  |

Table B
Anomaly Degree Variances From The Modified $1^{\circ}$ Covariance Function

| $\ell$ | $s^{-(\ell+2)}$ | $B_{\ell}$ | $\begin{gathered} \overline{\mathrm{C}}_{l} \dagger \\ \left(\mathrm{mgal}^{\infty}\right) \end{gathered}$ | $\begin{gathered} \mathrm{C}_{\ell}{ }^{*} \\ \left(\mathrm{mgal}^{\text {a }}\right) \end{gathered}$ | $\ell$ | $s^{-(l+2)}$ | $\beta_{l}$ | $\begin{gathered} \overline{\mathrm{C}_{l}} \dagger \\ \left(\mathrm{mgal}^{i s}\right) \end{gathered}$ | $\begin{gathered} \mathrm{C}_{\ell}{ }^{*} \\ \left(\mathrm{mgal}^{2}\right) \end{gathered}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| C | 1.00077 | 1.00000 | 0.07 | 0.07 | 50 | 1.02012 | 0.96943 | 5.61 | 6.09 |
| 1 | 1.00115 | 0.99998 | 0.02 | 0.02 | 51 | 1.02051 | 0.96822 | 3.48 | 3.70 |
| 2 | 1.00153 | 0.99993 | 7.54 | 7.56 | 52 | 1.02090 | 0.96699 | 4.03 | 4.40 |
| 3 | 1.00192 | 0.99985 | 33.88 | 33.95 | 53 | 1.02129 | 0.96573 | 5.72 | 6. 27 |
| 4 | 1.00230 | 0.99976 | 19.17 | 19.23 | 54 | 1.02168 | 0.96446 | 4.00 | 4.39 |
| 5 | 1.00269 | 0.99964 | 21.57 | 21.64 | 55 | 1.02208 | 0.46316 | 4.65 | 5.1 ? |
| 6 | 1.00307 | 0.99949 | 18.87 | 18.95 | 56 | 1.02247 | 0.96183 | 4.15 | 4.59 |
| 7 | 1.00345 | 0.99932 | 18.77 | 18.86 | 57 | 1.02286 | 0.96049 | 4.34 | 4.61 |
| 8 | 1.00384 | 0.99913 | 10.42 | 10.48 | 58 | 1.02325 | 0.95912 | 5.11 | 3.68 |
| 9 | 1.00422 | 0.99891 | 11.05 | 11.12 | 59 | 1.02364 | 0.95773 | 5.46 | 6.099 |
| 10 | 1.00461 | 0.99867 | 11.43 | 11.51 | 60 | 1.02403 | 0.95632 | 2.88 | 3.23 |
| 11 | 1.00499 | 0.99840 | 14.10 | 14.22 | 61 | 1. 02443 | 0.95489 | 3.94 | 4.42 |
| 12 | 1.00538 | 0.99811 | 3. 12 | 3.15 | 62 | 1.02482 | 0.95343 | 3. 813 | 4.32 |
| 13 | 1.00576 | 0.99780 | 9.47 | 9.56 | 63 | 1.02521 | 0.95195 | 3.83 | 4.34 |
| 14 | 1.00615 | 0.99746 | 5.76 | 5.82 | 64 | 1.02561 | 0.95045 | 4.14 | 4.69 |
| 15 | 1.00653 | 0.99710 | 7. 58 | 7.67 | 65 | 1.02600 | 0.94893 | 2.51 | 2.86 |
| 16 | 1.00692 | 0.99671 | 9.93 | 10.07 | 66 | 1.02639 | 0.94739 | 5.02 | 5.74 |
| 17 | 1.00730 | 0.99630 | 8.63 | 8.76 | 67 | 1.02678 | 0.94582 | 4.66 | 5.35 |
| 18 | 1.00769 | 0.99586 | 8.26 | 8.40 | 68 | 1.02718 | 0.94424 | 4.75 | 5.48 |
| 19 | 1.00808 | 0.99540 | 7.67 | 7.80 | 69 | 1.02757 | 0.94263 | 4.91 | 5.68 |
| 2.0 | 1.00846 | 0.99492 | 1.16 | 1.18 | 70 | 1.02797 | 0.94100 | 2.11 | 2.45 |
| 21 | 1.00885 | 0.99441 | 5.81 | 5.92 | 71 | 1.02836 | 0.93935 | 3.22 | 3.75 |
| 22 | 1.00924 | 0.99388 | 4.47 | 4.56 | 72 | 1.02875 | 0.93767 | 3.94 | 4.66 |
| 23 | 1.00962 | 0.99333 | 5.93 | 6.06 | 73 | 1.02915 | 0.93598 | 4.01 | 4.71 |
| 24 | 1.01001 | 0.99275 | 6.19 | 6.34 | 74 | 1.02954 | 0.93427 | 4.72 | 5.57 |
| 25 | 1.01040 | 0.99215 | 9.24 | 9.49 | 75 | 1.02994 | 0.93253 | 4.37 | 5.1.3 |
| 26 | 1.01078 | 0.99152 | 1.96 | 2.02 | 76 | 1.03033 | 0.93077 | 2.77 | 3.29 |
| 27 | 1.01117 | 0.99087 | 4.73 | 4.87 | 77 | 1.03073 | 0.92900 | 4.88 | 5.33 |
| 28 | 1.01156 | 0.99020 | 4.17 | 4.30 | 78 | 1.03112 | 0.92720 | 4.12 | 4.44 |
| 29 | 1.01195 | 0.98950 | 4.98 | 5.15 | 79 | 1.03152 | 0.92538 | 3.01 | 3.63 |
| 30 | 1.01233 | 0.98878 | 3.89 | 4.02 | 80 | 1.03191 | 0.92354 | 6.36 | 7.70 |
| 31 | 1.01272 | 0.98803 | 4.82 | 5.00 | 81 | 1.03231 | 0.92168 | 3.93 | 4.77 |
| 32 | 1.01311 | 0.98726 | 7.78 | 8.09 | 82 | 1.03270 | 0.91980 | 5.17 | 6.31 |
| 33 | 1.01350 | 0.98647 | 6.90 | 7.19 | 83 | 1.03310 | 0.91790 | 5.42 | 6.65 |
| 34 | 1.01389 | 0.98566 | 5.89 | 6. 15 | 84 | 1.03349 | 0.91598 | 3.44 | 4.24 |
| 35 | 1.01427 | 0.98482 | 7.63 | 7.98 | 85 | 1.03389 | 0.91403 | 4.80 | 5.44 |
| 36 | 1.01466 | 0.98395 | 6.42 | 6.73 | 86 | 1.03429 | 0.91207 | 6. 24 | 7.76 |
| 37 | 1.01505 | 0.98307 | 4.56 | 4.79 | 87 | 1.03468 | 0.91009 | 5.12 | 6. 39 |
| 38 | 1.01544 | 0.98216 | 7.39 | 7.77 | 88 | 1.03508 | 0.90809 | 4.80 | 6.03 |
| 39 | 1.01583 | 0.98122 | 5.64 | 5.95 | 89 | 1.03547 | 0.90607 | 3.68 | 4.55 |
| 40 | 1.01622 | 0.98027 | 5.41 | 5.72 | 90 | 1.03587 | 0.90403 | 4.63 | 5.87 |
| 41 | 1.01661 | 0.97929 | 5.45 | 5.78 | 91 | 1.03627 | 0.90197 | 4.18 | 5.32 |
| 42 | 1.01700 | 0.97828 | 6.46 | 6.87 | 92 | 1.03667 | 0.89989 | 4.95 | 6.34 |
| 43 | 1.01739 | 0.97726 | 5.01 | 5.34 | 93 | 1.03706 | 0.89779 | 2.92 | 3.76 |
| 44 | 1.01778 | 0.97621 | 5.56 | 5.93 | 94 | 1.03746 | 0.89567 | 3.39 | 4.38 |
| 45 | 1.01817 | 0.97514 | 7. 54 | 8.08 | 95 | 1.03786 | 0.89353 | 2.06 | 2.57 |
| 46 | 1.01856 | 0.97404 | 2.81 | 3.02 | 96 | 1.03825 | 0.89137 | 4.17 | 5. +4 |
| 47 | 1.01895 | 0.97292 | 5.78 | 6.22 | 97 | 1.03865 | 0.88920 | 2.56 | 3.36 |
| 48 | 1.01934 | 0.97178 | 3.64 | 3.93 | 98 | 1.03905 | 0.88700 | 4.72 | 6.24 |
| 49 | 1.01973 | 0.97062 | 5. 51 | 5.96 | 99 | 1.03945 | 0.88479 | 2.78 | 3.69 |

[^0]- From equation (16) with $s=1$

| 100 | 1.03985 | 0.88255 | 3.81 | 5.09 |
| :--- | :--- | :--- | :--- | :--- |
| 101 | 1.04025 | 0.88030 | 2.58 | 3.46 |
| 102 | 1.04064 | 0.87803 | 3.31 | 4.47 |
| 103 | 1.04104 | 0.87575 | 2.86 | 3.89 |
| 104 | 1.04144 | 0.87344 | 3.62 | 4.95 |
| 105 | 1.04184 | 0.87111 | 2.00 | 2.74 |
| 106 | 1.04224 | 0.86877 | 3.47 | 4.79 |
| 107 | 1.04264 | 0.86641 | 3.08 | 4.27 |
| 108 | 1.04304 | 0.86403 | 3.27 | 4.58 |
| 109 | 1.04344 | 0.86164 | 3.19 | 4.48 |
| 110 | 1.04384 | 0.85922 | 3.27 | 4.62 |
| 111 | 1.04424 | 0.85679 | 3.06 | 4.35 |
| 112 | 1.04464 | 0.85434 | 4.21 | 6.02 |
| 113 | 1.04504 | 0.85188 | 3.55 | 5.12 |
| 114 | 1.04544 | 0.84939 | 2.47 | 3.58 |
| 115 | 1.04584 | 0.84690 | 2.31 | 3.37 |
| 116 | 1.04624 | 0.84438 | 3.04 | 4.46 |
| 117 | 1.04664 | 0.84184 | 2.78 | 4.10 |
| 118 | 1.04704 | 0.83929 | 2.16 | 3.21 |
| 119 | 1.04744 | 0.83673 | 2.91 | 4.36 |
| 120 | 1.04784 | 0.83414 | 1.76 | 2.66 |
| 121 | 1.04825 | 0.83154 | 2.73 | 4.13 |
| 122 | 1.04865 | 0.82893 | 1.95 | 2.98 |
| 123 | 1.04905 | 0.82630 | 1.62 | 2.49 |
| 124 | 1.04945 | 0.82365 | 2.59 | 4.00 |
| 125 | 1.04985 | 0.82099 | 1.48 | 2.31 |
| 126 | 1.05026 | 0.81831 | 2.73 | 4.28 |
| 127 | 1.05066 | 0.81561 | 1.92 | 3.03 |
| 128 | 1.05106 | 0.81290 | 2.94 | 4.68 |
| 129 | 1.05146 | 0.81017 | 1.22 | 1.96 |
| 130 | 1.05187 | 0.80743 | 2.70 | 4.36 |
| 131 | 1.05227 | 0.80468 | 1.75 | 2.84 |
| 132 | 1.05267 | 0.80191 | 2.98 | 4.87 |
| 133 | 1.05308 | 0.79912 | 1.90 | 3.13 |
| 134 | 1.05348 | 0.79632 | 2.79 | 4.63 |
| 135 | 1.05388 | 0.79351 | 1.62 | 2.71 |
| 136 | 1.05429 | 0.79068 | 1.05 | 1.77 |
| 137 | 1.05469 | 0.78783 | 1.16 | 1.98 |
| 138 | 1.05509 | 0.78498 | 2.49 | 4.26 |
| 139 | 1.05550 | 0.78210 | 0.90 | 1.55 |
| 140 | 1.05590 | 0.77922 | 1.15 | 2.00 |
| 10 |  |  |  |  |

A subroutine COVA for the computation of the covariance of and between height anomalies, gravity anomalies and the longitudinal and transveral components of the deflections of the vertical is reproduced below.

The FORTRAN IV language of the IBM $360 / 370$ system has been used.
The subroutine can only be used for the computation of covariances corresponding to the degree-variance model given by equation (68).

By the execution of a DATA statement, the quantities $s, A$ and $B$ become equal to the values given in Table Seven. It is only necessary to change the values given in the DATA statement to obtain the covariances corresponding to a degreevariance model with other values of $\mathrm{s}, \mathrm{A}$ and B . The subroutine can be used to compute covariance values corresponding to both a local n'th order covariance function and to a covariance function, which has some of the degree-variances equal to empirical determined values.

The comments given in connection with the Fortran statements of the subroutine should give all details necessary for the application of the subroutine.

C
C

C THERE ARE THREE ENTRIES TO THE SUBROUTINE, WHICH HAVE TO bE CALLED IN C THE SEQUENCE COVA, COVB aND COVC.

C BY THE CALL OF COVA, THE KIND OF COVARIANCE FUNCTION TO BE USED IS C DETERMINED. THERE ARE THREE POSSIBILITIES:
C (1) THE COVARIANCE MODEL FOUR (EQUATIONS (130)-(132) AND (136)-(139)) IS USED WITHOUT MODIFICATIONS. IN THIS CASE EPS WILL BE A DUMMY ARRAY ANO NI MUST BE EQUAL TO GNE. the logical variable model will get the value true in this case.
(2) A NUMBER (N1) OF THE ANGMALY DEGREE-VARIANCES (DEGREE ZERO TO Nl-1) ARE PUT EQUAL TO EMPIRICAL DETERMINED DEGREE-VARIANCES. THE DEGREE-VARIANCE OF DEGREE K WILL HAVE TO BE STORED IN EPS (K+1) (IN UNITS OF MGAL**2).
(3) THE DEGREE-VARIANCES OF DEGREE ZERO TO N = N1-1 ARE PUT EQUAL TO ZERO, (AND THE OTHERS ARE THE SAME AS ABOVE DESCRIBED). THIS MEANS THAT AN N•TH ORDER LOCAL COVARIANCE FUNCTION WILL BE COMPUTED. IN THIS CASE EPS MUST HAVE NI ZERO VALUES STORED.
C IA ALL CASES NI MUST BE LESS THAN 300 AND EPS MUST HAVE DIMENSION
C NI.
IMPLICIT REAL *8(A-H,O-Z)
LOGICAL MODEL, NOTD, NOTDD
DIMENSION EPSC(300), EPS(1)
DATA RE,GM, A,SE,B,IB1,IB2,IBM1,EPSC(1),EPSC(2),D0,D1,D2,D3,D4,
※D5,RADSEC/6371.003,3.98D14,425.2800,0.99961700,24.000,25,26,23,
*3*0.000, 1.0D0,2.000,3.000,4.000,1.005,206264.80600/
$I B 12=I B 1$ 减 IB 2
RADSE2 $=$ RADSEC**2
RE2 $=$ RE*RE
RBJ2 $2=R E 2 \neq S E$
RBJ = DSQRT(RBJ2)
$A M=A / D 5$
$A M 2=A M / D 5$
C A IS IN UNITS OF MGAL**2, AM IN UNITS OF MGAL*M/SEC AND AMZ IN
C UNITS OF (M/SEC)**2. RBJ IS THE RADIUS OF THE BJERHAMMAR-SPHERE. MODEL = N1.EQ. 1
IF (MODEL) GO TO 20
C
C WE WILL NOW CGMPUTE THE MODIFIED (POTENTIAL) DEGREE-VARIANCES, CF.
C EQUATION (151).
IF (N1.LT.3) GO TO 20
DO $10 \mathrm{I}=3, \mathrm{Nl}$
$R I=$ DFLOAT $(I-1)$
IF (I.EQ.3) EPS(3) = EPS (3) *RBJ2*1.0D-10
10 IF (I.GT.3) EPS(I) =

* RBJ2*(EPS(I)/((RI-D1)**2)*1。OD-10-AM2/((RI-D1)*(RI-D2)*(RI+B)))

20 RETURN

ENTRY COVB(KTYPE)
BY THE CALL OF COVB, THE TYPE OF COVARIANCE TO BE COMPUTED IS DETER-
mined by the value of ktype, so that we get the covariance between:
THE GRAVITY ANOMALY AT P AND THE GRAVITY ANOMALY AT $Q$ FOK KTYPE=1,
THE - - $\quad$ - THE LONGITUDIONAL COMPONENT
OF THE DEFLECTION OF THE VERTICAL AT 0 FOR KTYPE=2,
THE - - AT P AND THE HEIGHT ANOMALY AT O FOR KTYPE=3,
THE LONGITUDIONAL COMPONENT OF THE DEFLECTION OF THE VERTI-
CAL AT P AND THE SAME TYPE OF QUANTITY AT 0 FOR KTYPE=4,
the transversal component of the oeflection of the verti-
CAL AT P AND THE SAME TYPE OF QUANTITY AT Q FOR KTYPE=5,
THE LONGITUDIONAL COMPONENT OF THE DEFLECTION OF THE VERTI-
CAL AT $P$ AND THE HEIGHT ANOMALY AT $Q$ FOR KTYPE=6,
AND THE HEIGHT ANOMALY AT $P$ AND THE HEIGHT ANOMALY AT $Q$ FOR KTYPE=7.
the value of ktype will then also determine which of the coefficients
(151)-(153), THAT WE WILL USE IN THE EVALUATIUN OF THF LEGENDRE-SERIES
AND Whether NO DIFFERENTIATION, DIFFERENTIATION ONE TIME OR DIFFEREN-
TIATION TWO TIMES WITH RESPECT TO THE VARIABLE T TAKES PLACE. TWO
LOGICAL VARIABLES NOTD AND NOTDU $\triangle R E$ USED TO UISTINGUISH BETWEEN THE
SITUATIONS.
IF (MODFL) GO TO 35
IF (KTYPE.EO.1) IP = 2
IF (KTYPE.EQ.2.OR.KTYPE.EQ.3) IP = 1
IF (KTYPE.GT.3) IP =
DO $30 \mathrm{I}=3, \mathrm{NI}$
$30 \operatorname{EPSC}(I)=\operatorname{EPS}(I) *((I-2) * D 5 / R B J) * * I P$

```
35 NOTD = KTYPE.EQ.1.OR.KTYPE.EQ.3.OR.KTYPE.EQ.7
    NOTDD = KTYPF.NE.5.AND.KTYPE.NE.4
    RETURN
```

    ENTRY COVC(PSI,HP,HQ,COV)
    by the call of covc the covariance of type ktype will be computed for
POINTS P AND D HAVING SPHERICAL DISTANCE (RADIANS) PSI, WHERE HP IS
THE HEIGHT OF $P$ above the earth and hQ the height of o above the
EARTH. THE COVARIANCE WILL BE RETURNED BY THE VARIABLE COV. UNITS aRE
PRODUCTS OF MGAL. METERS AND ARCSECONDS.

```
T= DCOS(PSI)
U = DSIN(PSI)
T2 = T*T
U2=U*U
RP=RE+HP
RQ = RE+HQ
S = RBJ2/(RP&RQ)
S2 = S*S
S3=S2*S
TS = T*S
P2 = (D3*T2-D1)/D2
GP = GM/(RP*RP)
GQ = GM/(RQ*RQ)
```

C THE QUANTITIES L.M AND N DEFINED IN EQ. (75) ARE HERE CALLED SL, SM C AND SN. L**2 $2=S L 2$.
$S L 2=01+S 2-1) 2 \star T S$
$S L=$ DSQRT(SL2)
$S L 3=S L 2 * S L$
$S N=D 1-T S+S L$
$S M=D 1-T S-S L$
$S L N=S L * S N$
SLNL $=-$ DLOG(SN/D2)
C WHEN WE ARE COMPUTING A LOCAL NITH ORDER COVARIANCE OR A COVARIANCE C FROM A GLOBAL MODEL WITH EMPIRICAL DEGREE-VARIANCES UP TO AND INCLUC SIVE DEGREE N, WE WILL HAVE TO COMPUTE THE SUM (154), THE SUM (155)
C (WHEN NOTD IS FALSE) AND THE SUN (156) (WHEN NOTDD IS FALSE). (154)

C WILL BE ACCUMMULATED IN BO, (155) IN DBO AND (156) IN DDBO.
C WHEN THE VARIABLE MODEL IS TRUE, BO, DBO AND DDBO WILL BE PUT EQUAL
C TO ZERO.
$B 0=00$
$D E O=00$
DOBO $=00$
IF (MODEL: GO TO 45
C
$B 1=D 0$
DB1 $=$ DO
DDB1 $=00$
$\mathrm{LI}=\mathrm{N} 1$
RLI $=$ DFLOAT(LI)
C
C WE WILL NOW USE THE RECURSION FORMULAE (183), (185) AND (186), WHERE
C THE TERM (176A) DIVIDED BY T IS CALLED EL AND FLI IS THE TERM (176B)
C FOR SUBSCRIPT $L+1$.
DO $40 \mathrm{I}=1, \mathrm{NI}$
$E L=(D 2 * R L I-D 1) * S / R L 1$
$F L 1=-R L 1 * S 2 /(R L 1+D 1)$
$R L 1=R L 1-D 1$
$B 2=B 1$
$B 1=B 0$
$B 0=B 1 * E L * T+B 2 * F L 1+E P S C(L 1)$
IF (NOTD) GO TO 40
C
$\mathrm{DB2}=\mathrm{DB} 1$
$D B 1=D B O$
$D B O=E L *(D B 1 * T+B 1)+F L 1 * D B 2$
IF (NOTDD) GO TO 40
C
DDB2 $=$ DDB1
DDB1 $=$ DDBO
DDBO $=E L *(D B 1 * D 2+D D B 1 * T)+F L 1 * D D B 2$
$40 \mathrm{Ll}=\mathrm{L} 1-1$
C

C COMPUTATION DF CLOSED EXPRESSIONS. FIRST SOME AUXILLIARY QUANTITIES.
C FM1 IS THF OUANTITY (86), FM2 IS (87), F1 IS (99) AND F2 IS (100)
$45 \mathrm{DPL}=\mathrm{DI}+\mathrm{SL}$
$D M L=O 1-S L$
P31 $=03 \div T S+D 1$
$B O=B O * S$
$F M I=S *(S M+T S * S L N L)$
$F M 2=S *(S M * P 31 / D 2+S 2 *(P 2 * S L N L+U 2 / D 4))$
$F l=D L O G(D 1+D 2 * S /(D 1-S+S L))$
$F 2=(S L-D 1+T * F 1) / S$
IF (NOTD) GO TO 48
C
$D B O=D B O * S$
C DFMI IS THE QUANTITY (90), DFM2 IS (92), DFI IS (101) AND DF2 IS
C (103).
DFM1 $=S 2 *(D M L / S L+S L N L+T S *(D 1 / S L N+D 1 / S N))$
DFM2 $=S 2 *((P 31 / S L+D 2-7.000 * T S-D 3 * S L) / D 2+S *(D 3 * T * S L N L$

* +S九P2ヶDPL/SLN))
$D F 1=S 2 / S L N$
$D F 2=-[1 / S L+T S / S L N+F 1 / S$
$D L=-S / S L$
IF (NOTDD) GO TO 48
C
DDBO $=$ DDBO $\because S$
C DDFM1 IS THE QUANTITY (91), DDFM2 IS (93), DDFI IS (102) AND DOF2 IS
C (104).
DDFM1 $=S 3 *(D 1 / S L 3+D 2 * D P L / S L N+T S *(D 1 /(S L 3 * S N)+(D P L / S L N) * * 2))$
C
DDFM2 $=S 3 *(16.0 D 0 / S L+P 31 / S L 3-7.000) / D 2+D 3 * S L N L+6.000 * T S * D P L / S L N$ $+P 2 * S 2 *((D P L / S L N) * * 2+D 1 /(S L 3 * S N)))$
DDF1 $=S 3 *(D P L / S L N * 2+D 1 /(S N * S L 3))$
DDF2 $=(-S 2 / S L 3+$ D2 $*$ DF $1+T * D D F 1) / S$
$D D L=-S 2 / S L 3$
C WE CAN NOW USE THE RECURSION FORMULAE (96), (97) AND (98) FOR THE
C CIMPUTATION OF THE QUANTITY (73) CALLED FB AND ITS DERIVATIVES DFB
C AND DDFB.
C
$48 \mathrm{DO} 50 \mathrm{I}=2,1 \mathrm{BMI}$
RI = DFLOAT(I)
DI2 $2=$ D2*RI-DI
$D 11=(R I-D 1) / S$
$F B=(S L+O I 2 * T * F 2-D I l * F 1) /(R I * S)$
$F 1=F 2$
$F 2=F B$
IF (NOTD) GO TO 50
$D F B=(D L+D I 2 *(F 1+T * D F 2)-D I l * D F 1) /(R I * S)$
$D F 1=D F 2$
$D F 2=D F B$
IF (NOTDD) GO TO 50
$D D F B=(D D L+D I 2 *(D 2 * D F 1+T * D D+2)-D I 1 * D D F 1) /(R I * S)$
DDF $1=$ DDF 2
DDF2 $=$ DDFB
50 CONTINUE

```
        IF (NOTD.OR.KTYPE.FO.2) GO TO 60
    C FROM (133) WE HAVE:
            DK=DBO+AM2*RBJ2*(IB1*OFM2-IB2*(DFM1-D3*T*S3)+DFB-S2/IB1-D 3*S 3*T/
        * IB2I/IB12
    60 GO TO (61,62,63,64,65,66,67),KTYPE
C EQUATION (132) AND (146) GIVES:
    61COV = S*RO+A*S*(IB1*(FB-S/B-S2*T/IB1-S3*P2/IB2)+FM2)/IB2
        GO TO 70
C. EQUATION (139) AND (150) GIVES:
    62C[IV = U*(DBO*RBJ/(RP*RQ)+AM*S*(DFM2-DFB+S2/IB1+D 3*S 3*T/IB2)/IR2)/
        * GQ*RADSEC
            GO TO 70
C EQUATION (131) AND (145) GIVES:
    63 COV = (BO*RBJ+AM*RBJ2*(FM2-FB+S/B+S2*T/IR1+S 3*P2/IB2)/IB2)/
        *(RP*GO)
            GO TO 70
C EQUATION (136) AND (147) GIVES:
    64COV = (T*DK/(RP*RQ)-U2*(DOBO/(RP*RO)+AM2*S*(IB1*DDFM2-IR2*(DOFM1
        * -D3*S3)+DDFB-D3*S3/IB2)/IB12))*RADSE2/(GP*GQ)
            GO TO 70
C EQUATION (137) AND (148) GIVES:
    65 COV = DK/(RP*RQ*GP*GQ)*RADSE2
        GO TO 70
C EQUATION (138) AND (149) GIVES:
        66 COV = U*DK/(GP*GO*RP)*RADSEC
        GO TO 70
C AND EQUATION (37), (130) AND (144) GIVES:
        67COV = (BO+AM2*RB\2*(IB1*FM2-IB2*(FM1-S 3*P2)+FB-S/B-S 2*T/IB1-S 3*P2
        * /IB2)/IB12)/(GP*GQ)
    70 RETURN
        END
```


[^0]:    * from equation (16A)

